Anions in Space and in the Laboratory

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lons are ubiquitous!



Interstellar Molecules

	СН	CN	CH⁺	OH	NH ₃	H ₂ O	H ₂ CO
	СО	H ₂	HCO⁺	CH₃OH	HC₃N	HCN	нсоон
	SiO	CS	CH₃CN	OCS	NH₂CHO	H ₂ S	HNCO
	CH₃CHO	CH₃CCH	CH₂NH	H ₂ CS	HNC	SO	CH ₃ OCH ₃
	CH₃NH₂	N_2H^+	C₂H	CH₂CHCN	CH₃CH₂OH	HCOOCH₃	SO ₂
	HDO	SiS	NS	NH ₂ CN	НСО	C ₃ N	H ₂ CCO
	C ₂	HNO	CH ₃ CH ₂ CN	HC ₇ N	HC₅N	HC ₉ N	C₄H
	NO	OCN⁻	CH₃SH	HNCS	C ₂ H ₄	HCS⁺	HOCO ⁺
ĺ	HOC⁺	CH ₃ C ₃ N	SiH ₄	CH₃C₄H	c-SiC ₂	C₃H	HCI
	C ₃ O	c-C ₃ H ₂	C₀H	HCNH ⁺	MgNC	C₅H	H_3O^+
	C ₂ S	C ₃ S	(CH ₃) ₂ CO	NaCl	AICI	KCI	AIF
	PN	CH ₃ NC	C ₃	c-C₃H	CH₂CN	HC₂CHO	C ₅
	SiC	C_2H_2	SiC ₄	CO ₂	CH₂	СР	$I-C_3H_2$
	HC₂N	NH	CH₄	C ₂ O	HCCNC	SiN	HNCCC
ĺ	SO⁺	NH ₂	CO ⁺	HC₃NH⁺	H₂CN	NaCN	N ₂ O
	MgCN	C ₈ H	H_3^+	H₂COH ⁺	C ₇ H	CH₃COOH	HC ₁₁ N
	HF	c-C ₂ H ₄ O	LiH	C₅N	SiC₃	SH	CH ₃
	CH₂OHCHO	SiCN	C_4H_2	C_6H_2	C ₆ H ₆	CH₂CHOH	AINC
	FeO	HOCH ₂ CH ₂ OH	NH ₂ CH ₂ COOH	N ₂	CH ₂ CHCHO	CH₃CH₂CHO	SiNC
	HC₄N	CO(CH ₂ OH) ₂	CH ₂ CCHCN	c-H ₂ C ₃ O	CH ₃ CONH ₂	CH ₃ C ₆ H	CH ₂ CNH
ĺ	CF⁺	CH ₃ C ₅ N	C ₆ H [−]	O ₂	C₄H⁻	НСР	CଃH⁻
	CH ₂ CHCH ₃	РО	CNCHO	ССР	C₃N⁻	NH ₂ CH ₂ CHN	PH ₃
	C₅N [−]	HCNO	AIO	HOCN	C ₂ H ₅ OCHO	C ₃ H ₇ CN	HSCN
	AIOH	CN⁻	H ₂ O ⁺	OH ⁺	C ₆₀	C ₇₀	H₂CI ⁺
	KCN	SH ⁺	FeCN				

http://www.astrochymist.org/astrochymist_ism.html

Laboratory and Astronomical Identification of the Negative Molecular Ion C₆H⁻ McCarthy, Gottlieb, Gupta and Thaddeus *Harvard-Smithsonian Center for Astrophysics*



Astrophysical Journal 652, L141 December 2006.

17 rotational lines between 8 & 187 GHz

Identified C_6H^- in IRC +10216 & TMC-1

 $C_4H^ C_6H^ C_8H^ CN^ C_3N$

Interstellar Molecules

СН	CN	CH⁺	ОН	NH ₃	H ₂ O	H ₂ CO
СО	H ₂	HCO⁺	CH₃OH	HC₃N	HCN	НСООН
SiO	CS	CH₃CN	OCS	NH ₂ CHO	H ₂ S	HNCO
CH₃CHO	CH₃CCH	CH₂NH	H₂CS	HNC	SO	CH ₃ OCH ₃
CH ₃ NH ₂	N₂H ⁺	C₂H	CH ₂ CHCN	CH ₃ CH ₂ OH	HCOOCH ₃	SO ₂
HDO	SiS	NS	NH ₂ CN	НСО	C₃N	H ₂ CCO
C ₂	HNO	CH ₃ CH ₂ CN	HC ₇ N	HC₅N	HC₃N	C₄H
NO	OCN [−]	CH₃SH	HNCS	C_2H_4	HCS⁺	HOCO⁺
HOC⁺	CH ₃ C ₃ N	SiH ₄	CH ₃ C ₄ H	c-SiC₂	C₃H	HCI
C ₃ O	$c-C_3H_2$	C ₆ H	HCNH⁺	MgNC	C₅H	H_3O^+
C ₂ S	C₃S	(CH ₃) ₂ CO	NaCl	AICI	KCI	AIF
PN	CH₃NC	C ₃	c-C₃H	CH ₂ CN	HC ₂ CHO	C ₅
SiC	C_2H_2	SiC ₄	CO ₂	CH ₂	СР	$I-C_3H_2$
HC ₂ N	NH	CH₄	C ₂ O	HCCNC	SiN	HNCCC
SO⁺	NH ₂	CO⁺	HC₃NH⁺	H ₂ CN	NaCN	N ₂ O
MgCN	CଃH	H_3^+	H₂COH⁺	C ₇ H	CH₃COOH	HC ₁₁ N
HF	c-C ₂ H ₄ O	LiH	C₅N	SiC ₃	SH	CH ₃
CH ₂ OHCHO	SiCN	C_4H_2	C_6H_2	C ₆ H ₆	CH ₂ CHOH	AINC
FeO	HOCH ₂ CH ₂ OH	NH ₂ CH ₂ COOH	N ₂	CH ₂ CHCHO	CH ₃ CH ₂ CHO	SiNC
HC₄N	CO(CH ₂ OH) ₂	CH ₂ CCHCN	c-H ₂ C ₃ O	CH ₃ CONH ₂	CH ₃ C ₆ H	CH ₂ CNH
CF⁺	CH ₃ C ₅ N	C ₆ H [−]	O ₂	C₄H⁻	НСР	CଃH⁻
CH ₂ CHCH ₃	PO	CNCHO	ССР	C₃N [−]	NH ₂ CH ₂ CHN	PH ₃
$C_5 N^-$	HCNO	AIO	HOCN	C₂H₅OCHO	C ₃ H ₇ CN	HSCN
AIOH	CN⁻	H_2O^+	OH⁺	C ₆₀	C ₇₀	H_2CI^+
KCN	SH ⁺	FeCN				

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Interstellar Molecular Synthesis



FA-SIFT-MS Instrument

Ion Production

Ion Selection

Reaction Flow Tube

Ion Detection

Flowing Afterglow-Selected Ion Flow Tube (FA-SIFT)



Features of the Flowing Afterglow

- > Thermal energy
- > Energy variability
- → Kinetic analysis
- High ion

density/sensitivity

 Coupling with other chemical versatility techniques lonic and neutral reactants

Negative Ion Chemistry

Reactions of Negative lons

- 1. Carbon chains (C_x⁻)
- 2. Hydrogenated carbon chains (HC_x⁻)
- 3. Organic anions $(H_w C_x N_y O_z)$
- 4. Nitrogen-containing carbanions
- (C_xN_y⁻) 5. Hydride anion (H⁻) → Summary

Future Directions (including PAH⁻)



Mass Spectrum for Cold Cathode DC Discharge



Mass Selecting and Injecting C₇⁻





Reactions of Carbon Chain Anions with H Atom

Reactant	Products	BR	k (cm³/s)
C ₄			
C ₅			
C ₆			
C ₇			
C ₈			
C ₉			
C ₁₀			

Reactions of Carbon Chain Anions with H Atom

Reactant	Products	BR	k (cm³/s)
C ₄	C₄H + e ⁻		6.2 x 10 ⁻¹⁰
C ₅	C₅H + e [¯]		6.2 x 10 ⁻¹⁰
C ₆	C ₆ H + e ⁻		6.1 x 10 ⁻¹⁰
C ₇			
C ₈			
C ₉			
C ₁₀			

Reactions of Carbon Chain Anions with H Atom

Reactant	Products	BR	k (cm ³ /s)
C ₄	C₄H + e ⁻		6.2 x 10 ⁻¹⁰
C ₅	C₅H + e ⁻		6.2 x 10 ⁻¹⁰
C ₆	C ₆ H + e ⁻		6.1 x 10 ⁻¹⁰
C ₇	C ₇ H + e ⁻	0.41	6.9 x 10 ⁻¹⁰
	C7H	0.59	
C ₈	C ₈ H + e ⁻	0.33	7.3 x 10 ⁻¹⁰
	CଃH	0.67	
C ₉	C ₉ H + e [¯]	0.17	7.2 x 10 ⁻¹⁰
	C₃H	0.83	
C ₁₀	C ₁₀ H + e	0.24	7.5 x 10 ⁻¹⁰
	C ₁₀ H ⁻	0.76	

Reactions of Cx with O Atom

Reactant	Products	k (cm³/s)
C ₂		
C4		
C ₅		
C ₆		
C ₇		

Reactions of Cx with O Atom

Reactant	Products	k (cm³/s)
C ₂	C + CO	5.8 x 10 ⁻¹⁰
C ₄	C_3 + CO	5.6 x 10 ⁻¹⁰
C ₅	C ₄ + CO	6.4 x 10 ⁻¹⁰
C ₆	C ₅ + CO	4.7 x 10 ⁻¹⁰
C ₇	C ₆ + CO	5.3 x 10 ⁻¹⁰

Reactions of Cx with N Atom

Reactant	Products	k (cm³/s)
C ₂		
C ₄		
C ₅		
C ₆		
C ₇		

Reactions of Cx with N Atom

Reactant	Products	k (cm ³ /s)	EA(C _x) eV
C ₂	CN ⁺ C	2.3 x 10 ⁻¹⁰	3.27
C ₄	$CN^{-} + C_3$	2.0 x 10⁻¹⁰	3.88
	C_3 + CN		
C ₅	CN + C4	2.7 x 10 ⁻¹⁰	2.84
	C_4 + CN		
	$C_3N + C_2$		
C ₆	$CN^{-} + C_{5}$	1.5 x 10 ⁻¹⁰	4.19
	C₅ + CN		
	$C_3N + C_3$		
C ₇	$CN^{-} + C_{6}$	2.2 x 10 ⁻¹⁰	3.39
	C_6 + CN		
	$C_3N + C_4$		
	$C_5N + C_2$		

Negative Ion Chemistry

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Future Directions (including PAH⁻)



Nitrogen-Containing Carbanions





Reactions of C_x N_y^- with H atoms $C_x N^- (x = 1 - 6)$



Reactions of C_xN⁻ with H atoms







Reactions of C_x N_y^- with H atoms $C_x N_2^-$ (x = 1, 3 – 5)





kcal/mol, CCSD(T)/aug-cc-pVDZ//B3LYP/aug-cc-pVTZ

Reactions of $C_x N_y^-$ with H atoms $C_x N_3^-$ (x = 2, 4) b) $C_4 N_3^-$ + H



High energy TS and Reaction

Endothermic pathways



kcal/mol, CCSD(T)/aug-cc-pVDZ//B3LYP/aug-cc-pVTZ

Interstellar Molecular Synthesis



Summary – General Themes Reactions of Negative Ions + Unreactive with H₂

- → React with H by associative detachment Rate often correlates with exothermicity Some fragmentation pathways Alternation in reactivity for x=even or odd
- → React with N and O Rich variety of pathways O-atom more rapid than N-atom
- Computations

 C_x^{-}

HC_v⁻

Nitriles

Aldehydes

Ketones

Esters

Acids

Alcohols

Glycine

 $C_x N_v^{-}$

- Provide insight to products & energies Importance of spin conservation
- > Processes provide routes to neutrals and ions observed in interstellar clouds

Future Directions

- Quantify product ratios for anion reactions
 Account for associative detachment and ionic products,
 mass discrimination, secondary reactions
- → Study additional reactions of C_xN_y⁻ With N and O and other reagents

Study PAH⁻ (and larger PAH⁺/PAH⁻)
 Develop and implement LIAD and ESI sources

PAH Anions Inclusion of PAHs in dense clouds

- PAH⁻ become the dominant carriers of negative charge
- Free e⁻ are replaced by PAH⁻
 Reduces overall ionization fraction
 - Neutralization of atomic cations is enhanced

Wakelam & Herbst, ApJ 680, 371 (2008)

 Deprotonated PAHs are more stable than the parent radical anions

Hammonds & Sarre,

Poster 1.34

Electron affinities of PAHs and dehydrogenated PAH radicals



Laser Induced Acoustic Desorption



Electrospray Ionization



•High voltage (1-5 kV) supplied to needle.





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Reactions of Organic Anions with H atoms

