

**Gas-Phase Reactions in the ISM:
Rate coefficients, temperature-dependences,
and reaction products**

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A little (personal) history

1987: Rate Coefficients in Astrochemistry (TJ Millar & DA Williams)

(IWMS) Experimental Measurements of the Rate Constants for Neutral-Neutral Reactions

1996: IAU Symposium 178: Molecules in Astrophysics: Probes and Processes (E F van Dishoeck)

(IWMS) Reactions between Neutral Species at Low Temperatures: Laboratory Results and Astrophysical Modelling

2011: IAU Symposium 280: Gas-Phase Chemistry in the ISM: Rate coefficients, temperature-dependences, and reaction products

Gas-phase reactions of different types included in the OSU kinetic database (version osu-09-2008) for astrochemistry.

Type of process	Example	Number in model
Gas-grain interactions:	$\text{H} + \text{H} + \text{grain} \rightarrow \text{H}_2 + \text{grain}$	14
Direct cosmic ray processes	$\text{H}_2 + \zeta \rightarrow \text{H}_2^+ + \text{e}$	11
Cation-neutral reactions	$\text{H}_2^+ + \text{H}_2 \rightarrow \text{H}_3^+ + \text{H}$	2933
Anion-neutral reactions	$\text{C}^- + \text{NO} \rightarrow \text{CN}^- + \text{O}$	11
Radiative associations (ion)	$\text{C}^+ + \text{H}_2 \rightarrow \text{CH}_2^+ + h\nu$	81
Associative detachment	$\text{C}^- + \text{H}_2 \rightarrow \text{CH}_2 + \text{e}$	46
Chemi-ionization	$\text{O} + \text{CH} \rightarrow \text{HCO}^+ + \text{e}$	1
Neutral-neutral reactions	$\text{C} + \text{C}_2\text{H}_2 \rightarrow \text{C}_3\text{H} + \text{H}$	382
Radiative association (neutral)	$\text{C} + \text{H}_2 \rightarrow \text{CH}_2 + h\nu$	16
Dissociative recombination	$\text{N}_2\text{H}^+ + \text{e} \rightarrow \text{N}_2 + \text{H}$	539
Radiative recombination	$\text{H}_2\text{CO}^+ + \text{e} \rightarrow \text{H}_2\text{CO} + h\nu$	16
Anion-cation recombination	$\text{HCO}^+ + \text{H}^- \rightarrow \text{H}_2 + \text{CO}$	36
Electron attachment	$\text{C}_6\text{H} + \text{e} \rightarrow \text{C}_6\text{H}^- + h\nu$	4
External photo-processes ^a	$\text{C}_3\text{N} + h\nu \rightarrow \text{C}_2 + \text{CN}$	175
Internal photo-processes ^a	$\text{CO} + h\nu \rightarrow \text{C} + \text{O}$	192

Kinetic information needed for Chemical Models

- Rate coefficient and its dependence on temperature. For reaction between *A* and *B*:

$$- d[A] / dt = - d[B] / dt = k(T) [A] [B]$$

- In models: $k(T) = \alpha (T/300)^\beta \exp(-\gamma / T)$ — **but caution!**

- **Products of reaction**

- (a) thermodynamics **may** demonstrate that only one channel is open;
- (b) determination of 'branching ratios' can be 'challenging'

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Ion-Neutral Reactions: Experimental Methods

■ 'Trapping Methods':

(a) ICR – Ion Cyclotron Resonance: **300 K**

(b) Ion Traps: especially 22-pole trap (Gerlich): **down to 10 K**, especially useful for study of *radiative association*

■ 'Flow Methods':

(a) FA – Flowing Afterglow: **$82 \leq (T/K) \leq 600$** ; later **$300 \leq (T/K) \leq 1800$**

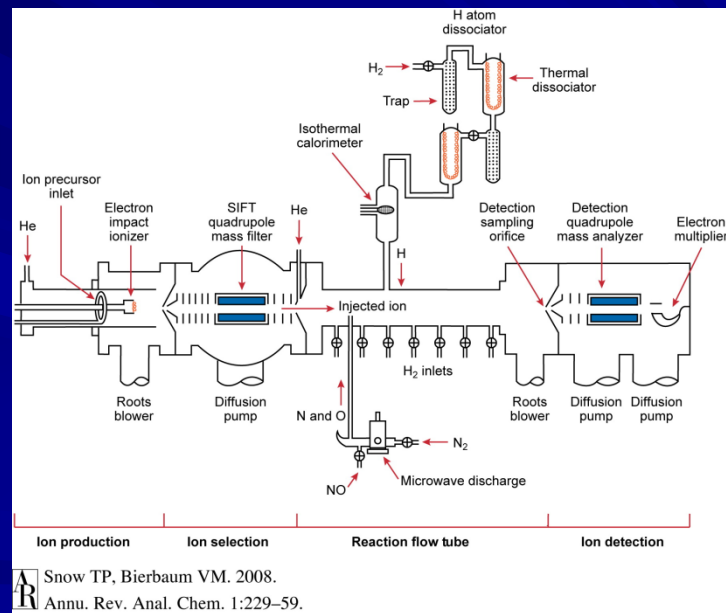
(b) SIFT – Selected Ion Flow Tube: typically, **down to 80 K**, also one study from **18 K to 295 K**

(c) SIFDT – Selected Ion Flow Drift Tube extends measurements to much higher collision energies; **up to ≈ 1800 K**;

(d) CRESU – Reaction Kinetics in Uniform Supersonic Flows: **$8 \leq (T/K) \leq 300$**

Ion-Neutral Reactions: Experimental Methods

- As **mass spectrometry** is used to observe loss of ionic reactant, ionic product(s) can be observed – hence, **branching ratios**
- Methods used for reactions of cations can be adapted to study the **reactions of anions** (*Veronica Bierbaum*)
- In most studies of ion-neutral reactions, the neutral reactant is a ‘stable’ (i.e., non-radical) species. However, SIFT method adapted to study of **reactions between ions and radical atoms** (*Veronica Bierbaum*)



(Snow & Bierbaum, *Ann. Rev. Anal. Chem.* 44, 367 (2006))

Ion-Neutral Reactions: T -dependence of rate coefficients

- For **most** ion-neutral –reactions, there is no *activation barrier*, rather the rate coefficient is determined by ‘capture’: that is, the ability of the long-range ‘electrostatic potential’ to bring the reactants into close contact ‘against’ the requirement to conserve angular momentum
- The ‘**Langevin model**’ assumes attraction between charge and induced dipole, leading to:

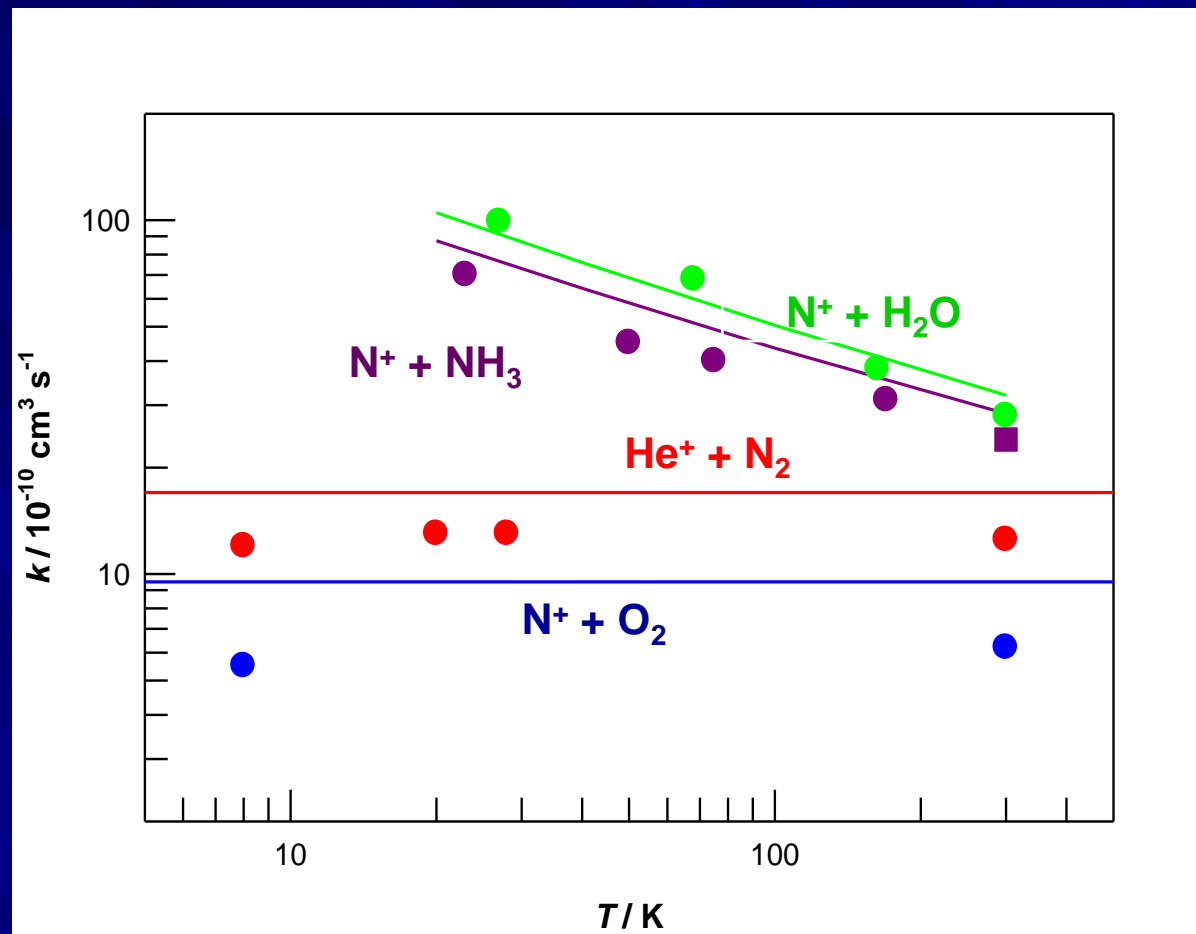
$$k_L = 2\pi e (\alpha / \mu)^{1/2}$$

- For reactions between ions and dipolar neutral molecules, the long-range potential is more complex (depends on orientation). Ratio of rate coefficient to the Langevin value has been parameterised:

$$k_D / k_L = ax + b \text{ where } x = \mu_D / (2\alpha k_B T)^{1/2}$$

Ion-Neutral Reactions: tests of simple capture models

CRESU experiments provide a good test of predictions of T-dependence for reactions between simple ions and neutral (non-polar and polar) molecules



Neutral-Neutral Reactions: Experimental Methods

- **'Discharge-Flow Methods':**
 - (a) Reaction of atomic (H, O, N, Cl) and diatomic radicals:
typically, $200 \leq (T/K) \leq 500$
 - (b) optical and mass spec. methods to observe loss of reactants
(formation of products?)
 - (c) HTFFR (Fontijn) reaches ca. **1320 K**
- **'Pulsed Laser Photolysis Methods':**
 - (a) coupled to optical methods for detection (e.g. LIF):
typically $200 \leq (T/K) \leq 600$ some experiments **down to 80 K**
 - (b) HTPR (Fontijn) reaches ca. **1430 K**
 - (c) CRESU – Reaction Kinetics in Uniform Supersonic Flows:
 $13 \leq (T/K) \leq 300$

Both methods provide data on (radical + molecule) and (radical + radical) reactions but are difficult at extremes of temperature

Measuring product yields/branching ratios is generally difficult

(<http://www.iupac-kinetic.ch.cam.ac.uk>; <http://jpldataeval.jpl.nasa.gov>)

(Baulch, D.L. et al., 2005, *J. Phys. Chem. Ref. Data*, 34, 757)

Neutral-Neutral Reactions: in the ISM and CRESU Experiments

- Of the observed 159 species in the ISM (July 2010): 17 are cations, 6 anions, 136 electrically neutral
- Of the 136 neutral species: many are free radicals (e.g., CH, CN, C₂H, C₄H, etc) and many are 'unsaturated' (C₂H₂, C₂H₄, HC₂CN, HC₄CN etc)
- Clearly scope for neutral-neutral reactions
- In CRESU experiments, radicals (C, Al, Si, B, O, CN, OH, CH, C₂H, C₂, C₄H) are formed by pulsed laser photolysis, removal followed by LIF or using chemiluminescent techniques
- Co-reactants: principally HC's, CH₄, C₂H₄, C₂H₂,.....
- Radical-radical reactions difficult to study – only one $O + OH \rightarrow O_2 + H$

Neutral-Neutral Reactions: T -dependence of rate coefficients

- For most reactions between radicals and saturated molecules, $\gamma > 0$, $k(T)$ increases with T : e.g. $\text{CN} + \text{H}_2$ and $\text{C}_2\text{H} + \text{H}_2$
- For reactions where $k(298 \text{ K}) \geq 10^{-11} \text{ cm}^3 \text{ s}^{-1}$: general trend is for $k(T)$ to increase as T is lowered
- However, the form of $k(T)$ is variable
- Suggests the absence of an activation barrier and that values of $k(T)$ are determined by capture
- Product yields/branching ratios difficult to measure

One example: $\text{CN} + \text{NH}_3 \rightarrow \text{products}$

- Can we explain the T -dependence?
- Can we determine the products? That is $\text{HCN} + \text{NH}_2$ or $\text{NCNH}_2 + \text{H}$?

Neutral-Neutral Reactions: branching ratios

- No 'universal' method
- *Ab initio* calculations
- Experiments on reactions that yield H atoms
 - a) Bordeaux group (Bergeat, Loison, et al) use a flow system and create CH radicals ($\text{CHBr}_3 + 3 \text{K} \rightarrow \text{CH} + 3\text{KBr}$) and observe H atoms by 'resonance fluorescence'
 - b) Leeds group (Blitz, Seakins, et al) carry out time-resolved experiments $\text{CHBr}_3 + 3 h\nu \rightarrow \text{CH} + 3\text{Br}$ and observe H atoms by LIF 'laser-induced fluorescence'

Note that a number of important low T reactions involve addition of radical followed by H-atom elimination



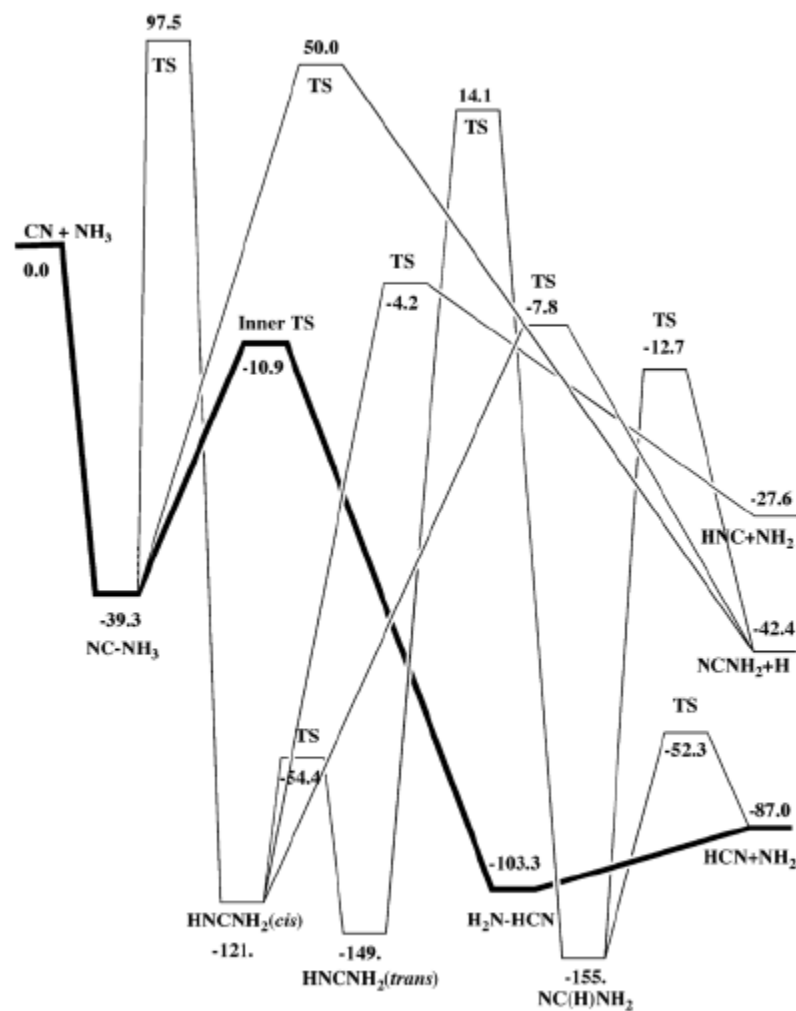
Neutral-Neutral Reactions: Theory and Experiments on $\text{CN} + \text{NH}_3 \rightarrow \text{products}$

- *Ab initio* calculations of reaction path (Dahbia Talbi)

No low energy path to $\text{NCNH}_2 + \text{H}$;
reaction proceeds 100% to $\text{HCN} + \text{NH}_2$

- Calculations of $k(T)$ using two-transition state method of Klippenstein & Georgievskii

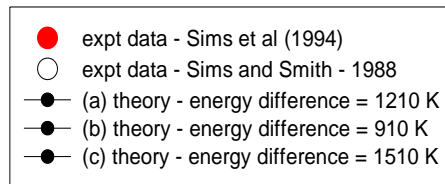
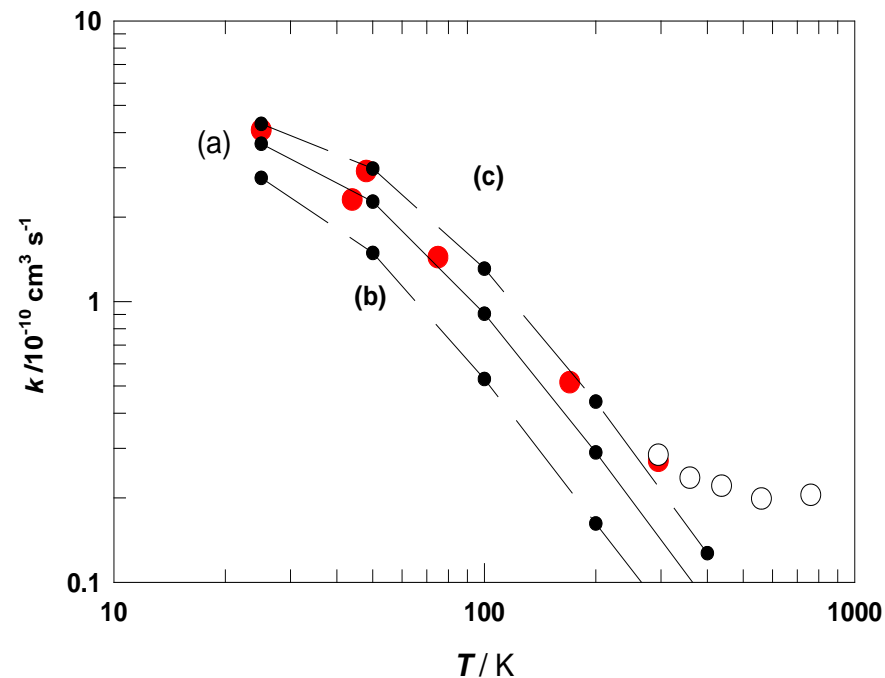
(D Talbi & IWMS, *PCCP*, 11, 8477 (2009))



Neutral-Neutral Reactions: Theory and Experiments on $\text{CN} + \text{NH}_3 \rightarrow \text{products}$

Large points show experimental rate coefficients from 23 to ca. 750 K

Small points and lines show results of TST calculations with different choices for the inner transition state barrier

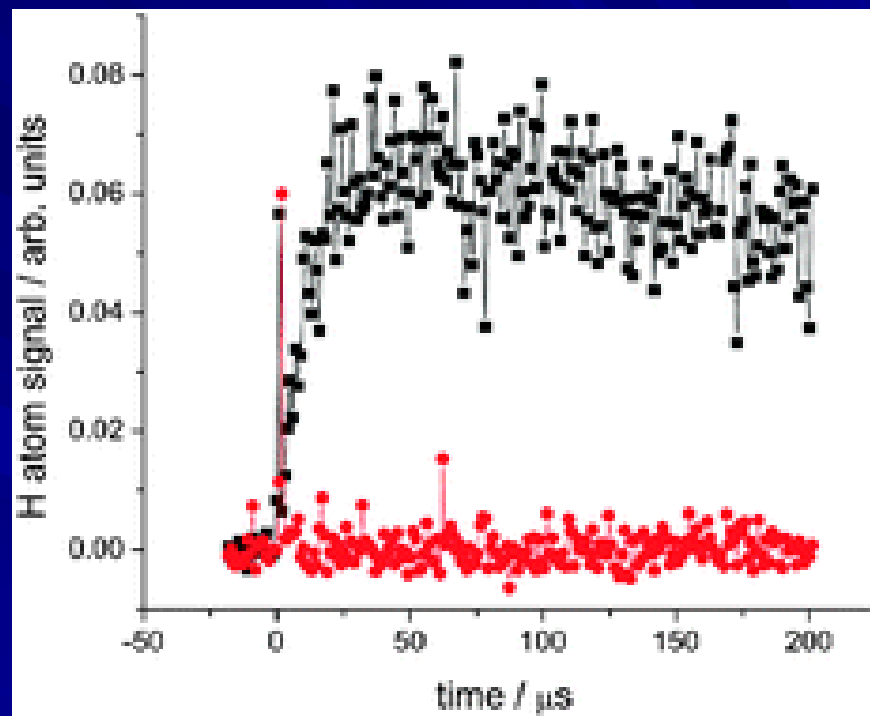


Neutral-Neutral Reactions: Theory and Experiments on $\text{CN} + \text{NH}_3 \rightarrow \text{products}$

- CN produced by PLP from ICN
- H atoms observed by LIF

● from $\text{CN} + \text{C}_2\text{H}_2$

● from $\text{CN} + \text{NH}_3$



Yield of H atoms from $\text{CN} + \text{NH}_3 < 5\%$

Gas-Phase Chemistry: role of modelling

- Main aim is to reproduce the observed molecular abundances in different regions of ISM
- Inclusion of errors on the rate coefficients allows one to identify significant discrepancies
- Modelling can also be used to identify 'important reactions'
- In KIDA (Kinetic Database for Astrochemistry) efforts are made to give estimate of uncertainties
- In KIDA, 'data sheets' are provided for important reactions

Gas-Phase Chemistry: some 'take home messages'

- Ion-neutral reactions measured over wide T -range and quite well-understood – it seems that charge exchange be treated by same (Langevin) methods as atom/ion transfers?
- Extension of CRESU experiments to lower temperatures: 5 K?
- Low temperature trapping experiments on radiative association largely restricted to cation + H_2 reactions. No measurements on neutral-neutral associations: e.g. $C + H_2 \rightarrow CH_2 + h\nu$
- More details in IWM Smith, *Ann Rev Astronom Astrophys*, to be published this year.

Radical-Molecule Reactions: CN, C₂H + H₂

$$k(T) = \alpha (T / 300)^\beta \exp(-\gamma / T)$$

	CN + H₂ → HCN + H	C₂H + H₂ → HCCH + H
$\alpha / \text{cm}^3 \text{s}^{-1}$	5.0×10^{-13}	1.95×10^{-12}
β	2.60	2.32
γ / K	960	444
$k(1000 \text{ K}) / \text{cm}^3 \text{s}^{-1}$	4.4×10^{-12}	2.0×10^{-11}

Reported Interstellar and Circumstellar Molecules

N=2	N=2	N=3	N=3	N=4	N = 5	N = 6	N = 7	N = 8	N = 9	N = 10
H ₂	AlCl	H ₃ ⁺	HDO	NH ₃	CH ₄	CH ₃ OH	CH ₃ NH ₂	HCOOCH ₃	(CH ₃) ₂ O	(CH ₃)CO
CH	PN	CH ₂	OCS	H ₃ O ⁺	SiH ₄	CH ₃ SH	CH ₃ CCH	CH ₃ C ₂ CN	C ₂ H ₅ OH	CH ₃ C ₄ CN
CH ⁺	SiN	NH ₂	MgCN	H ₂ CO	CH ₂ NH	C ₂ H ₄	CH ₃ CHO	C ₆ H ₂	C ₂ H ₅ CN	
NH	SiO	H ₂ O	MgNC	H ₂ CS	H ₂ C ₃	CH ₃ CN	c-CH ₂ OCH ₂	C ₇ H	CH ₃ C ₄ H	CH ₃ CH ₂ CH ₂ O
OH	SiS	H ₂ S	NaCN	l-C ₃ H	l-C ₃ H ₂	CH ₃ NC	CH ₂ CHCN	HOCH ₂ CHO	C ₈ H	(CH ₂ OH) ₂
HF	CO ⁺	C ₂ H	SO ₂	c-C ₃ H	c-C ₃ H ₂	H ₂ CCHO	HC ₄ CN	CH ₃ COOH	HC ₆ CN	
C ₂	SO ⁺	HCN	N ₂ O	HCCH	H ₂ CCN	NH ₂ CHO	C ₆ H	H ₂ CCCHCN	CH ₃ CONH ₂	
CN	PO	HNC	SiCN	HCNH ⁺	H ₂ NCN	HC ₃ NH ⁺	H ₂ CCHOH	H ₂ C ₆	C ₈ H ⁻	
CO	SH	HCO	CO ₂	H ₂ CN	CH ₂ CO	H ₂ C ₄	C ₆ H ⁻	CH ₂ CHCHO	CH ₃ CHCH ₂	N = 11
CS	AlF	HCO ⁺	c-SiC ₂	c-C ₃ H	HCOOH	C ₅ H		CH ₂ CCHCN		HC ₈ CN
CP	FeO	HOC ⁺	SiNC	HCCN	C ₄ H	C ₅ N				HCOOC ₂ H ₅
NO	SiC	HN ₂ ⁺	AlNC	HNCO	HC ₂ CN	HC ₄ H				N = 12
NS	CF ⁺	HNO	HCP	HOCO ⁺	HC ₂ NC	HC ₃ CN				C ₆ H ₆
SO	N ₂	HCS ⁺	C ₂ P	HNCS	C ₄ Si	c-C ₃ H ₂ O				C ₃ H ₇ CN
HCl	LiH	C ₃	AlOH	C ₂ CN	C ₅	H ₂ CCHNH				N = 13
NaCl	SiH	C ₂ O	H ₂ O ⁺	C ₃ O C ₃ S	C ₄ N	C ₅ N ⁻				HC ₁₀ CN
KCl AlO	O ₂	C ₂ S	H ₂ Cl ⁺	SiC ₃ PH ₃	H ₂ COH ⁺					C ₂ H ₅ OCH ₃
ClH ⁺	CN ⁻			C ₂ N ⁻	C ₂ H ⁻					

CRESU apparatus configured for radical-neutral reactions

