

# The NGTS Survey Transiting Neptunes & Super-Earths

Daniel Bayliss  
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# Overview

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- In coordination with François Bouchy & Christophe Lovis at the Observatory of Geneva (WG A1).

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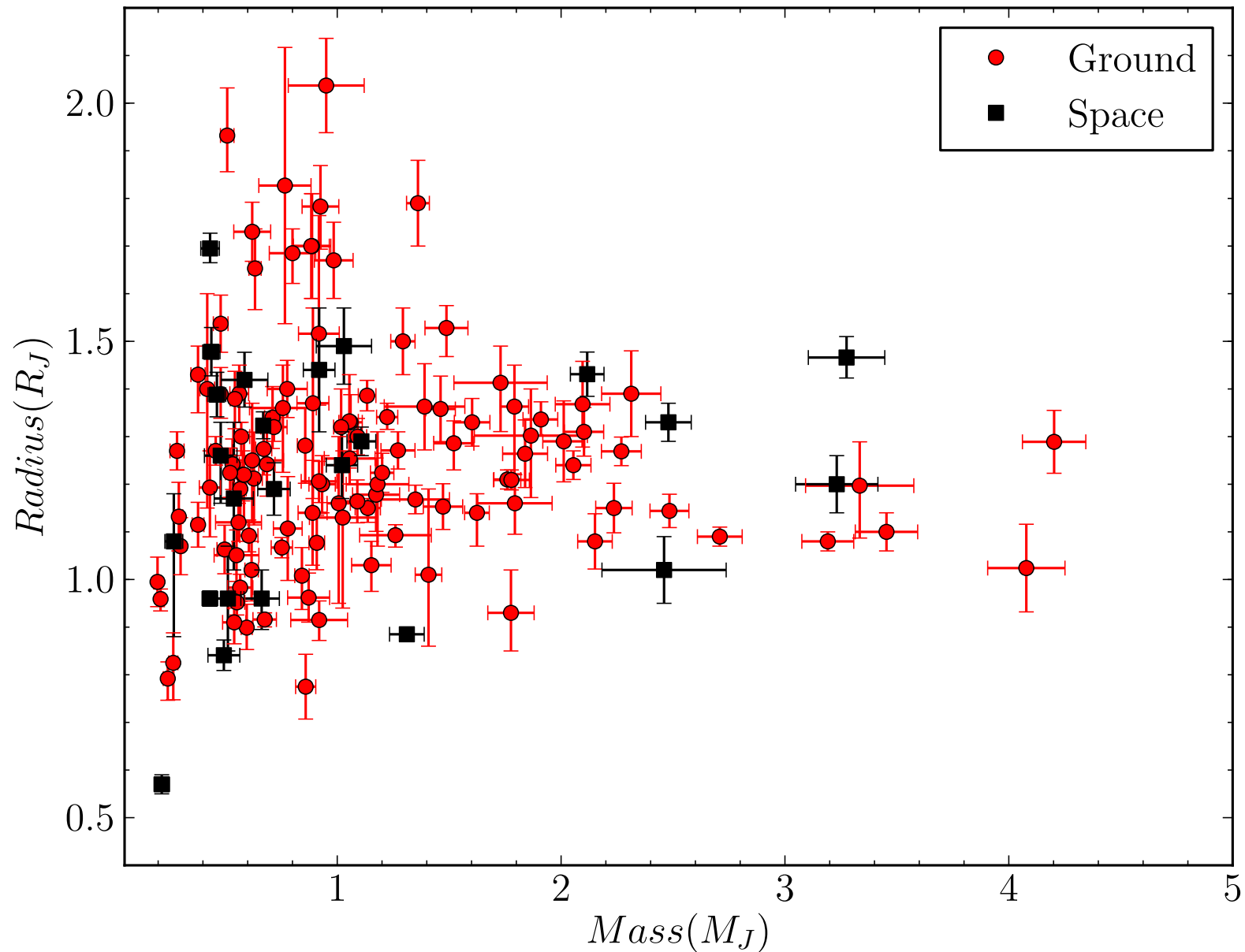
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- Focus primarily on transiting gas-planets (Neptune to Jupiter radii).
- Discuss more than only NGTS: current state of field and prospects over the next 2-3 years in lead-up to CHEOPS launch.

# Current State

# Transiting “giant” planets



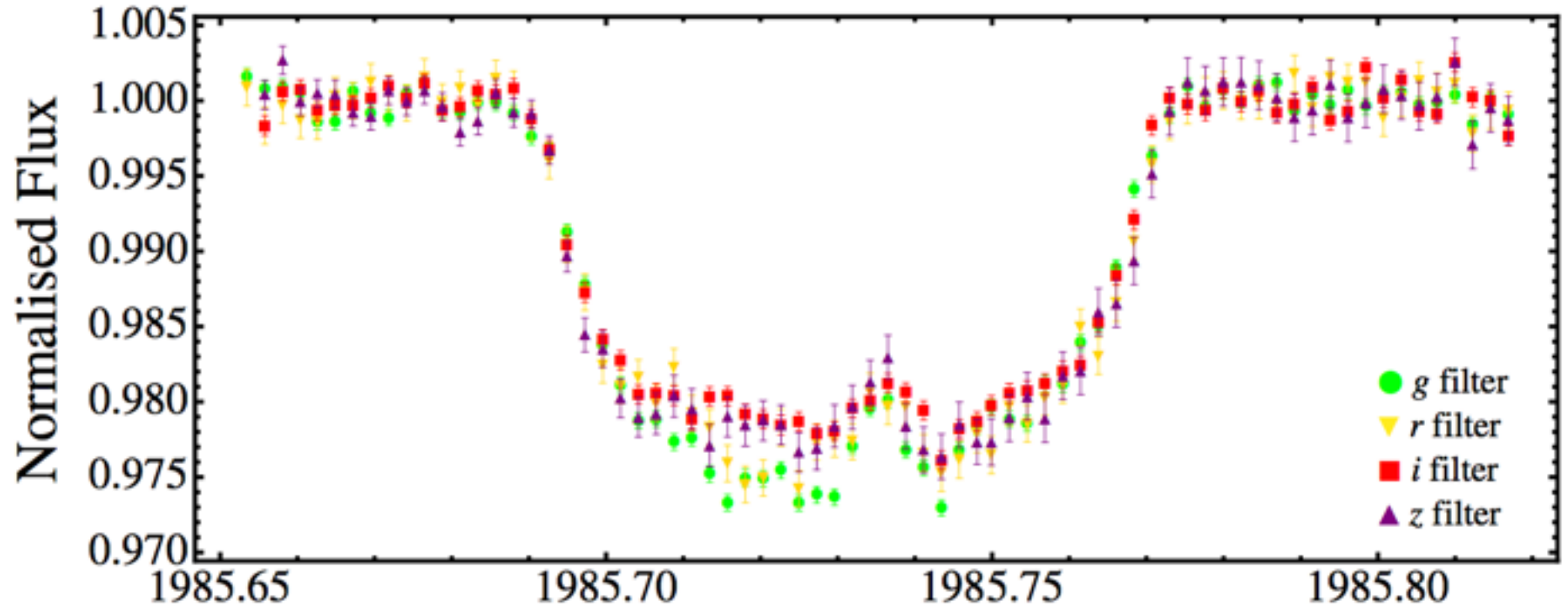
# Transiting exoplanet discoveries

- Currently driven by WASP/SuperWASP (87 TEPs) and HATNet/HATSouth (55 TEPs). Also KELT for bright stars (5 TEPs)

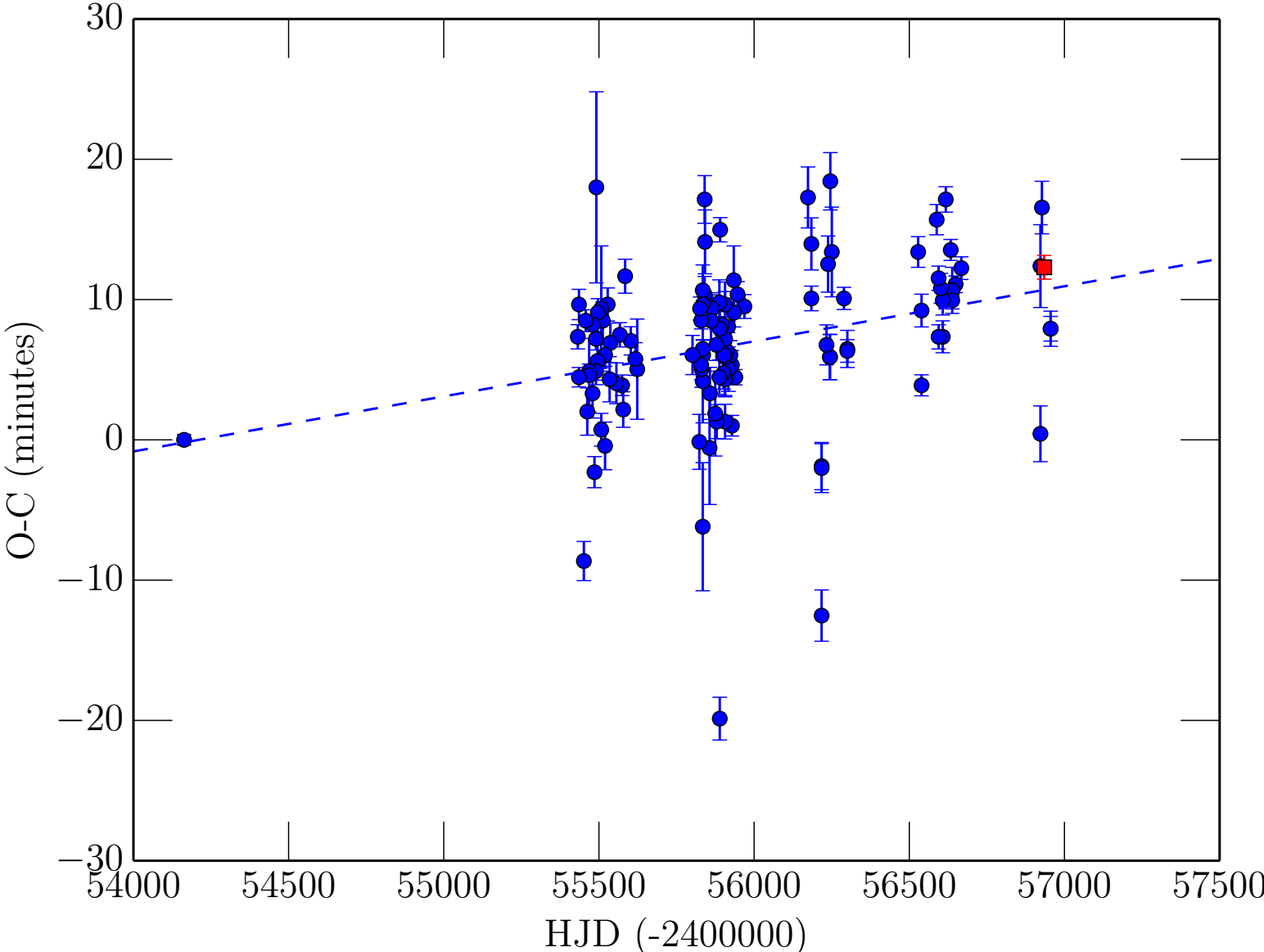




# Star-spots (e.g. HATS-2b)

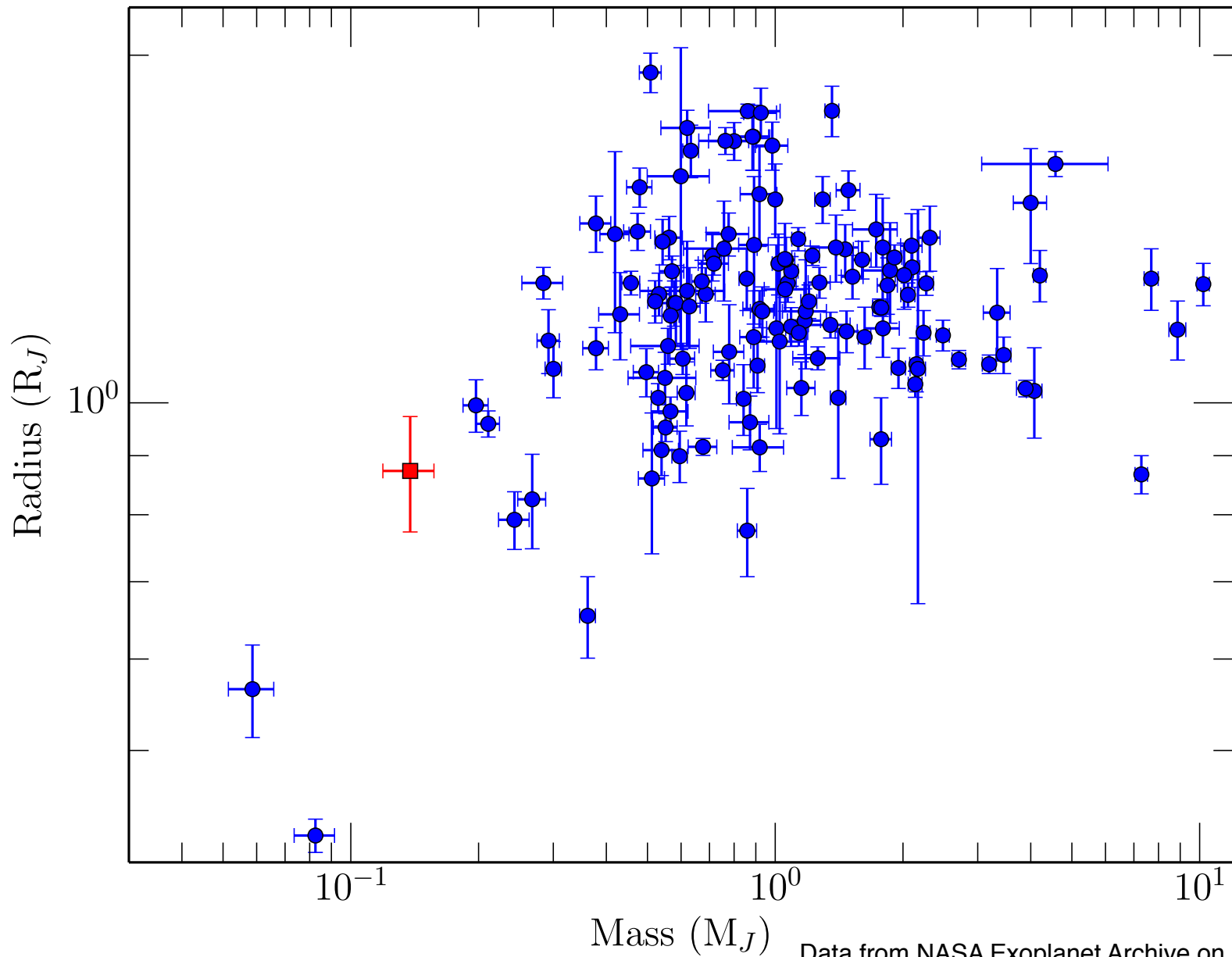


# Long-baseline transit timing variations

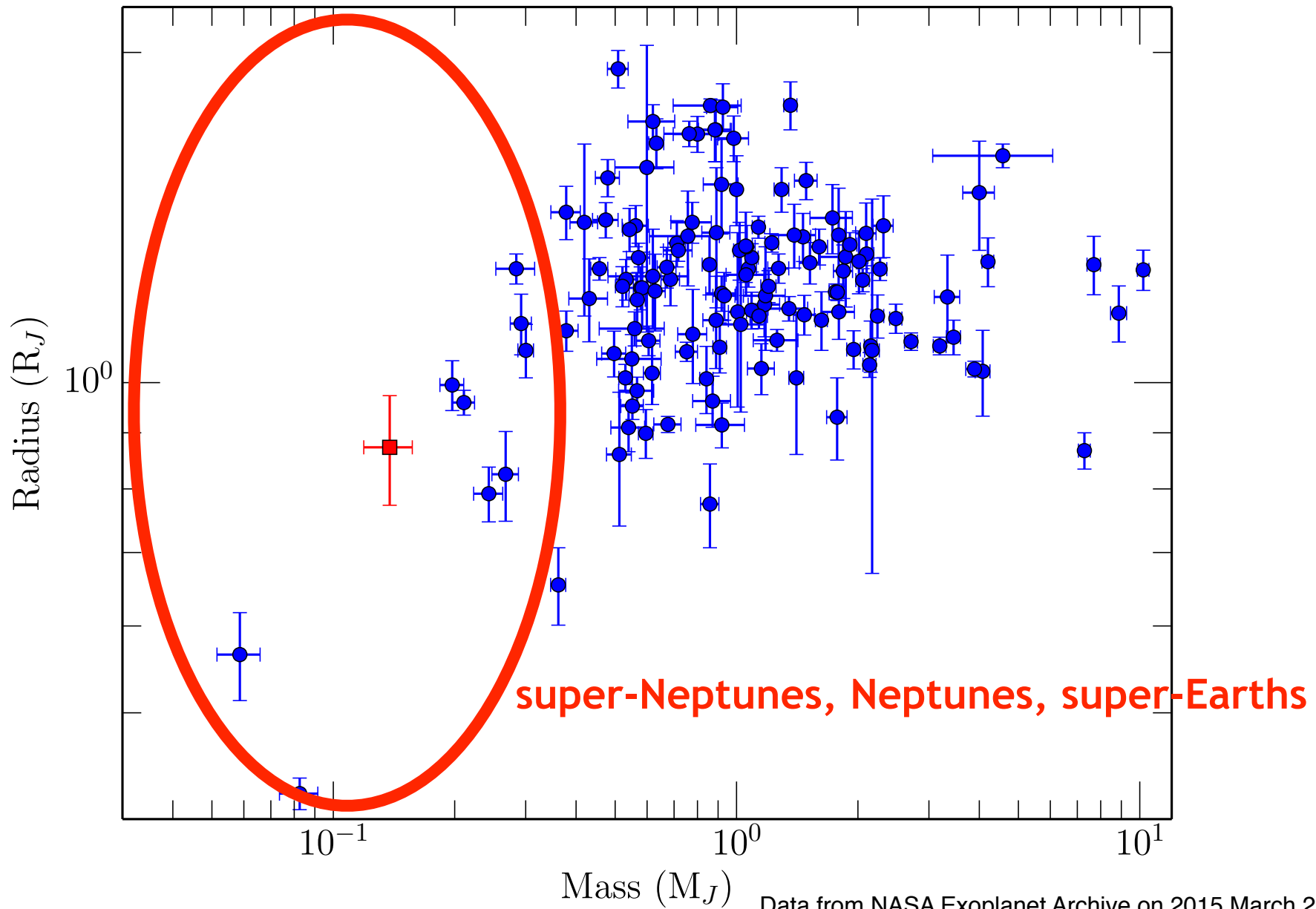


# Future Prospects (2-3 years)

# Ground-based Detections



# Ground-based Detections



# HATSouth - Improvements:

- larger telescopes (20cm)
- more telescopes (24 units)
- global network (3 sites)

HAT Surveys – HATNet – HATSouth

 **PRINCETON UNIVERSITY**

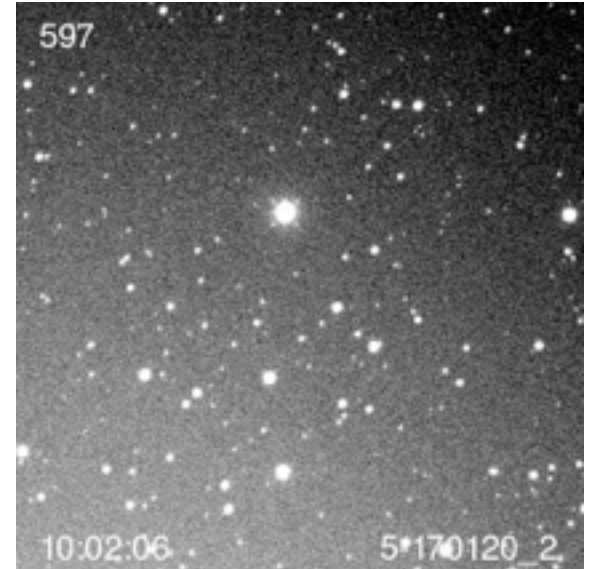


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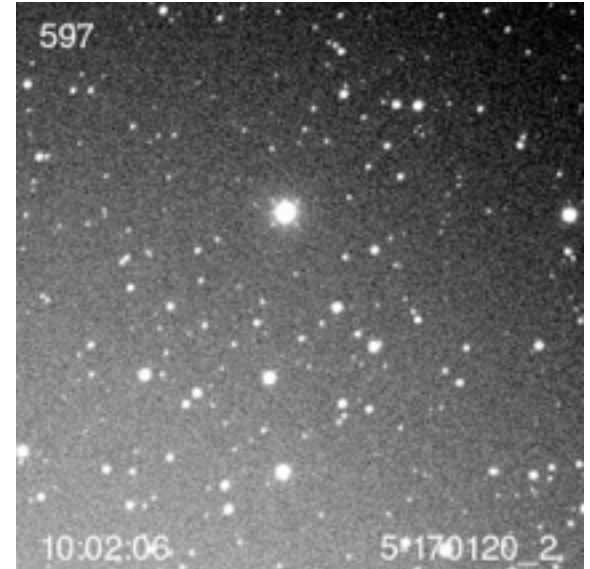


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# HATSouth



# HATSouth





# Results to Date

Planet name	Period [days]	Semimajor axis [AU]	Eccentricity	Mass [ $M_{\text{Jupiter}}$ ]	Radius [ $R_{\text{Jupiter}}$ ]	Discovery paper
HATS-14b	2.76676	0.03815	< 0.142	1.071	1.039	<a href="#">Mancini+ 2015</a>
HATS-13b	3.04405	0.04057	< 0.181	0.543	1.212	<a href="#">Mancini+ 2015</a>
HATS-10b	3.31285	0.04491	< 0.501	0.526	0.969	<a href="#">Brahm+ 2015</a>
HATS-9b	1.91531	0.03048	< 0.129	0.837	1.065	<a href="#">Brahm+ 2015</a>
HATS-8b	3.58389	0.04667	< 0.376	0.138	0.873	<a href="#">Bayliss+ 2015</a>
HATS-6b	3.32527	0.03621	0 (fixed)	0.333	1.000	<a href="#">Hartman+ 2014</a>
HATS-5b	4.76339	0.0542	0.019	0.237	0.912	<a href="#">Zhou+ 2014</a>
HATS-4b	2.51673	0.0362	0.013	1.323	1.020	<a href="#">Jordan+ 2014</a>
HATS-3b	3.57485	0.0485	0 (fixed)	1.071	1.381	<a href="#">Bayliss+ 2013</a>
HATS-2b	1.13541	0.0230	0 (fixed)	1.345	1.168	<a href="#">Mohler-Fischer+ 2013</a>
HATS-1b	3.44646	0.0444	0.120	1.855	1.302	<a href="#">Penev+ 2013</a>

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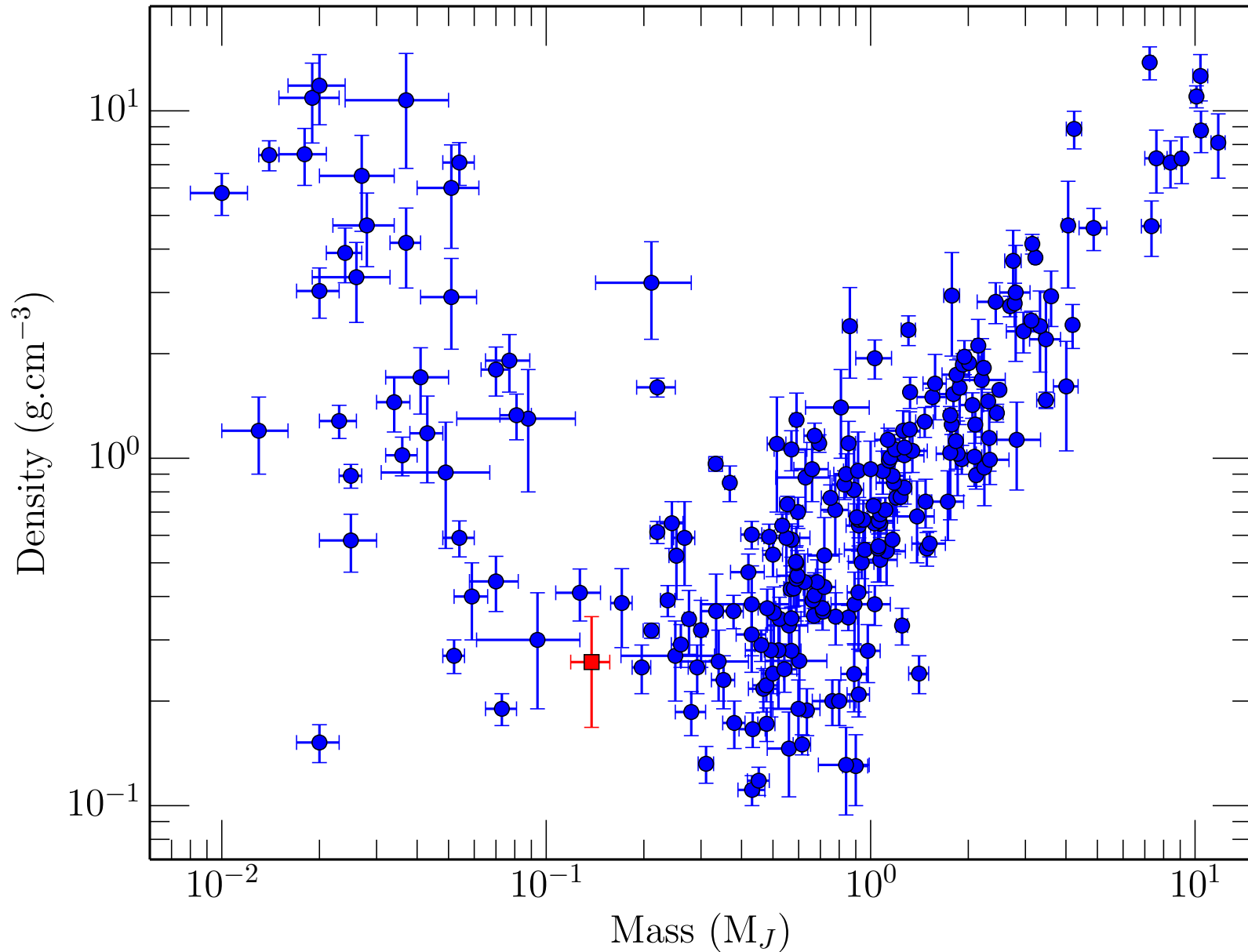
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# Exoplanet Densities



# NGTS - Improvements

- Larger telescopes (20cm).
- High quality CCDs (red sensitive).
- Individual mounts and high precision tracking/  
guiding (stars stay on same pixel).
- More telescopes (12 units).
- Excellent site (Paranal).

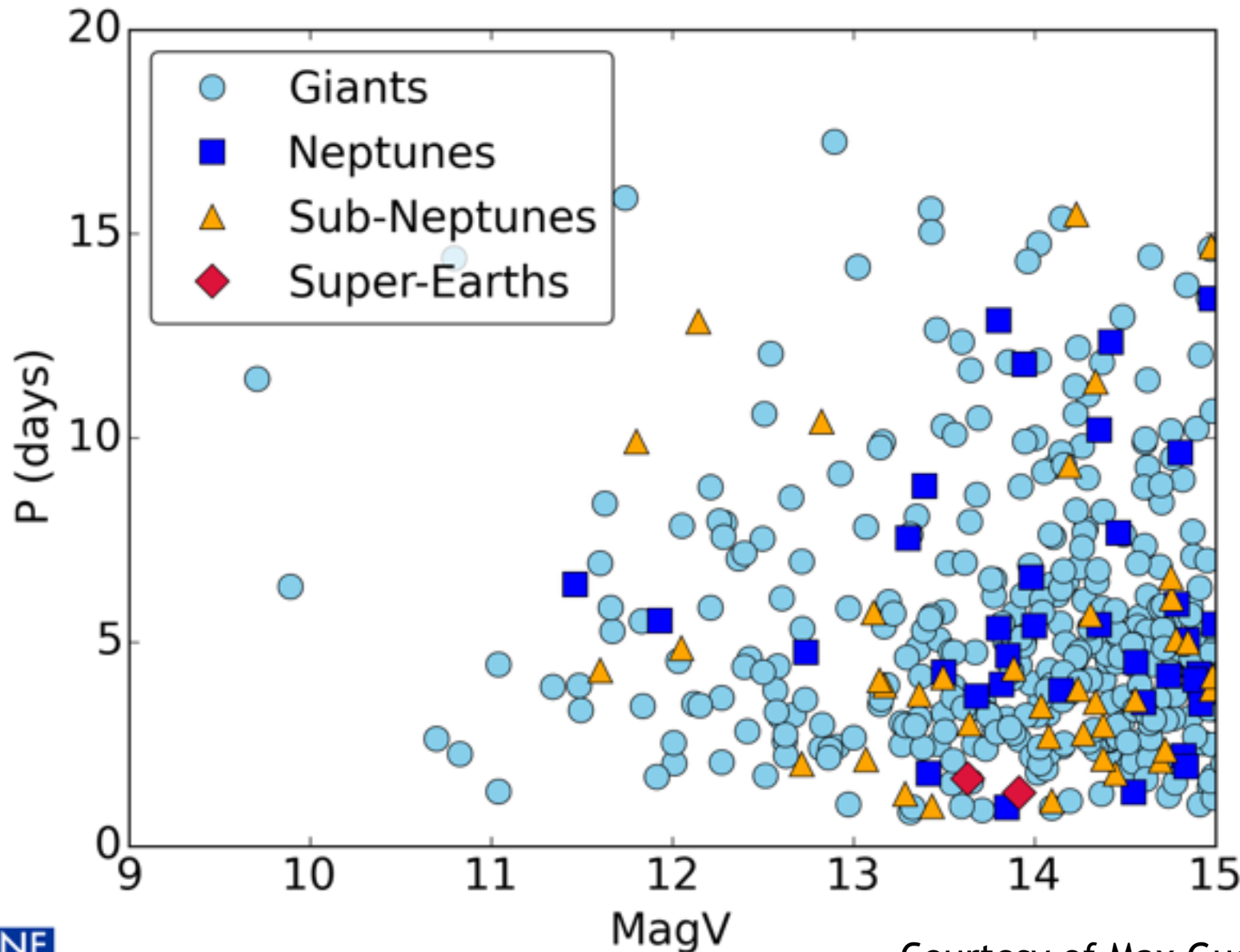








# NGTS Yields



# Summary

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- We need to think about what CHEOPS can do for existing TEPs around bright stars: TTV, star-spots, moons/rings, unseen companions, visible light variations, oblateness, etc
- Many interesting systems may be discovered below  $V=12$ . What can CHEOPS do for these?