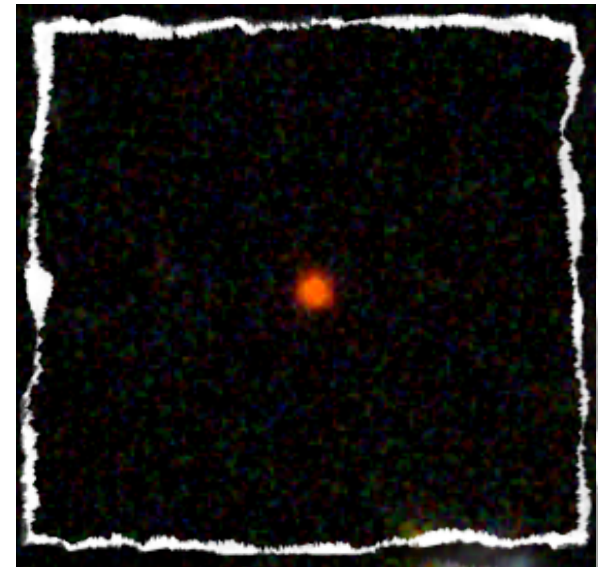
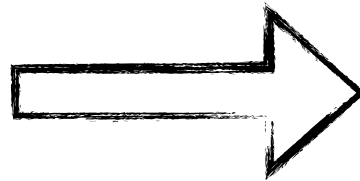
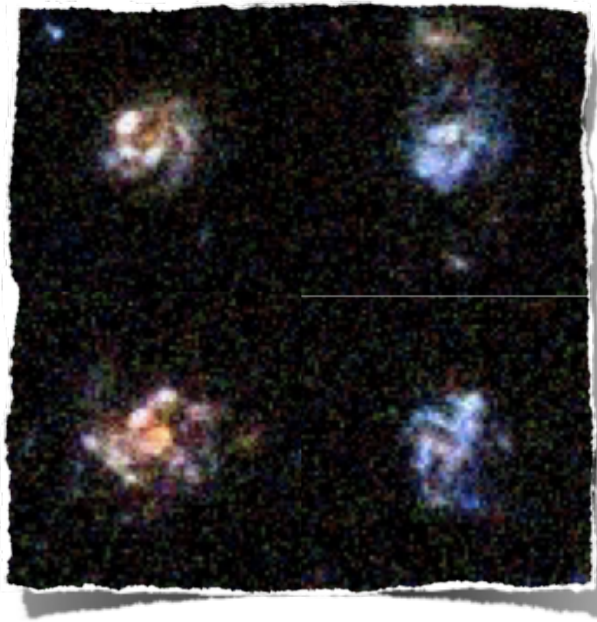


Live fast, die... small:

The progenitors of
the first quiescent galaxies



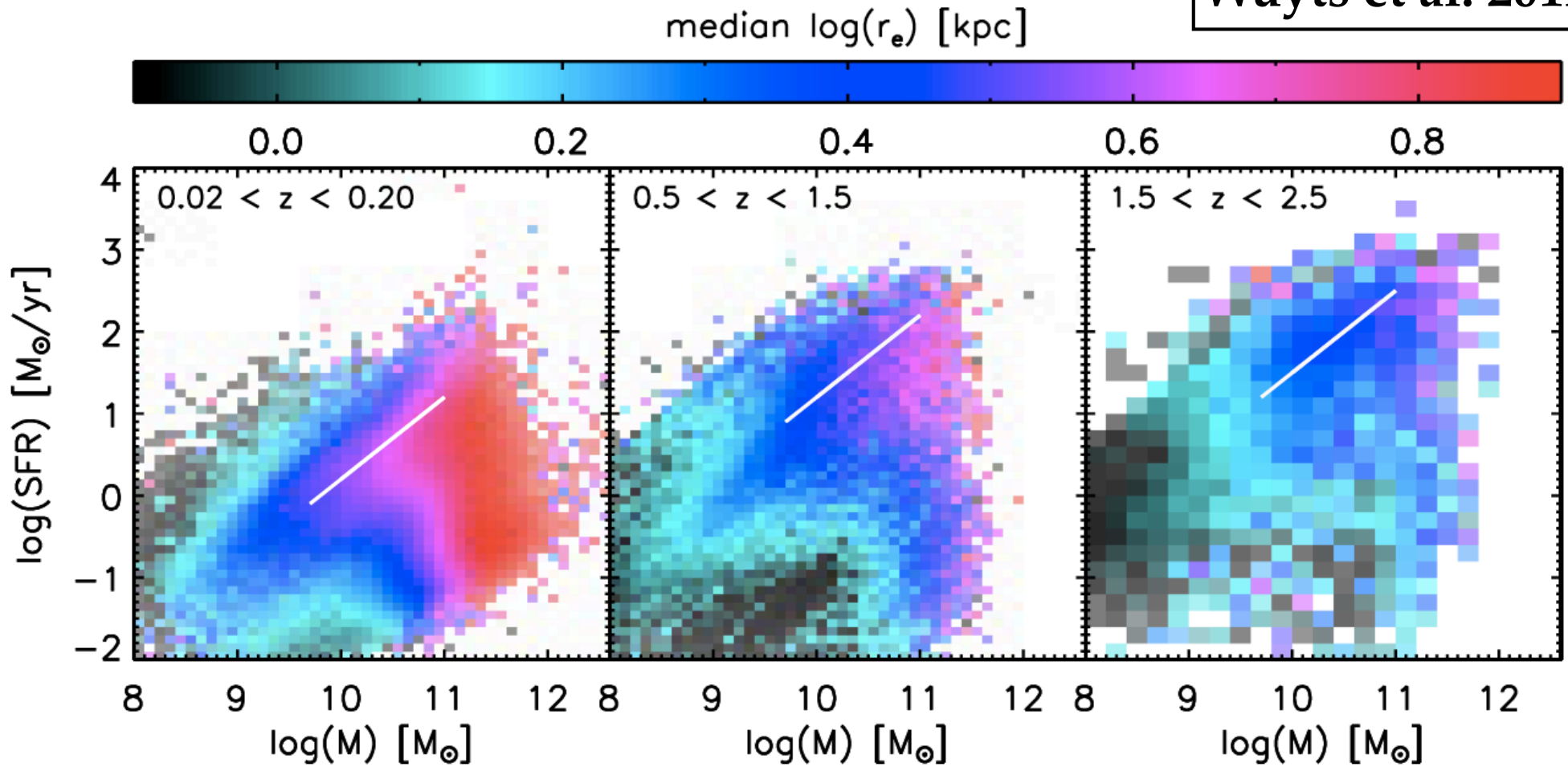
Guillermo Barro (UCSC)

S. Faber, P. Perez-Gonzalez, D. Koo, J. Trump, D. Kocevski, E. McGrath, L. Porter, J. Primack, C. Pacifici, C. Moody, P. Kollipara, A. van der Wel, S. Wuyts + CANDELS

June 21th 2013, UCM - SHARDS team meeting

The SFR-M plane

Wuyts et al. 2011

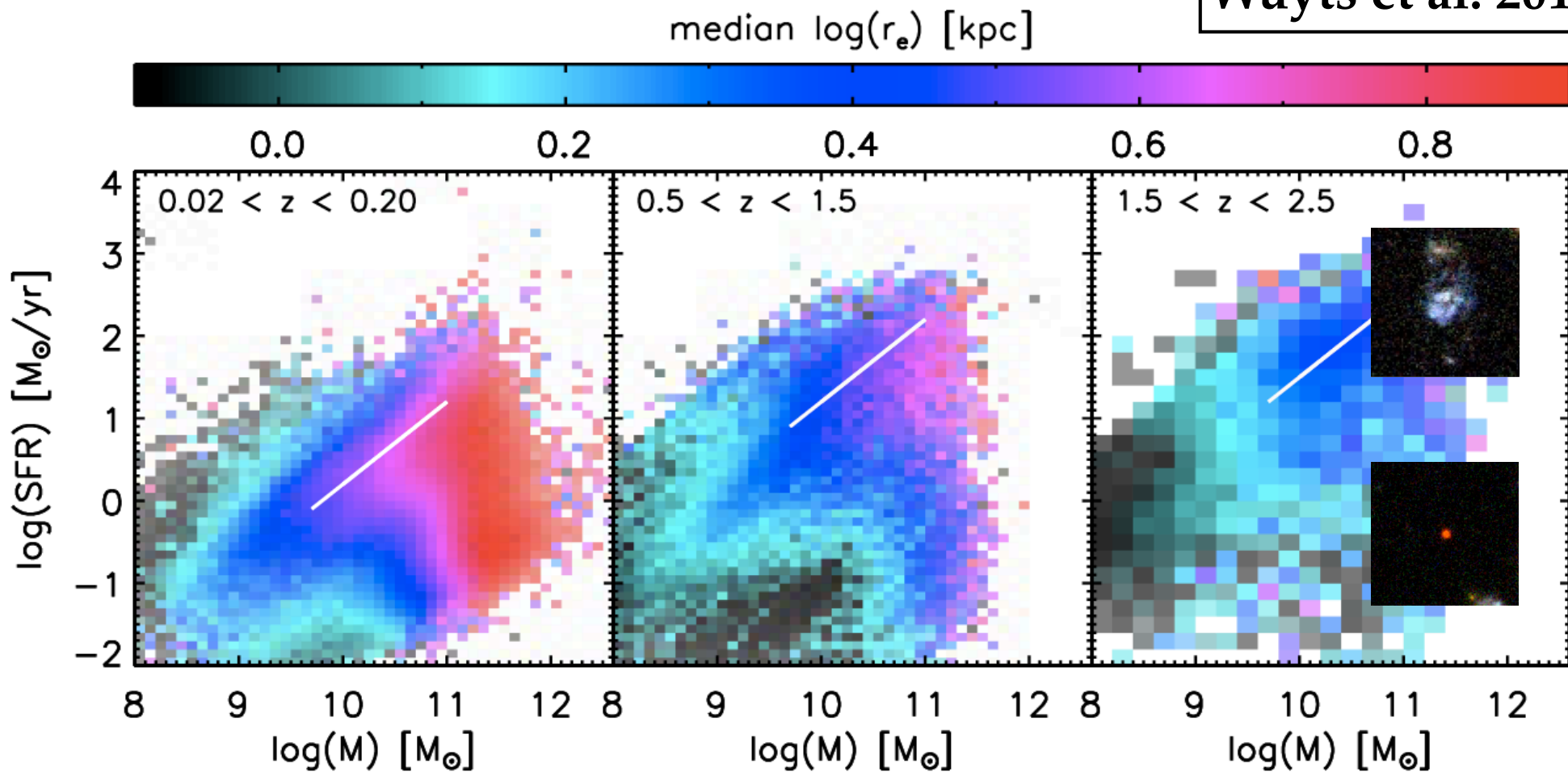


$$\text{SFR} \approx M^{\alpha}$$

- SFGs are bigger at a given mass

The SFR-M plane

Wuyts et al. 2011

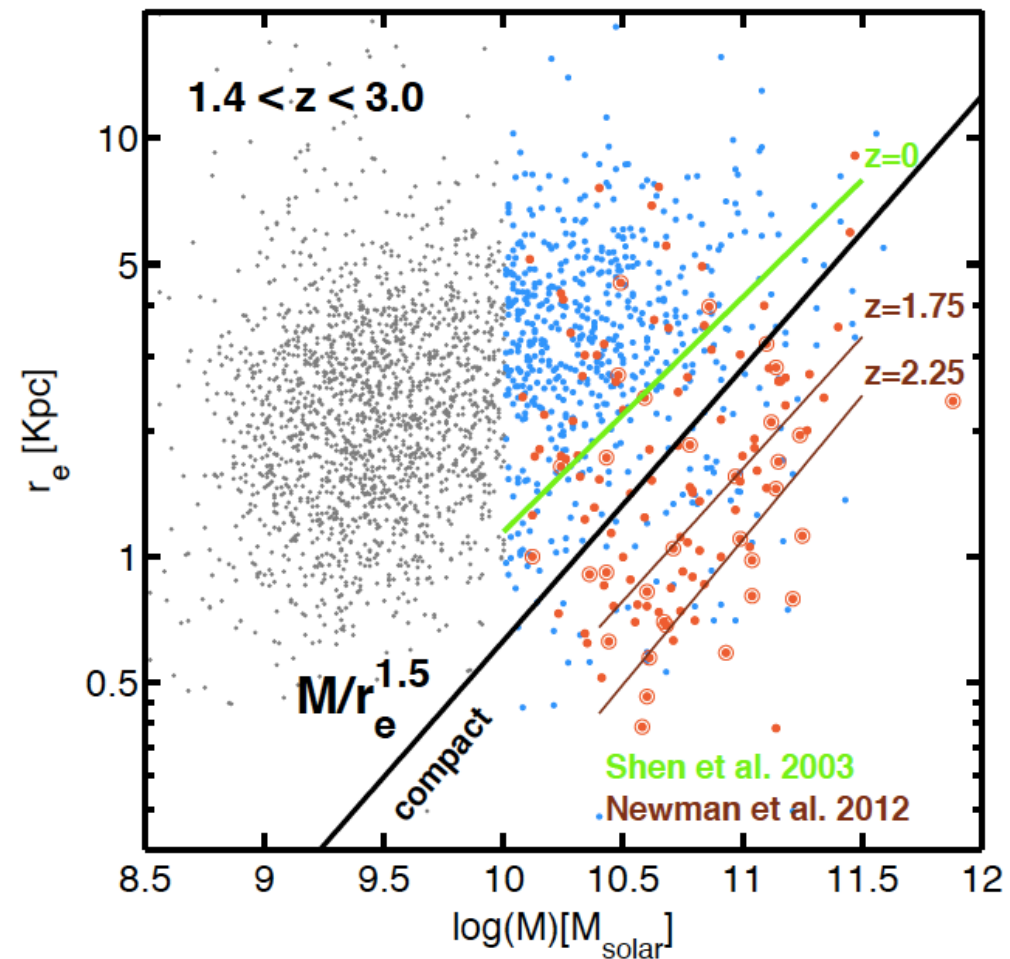
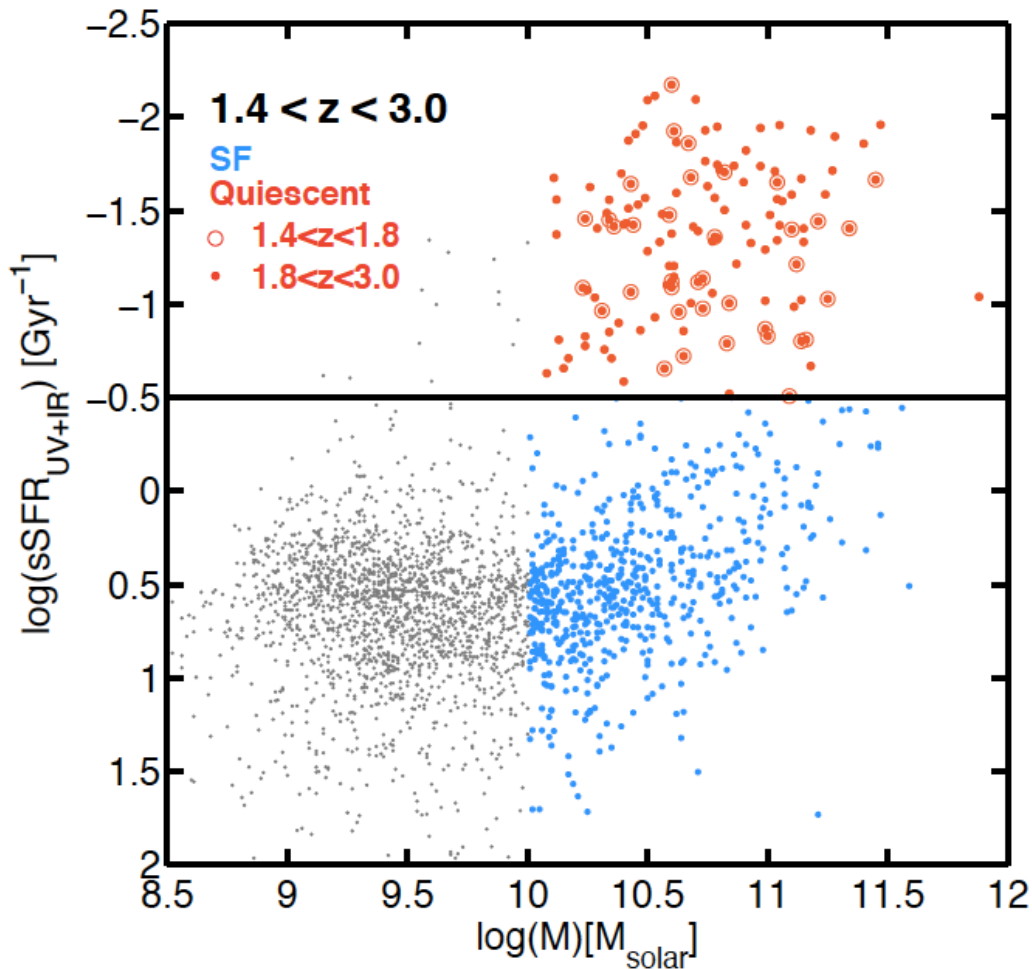


$$\text{SFR} \approx M^{\alpha}$$

- SFGs are bigger at a given mass

sSFR-M and Mass-Size

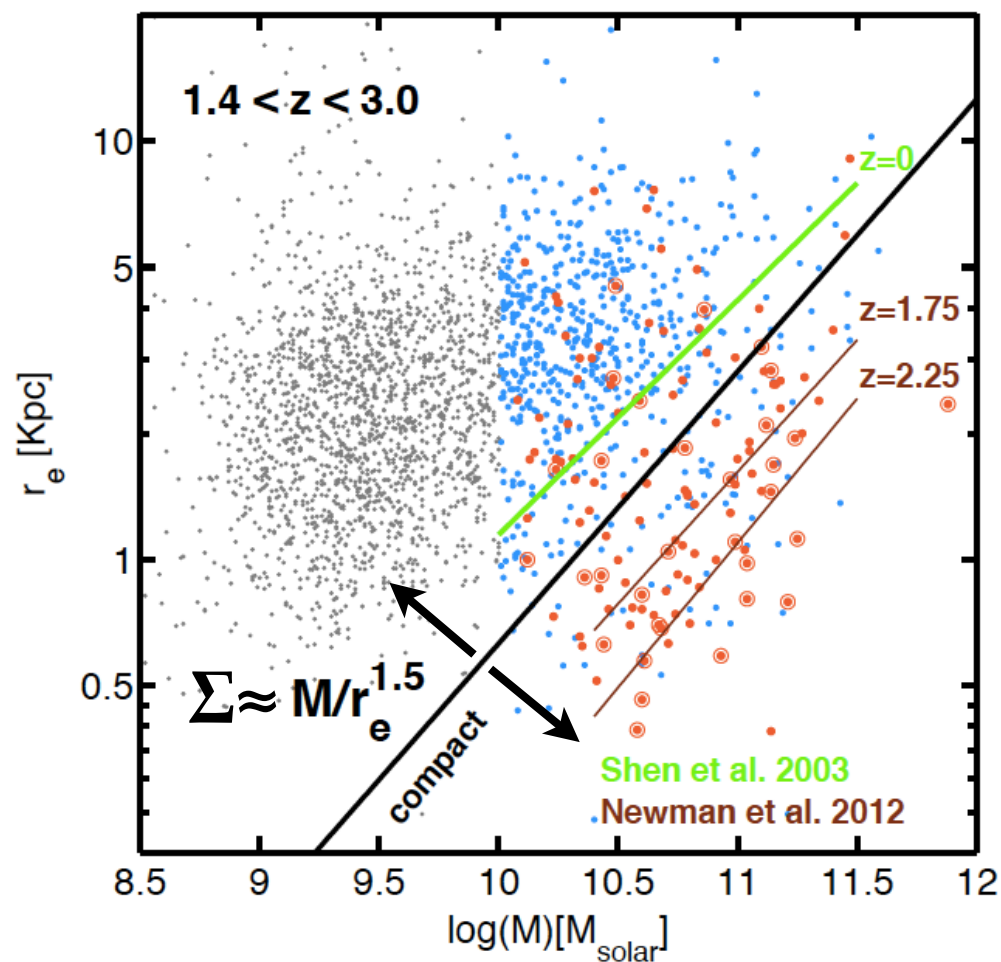
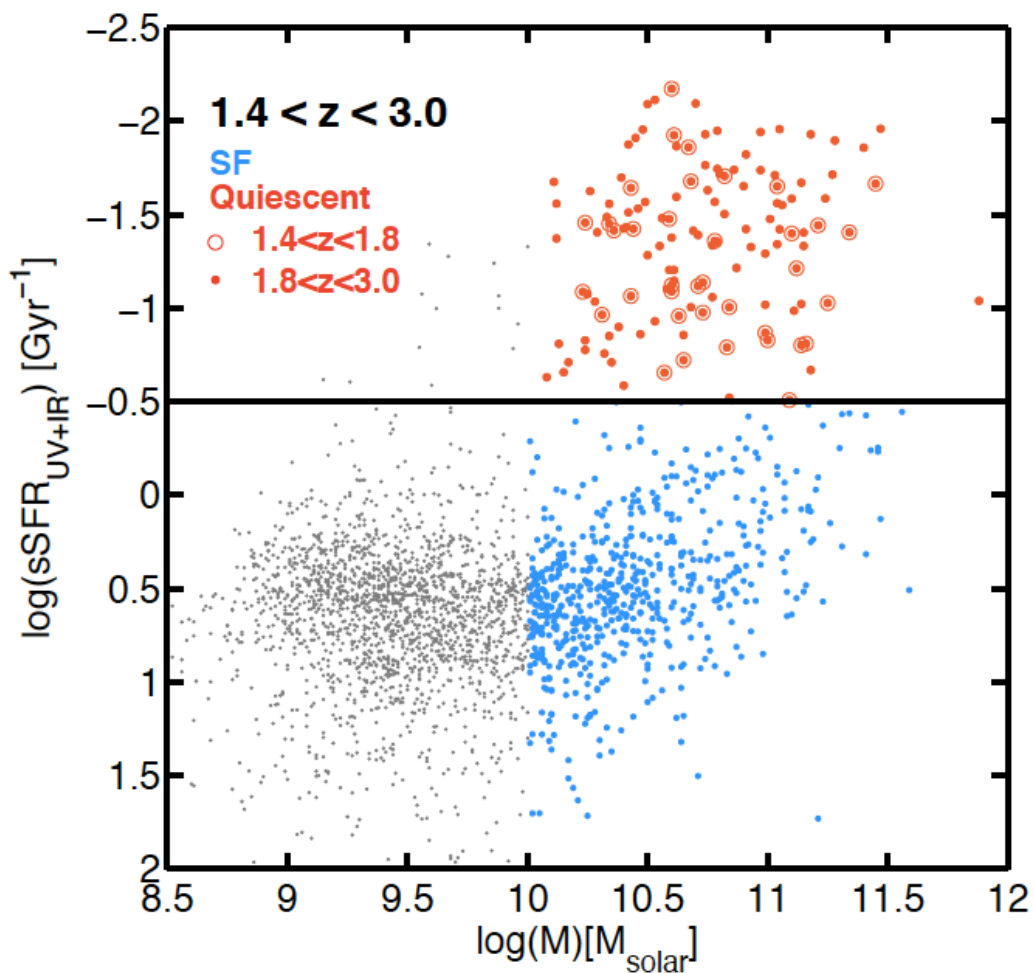
Barro et al. 2013

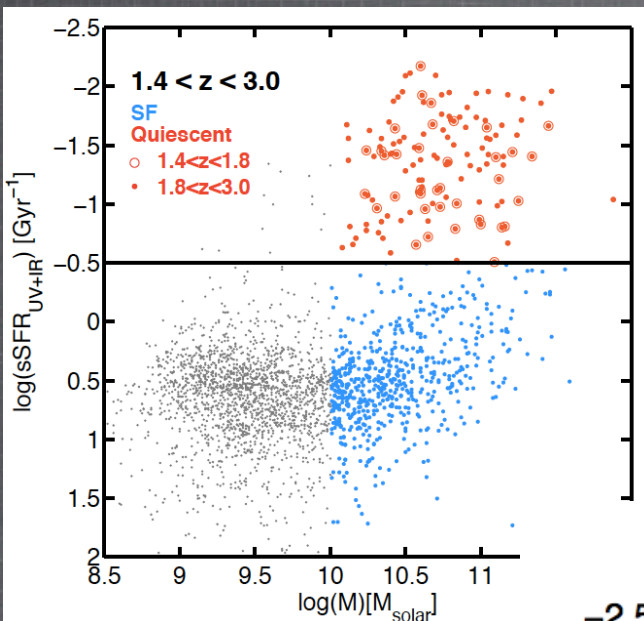


- CANDELS H-band selected in GOODS-S & UDS , $\log(M) > 10$
- Photo-z's (spec-z), stellar masses, (UV+IR) SFRs, GALFIT morphologies

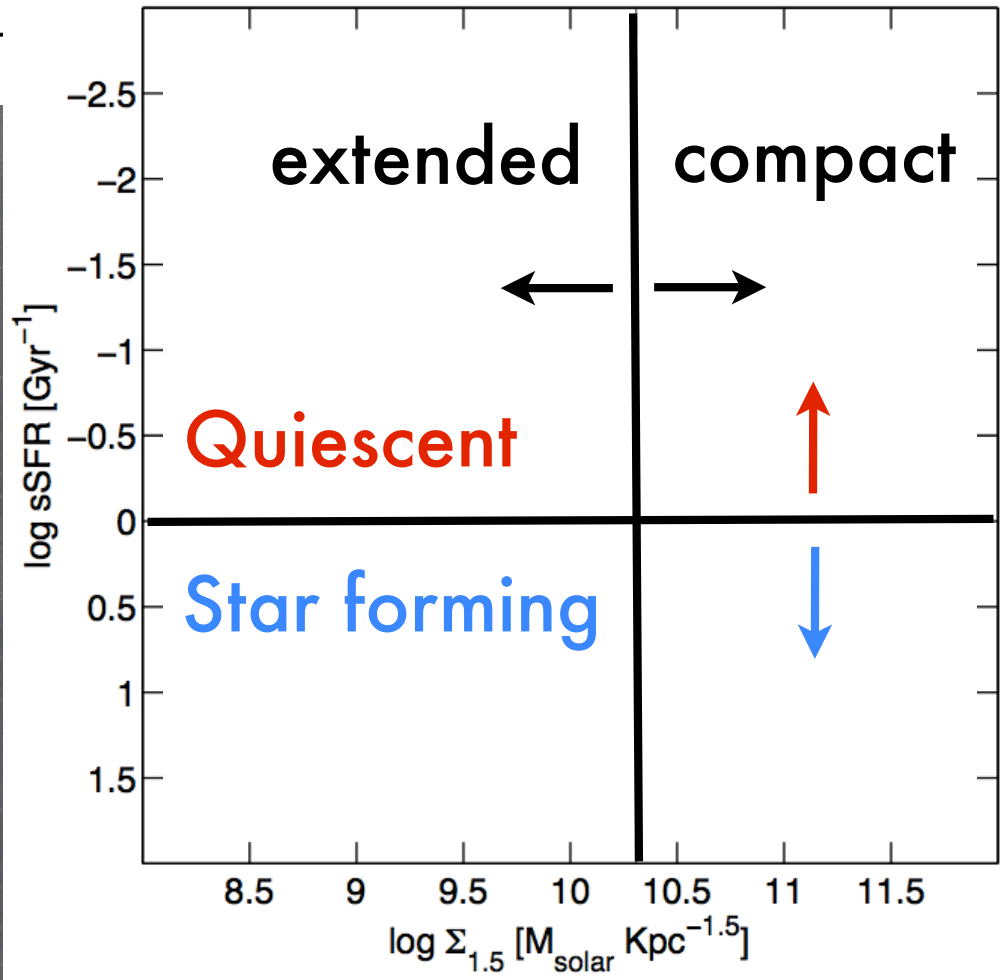
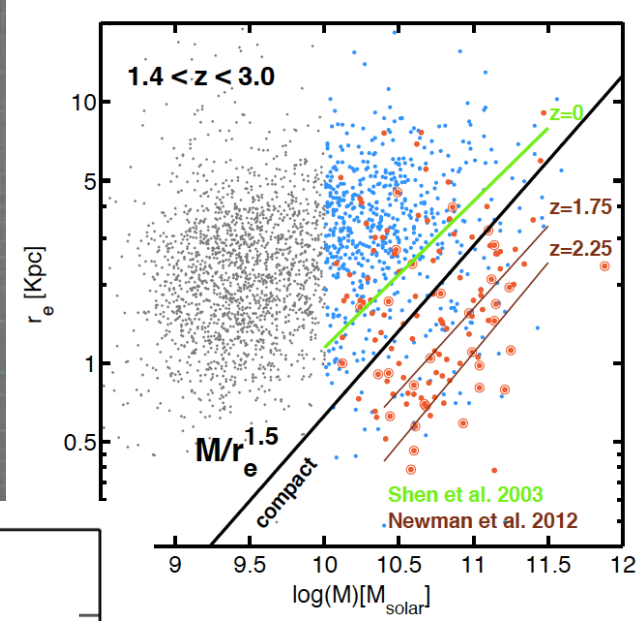
sSFR-M and Mass-Size

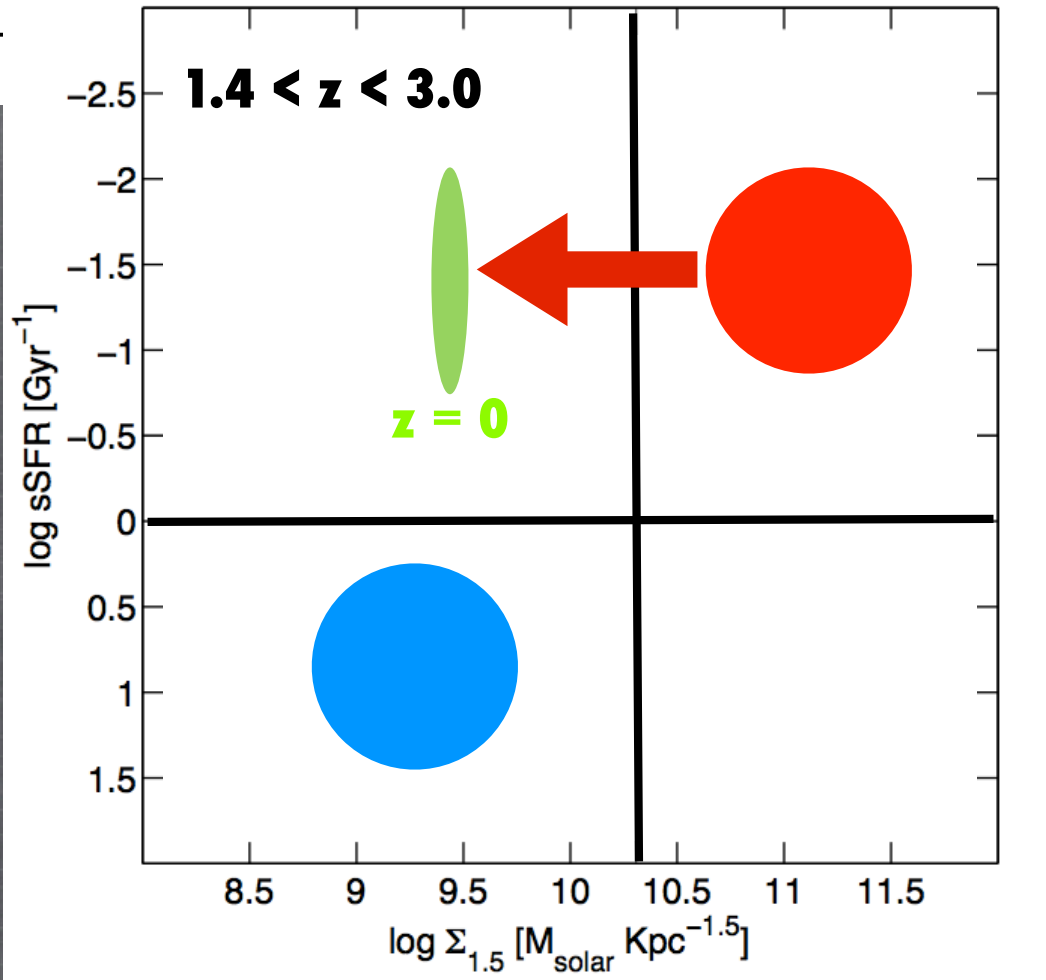
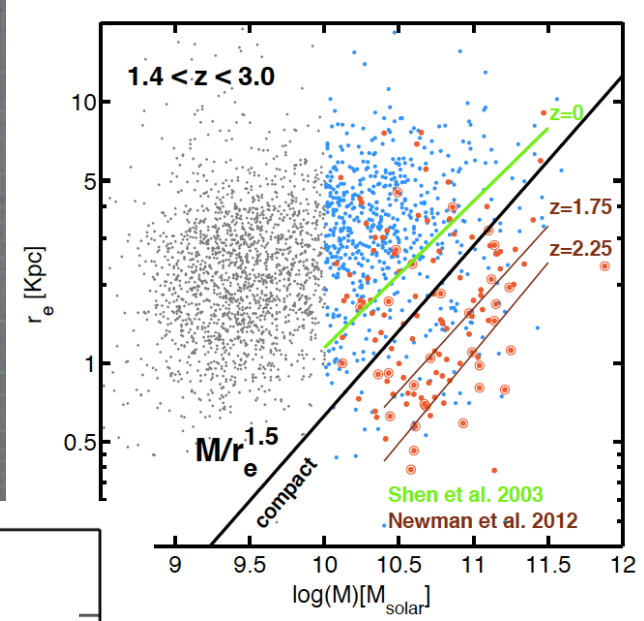
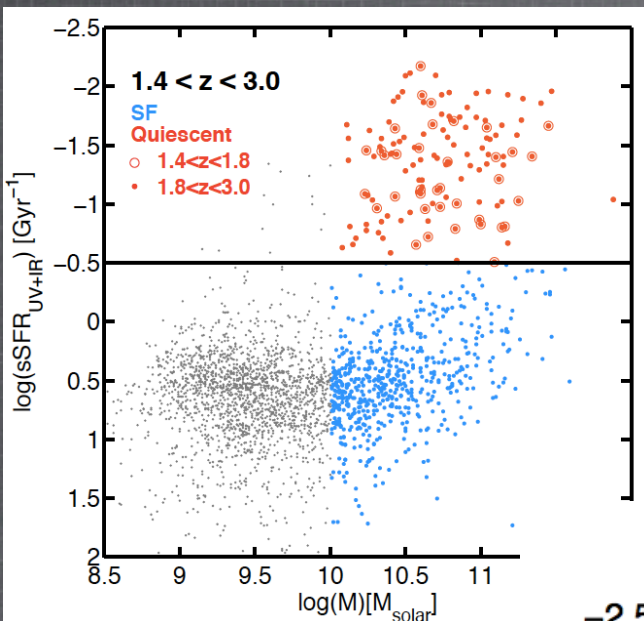
Barro et al. 2013

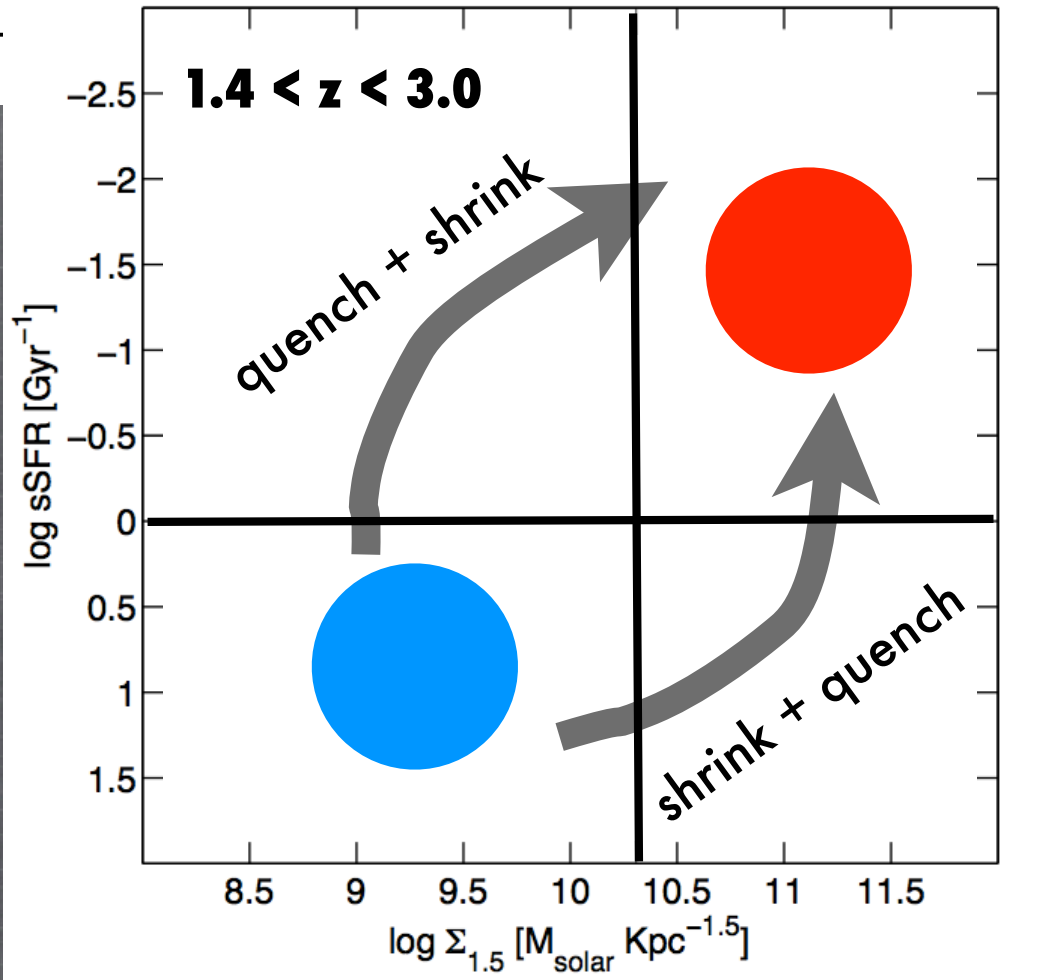
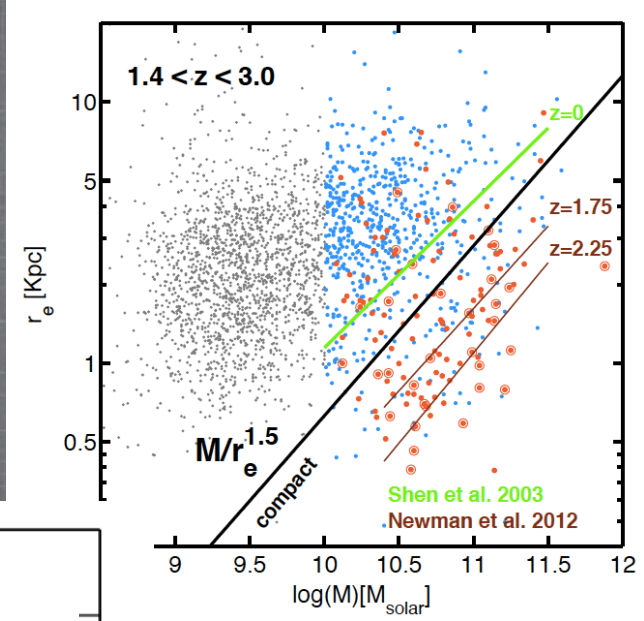
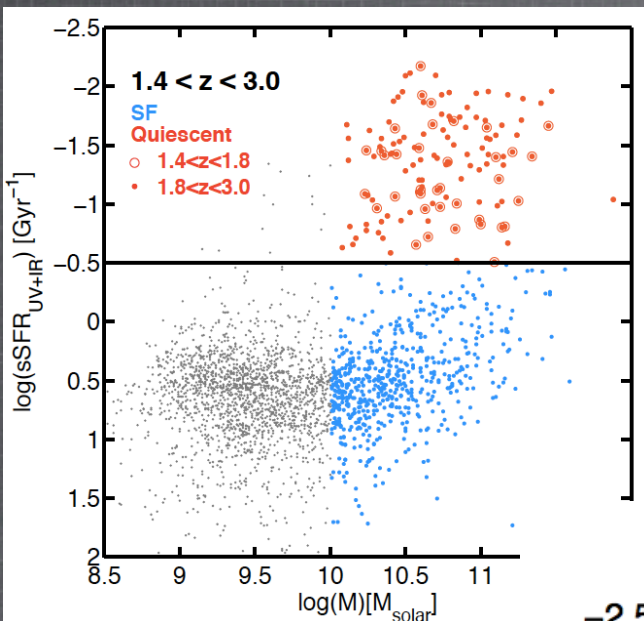


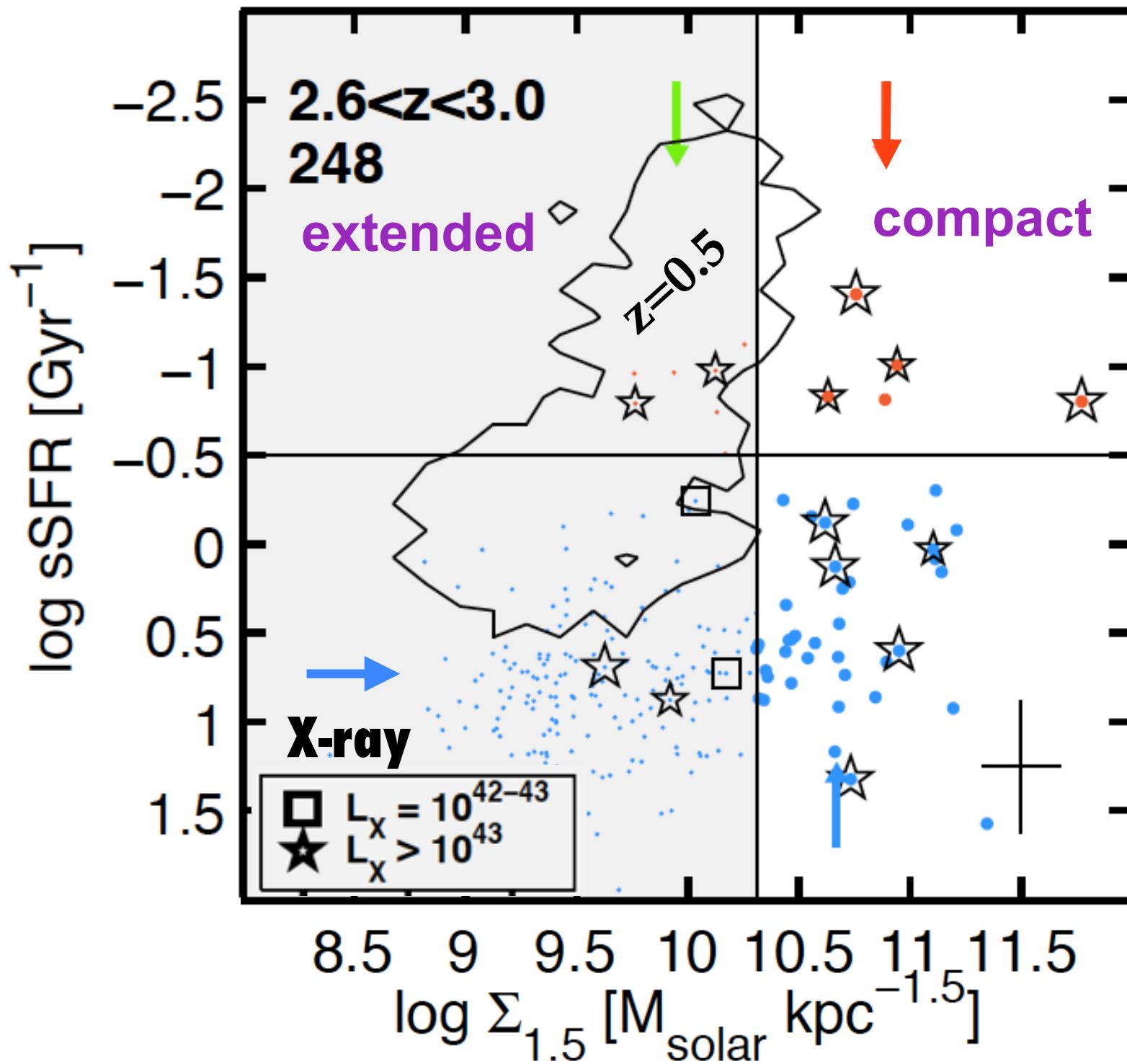


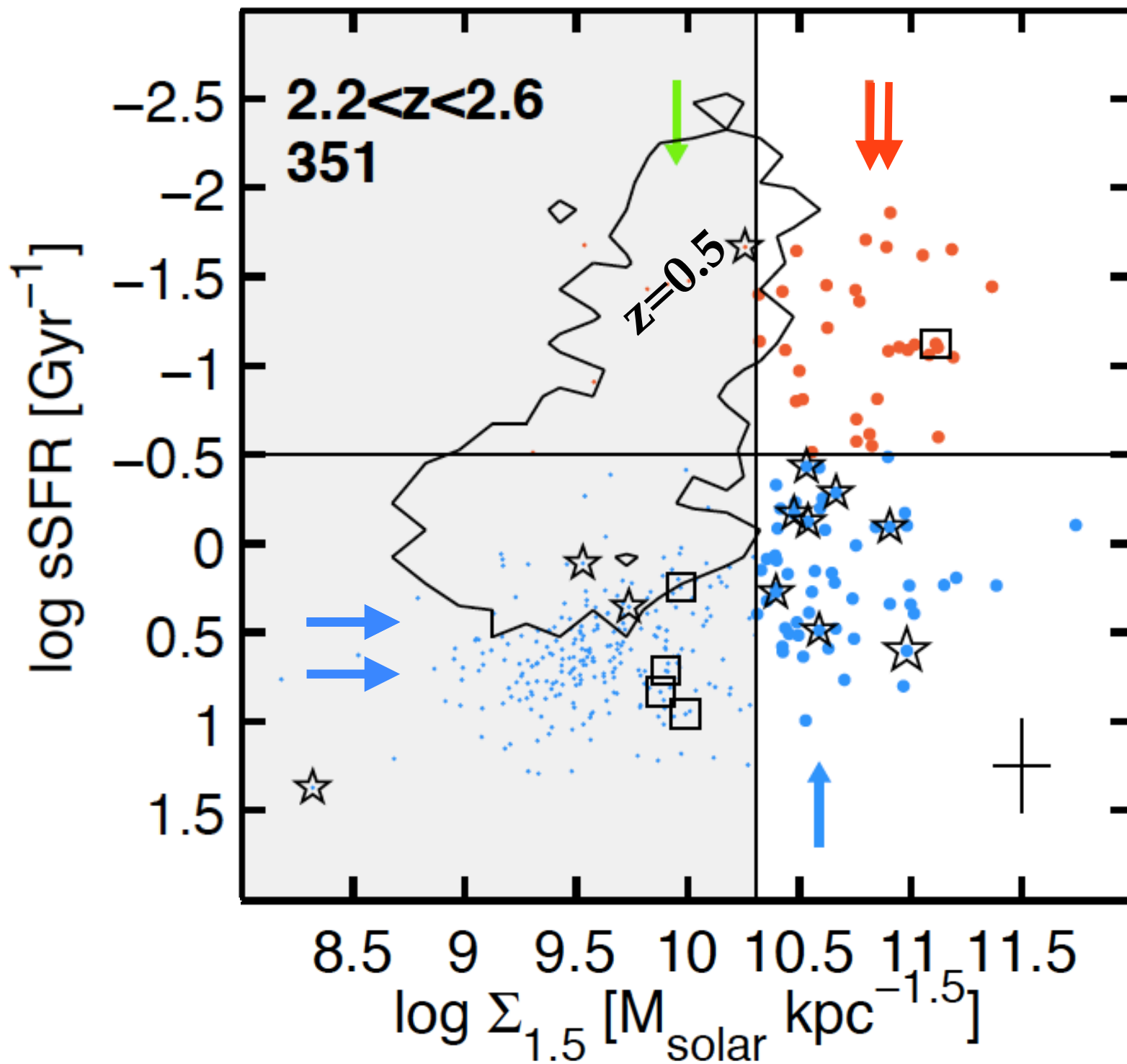
$1.4 < z < 3.0$

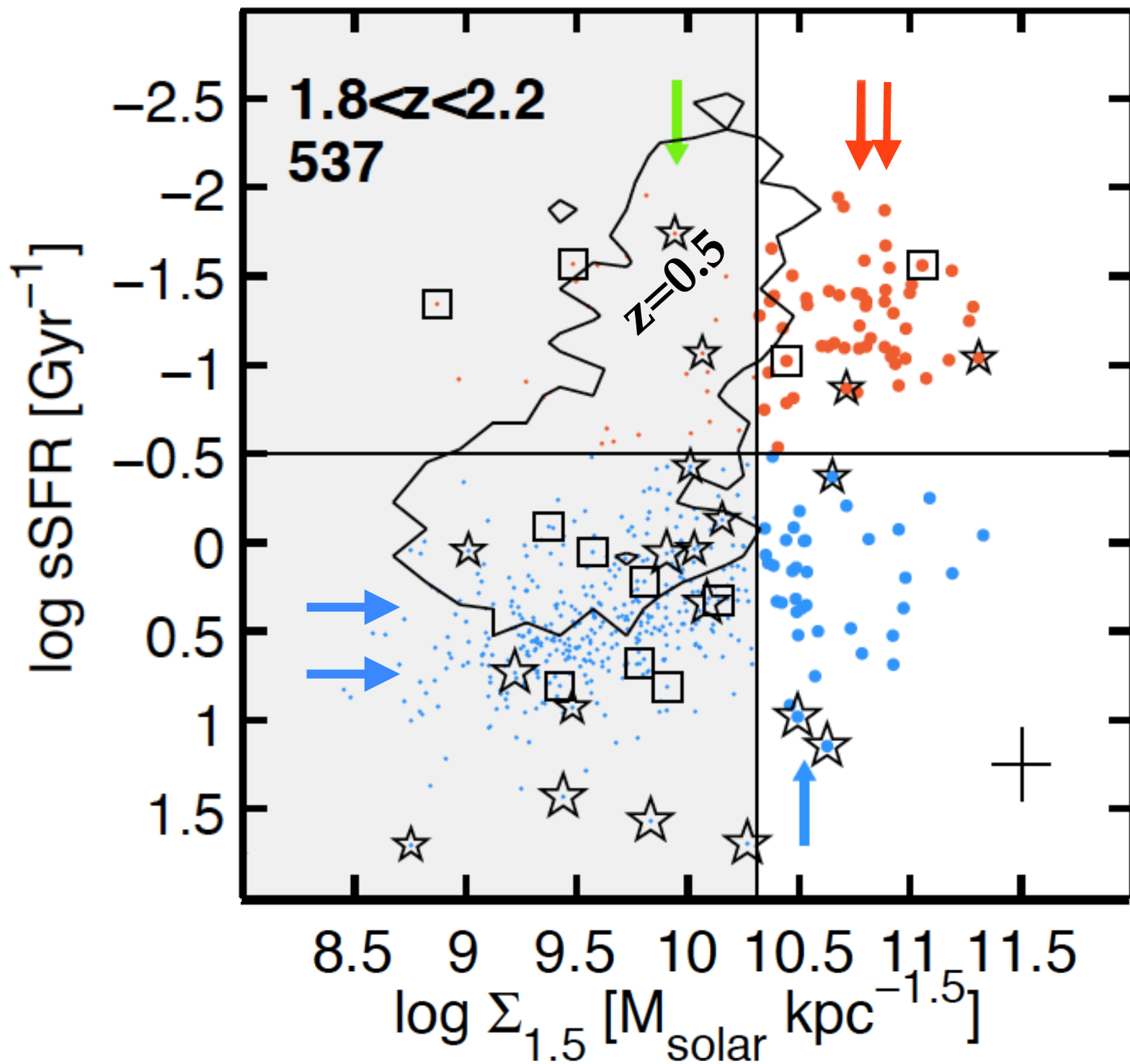


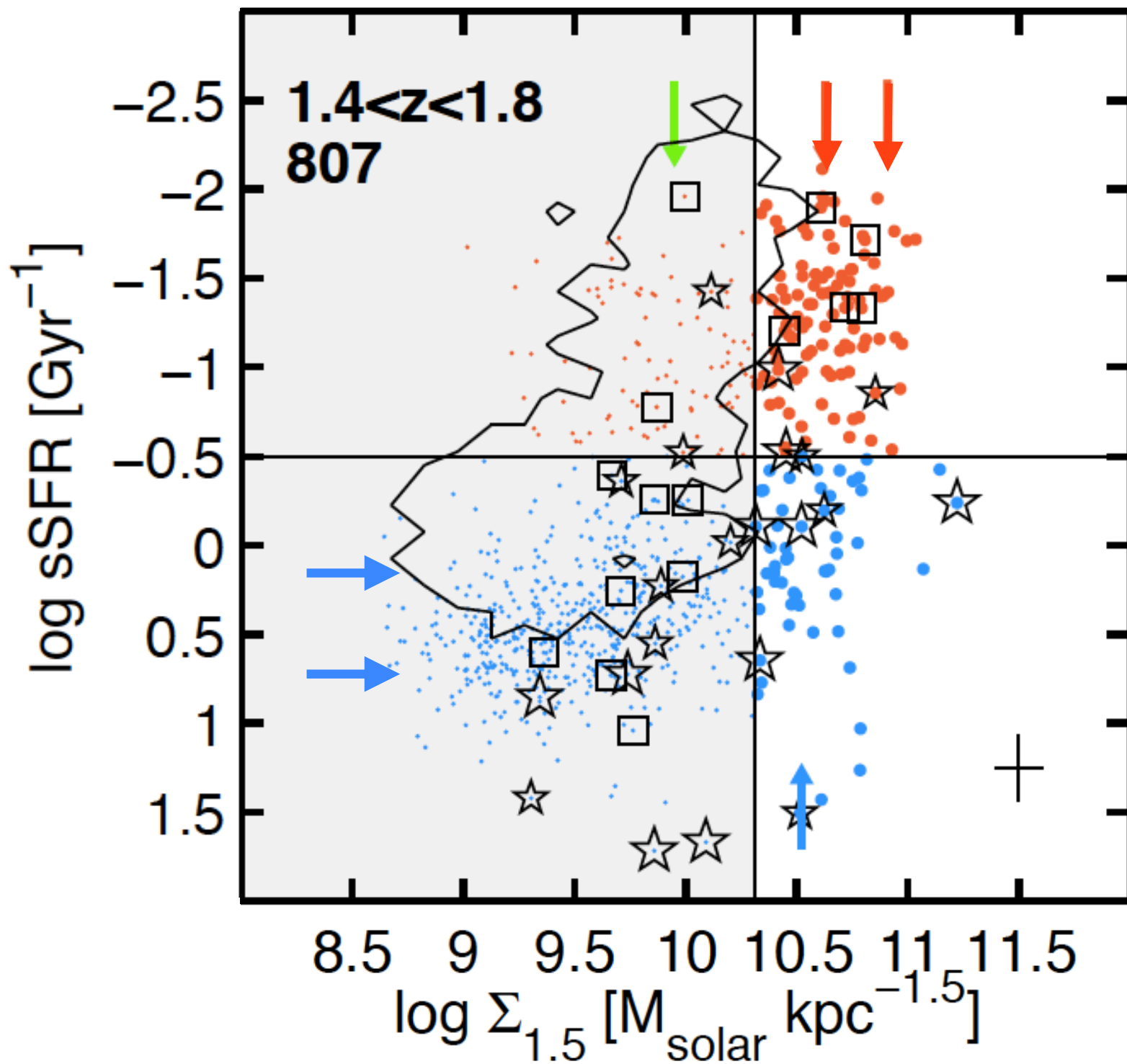


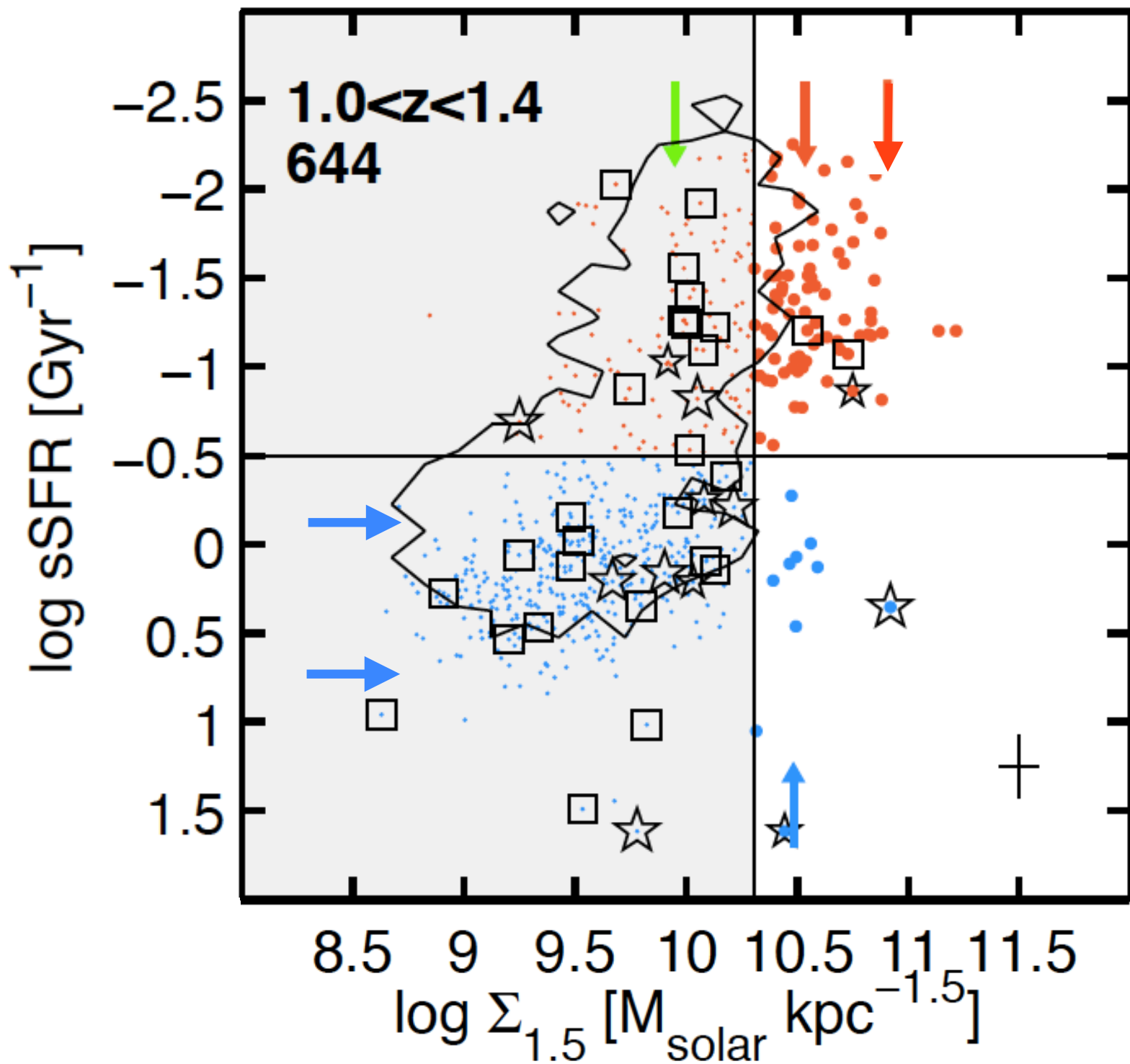


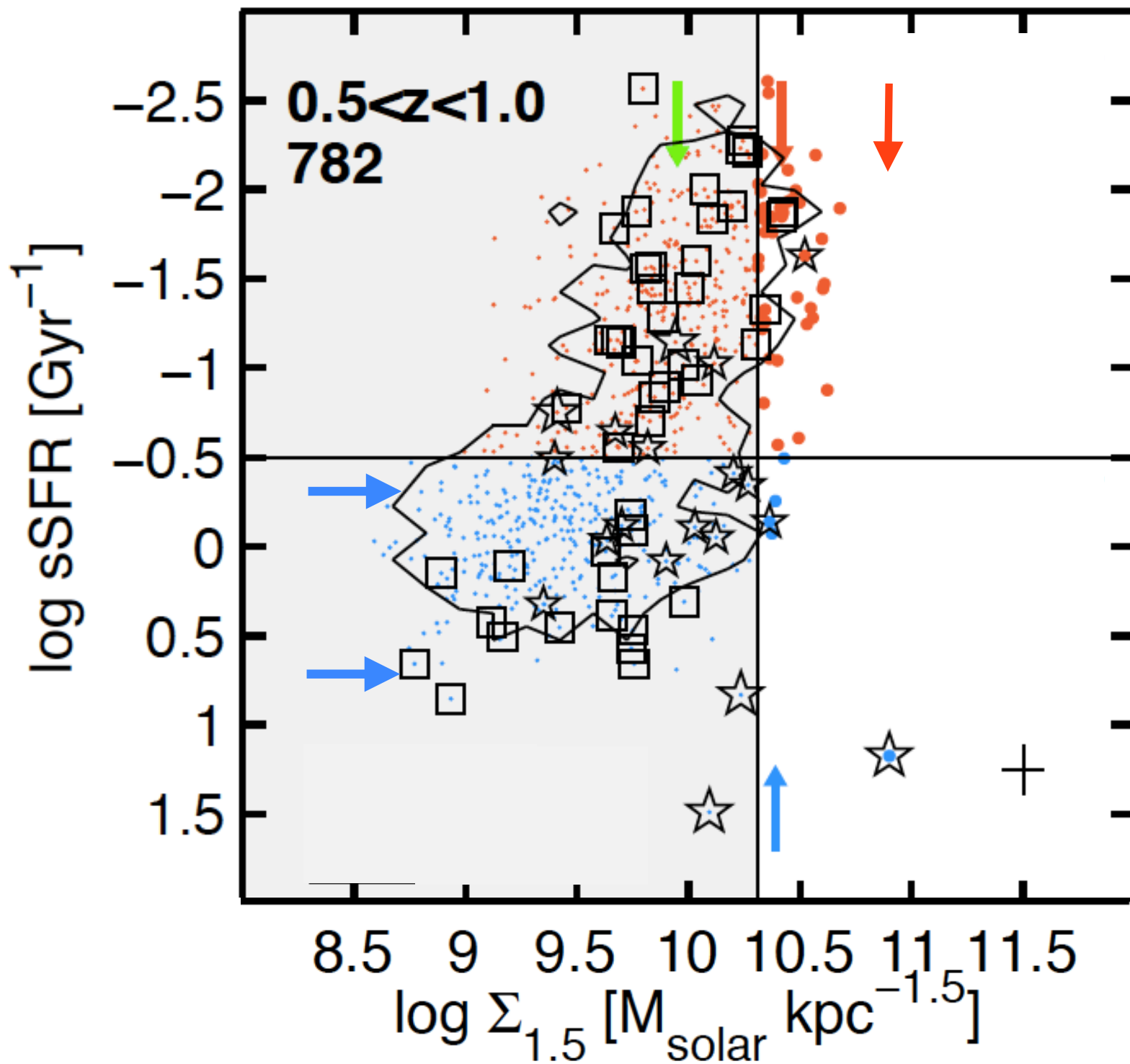






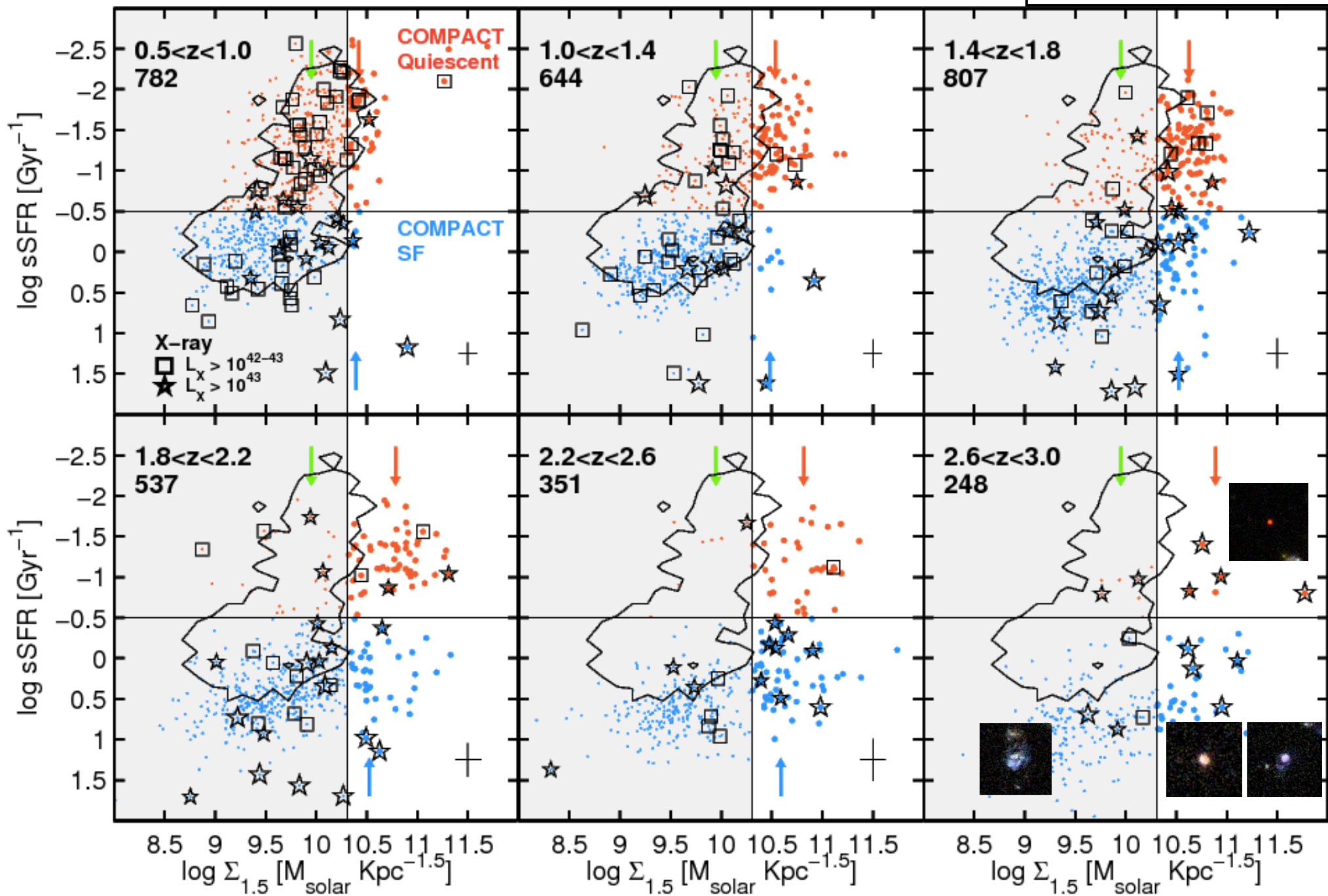




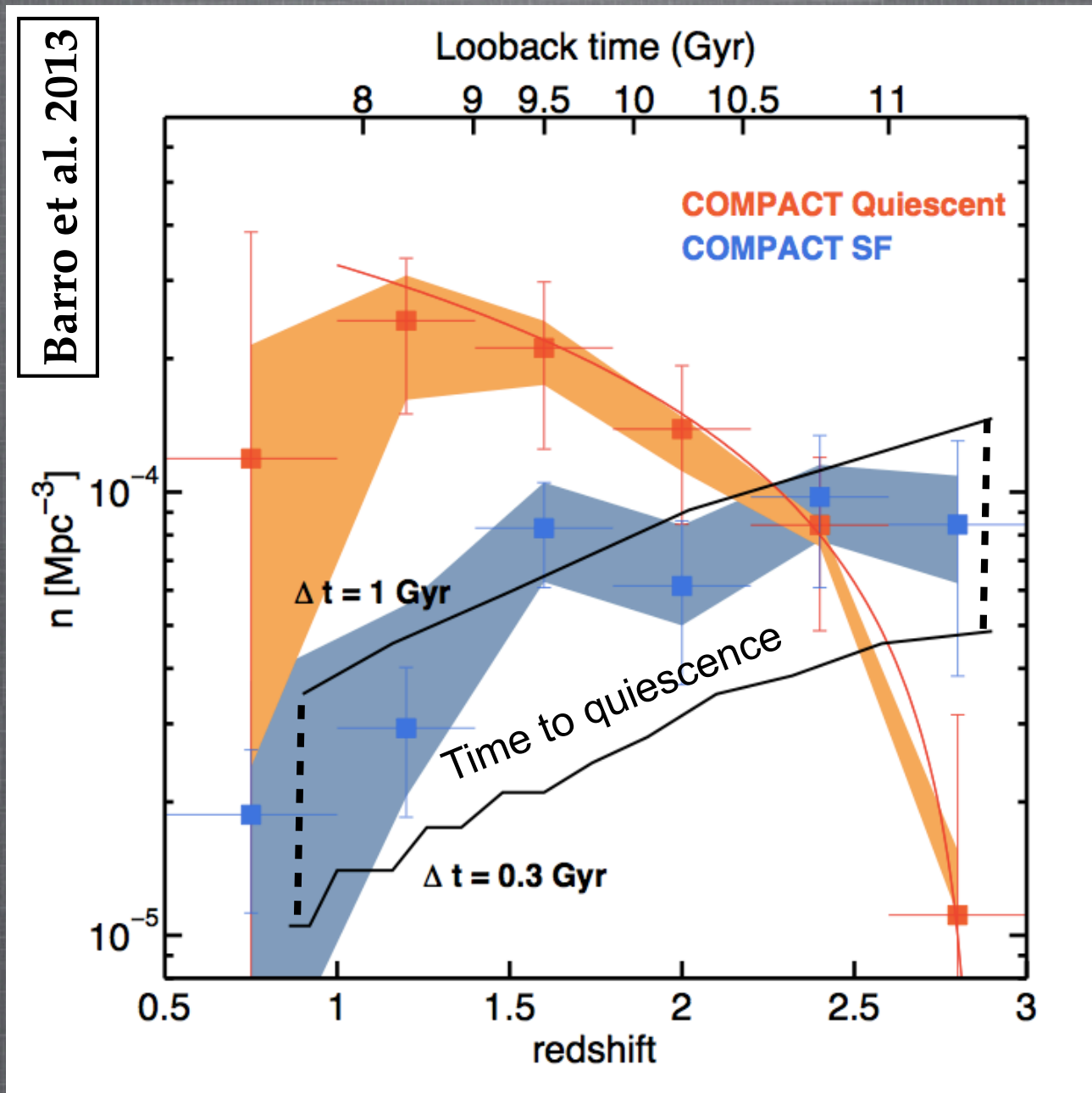


Compact quiescent and SFGs

Barro et al. 2013

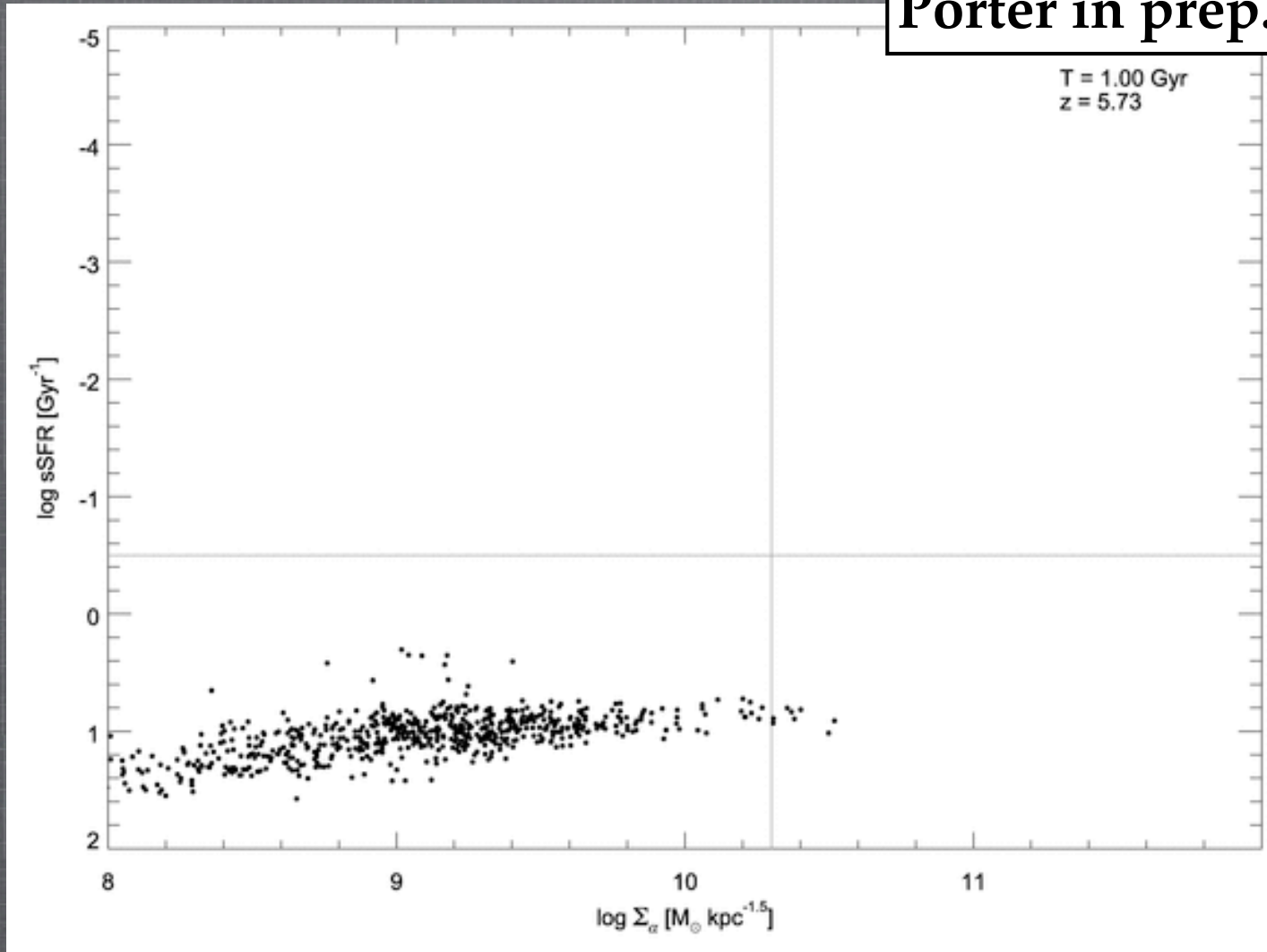


Density of compact galaxies



Life-paths of α SFGs from SAMs

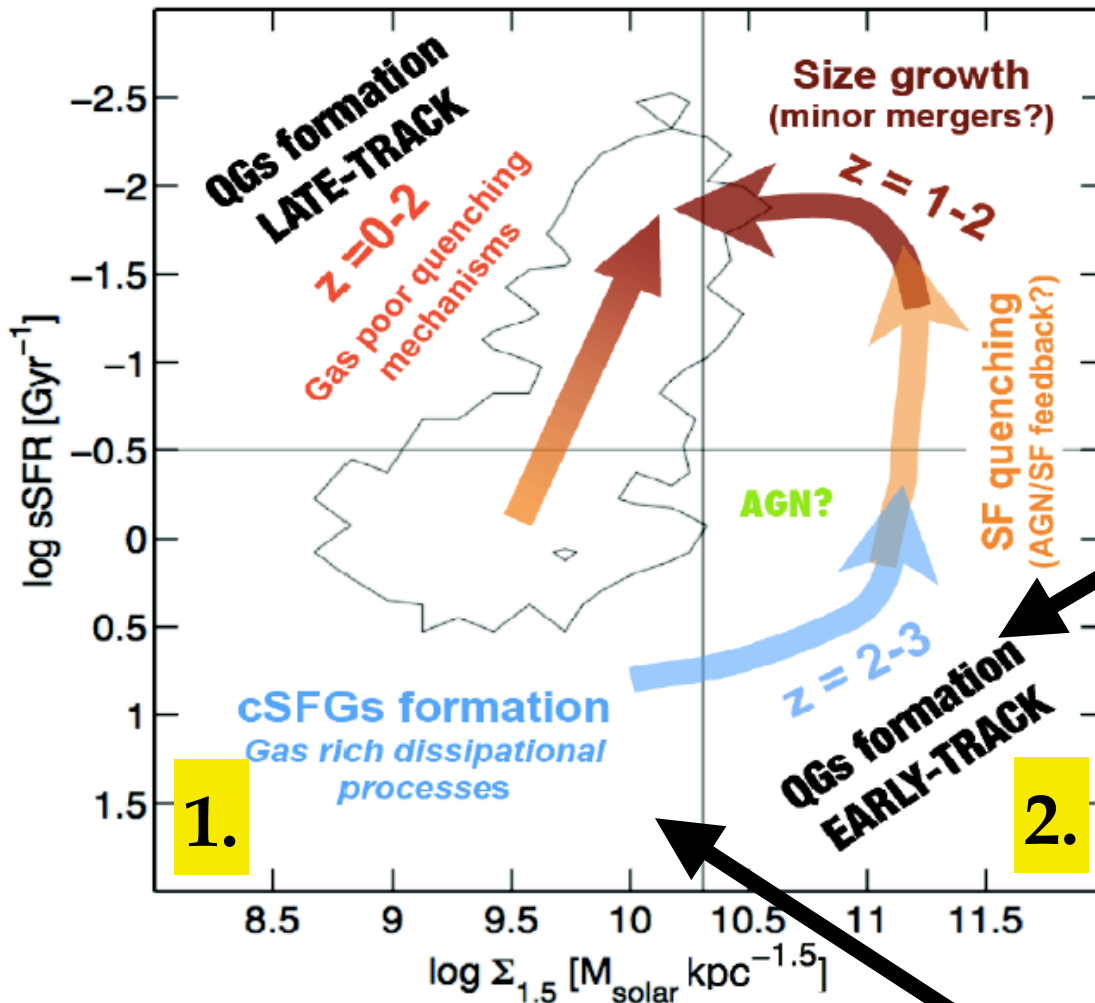
Porter in prep.



Semi-analytic model (Somerville+2010)
Bolshoi DM simulation (Klyping+2011)
Halo merger tree (Behroozi +2011)

Gas-rich merger in past Gyr
Gas-poor merger in past Gyr

Conclusions



Compact SFGs properties

- ❖ 80% dusty (IR-) star-formation. **2.**
- ❖ high-sersic, undisturbed app.
- ❖ 40% AGN det. fraction.
- ❖ 300 Myr - 1 Gyr quenching times.
- ❖ AGN/SF feedback (outflows?)

Compact SFGs formation

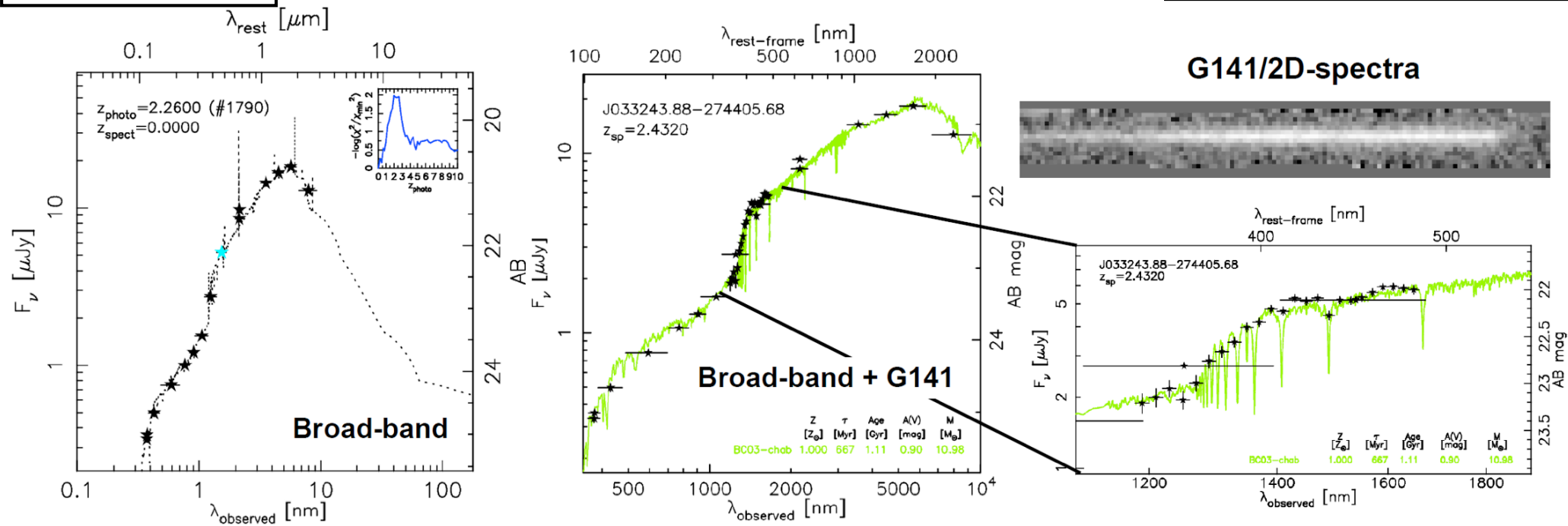
- ❖ SAMs - DI (60%) % wet mergers **1.**
- ❖ SAMs - Preferentially in already compact gal.
- ❖ ART-hydro - VDI time-scale 300 - 500 Myrs.

Improving the age estimates for cSFGs

- Will they quench in 300 Myr - 1 Gyr ?
- SED-based stellar properties for $2 < z < 3$ cSFGs

Barro in prep.

Higher resolution SEDs

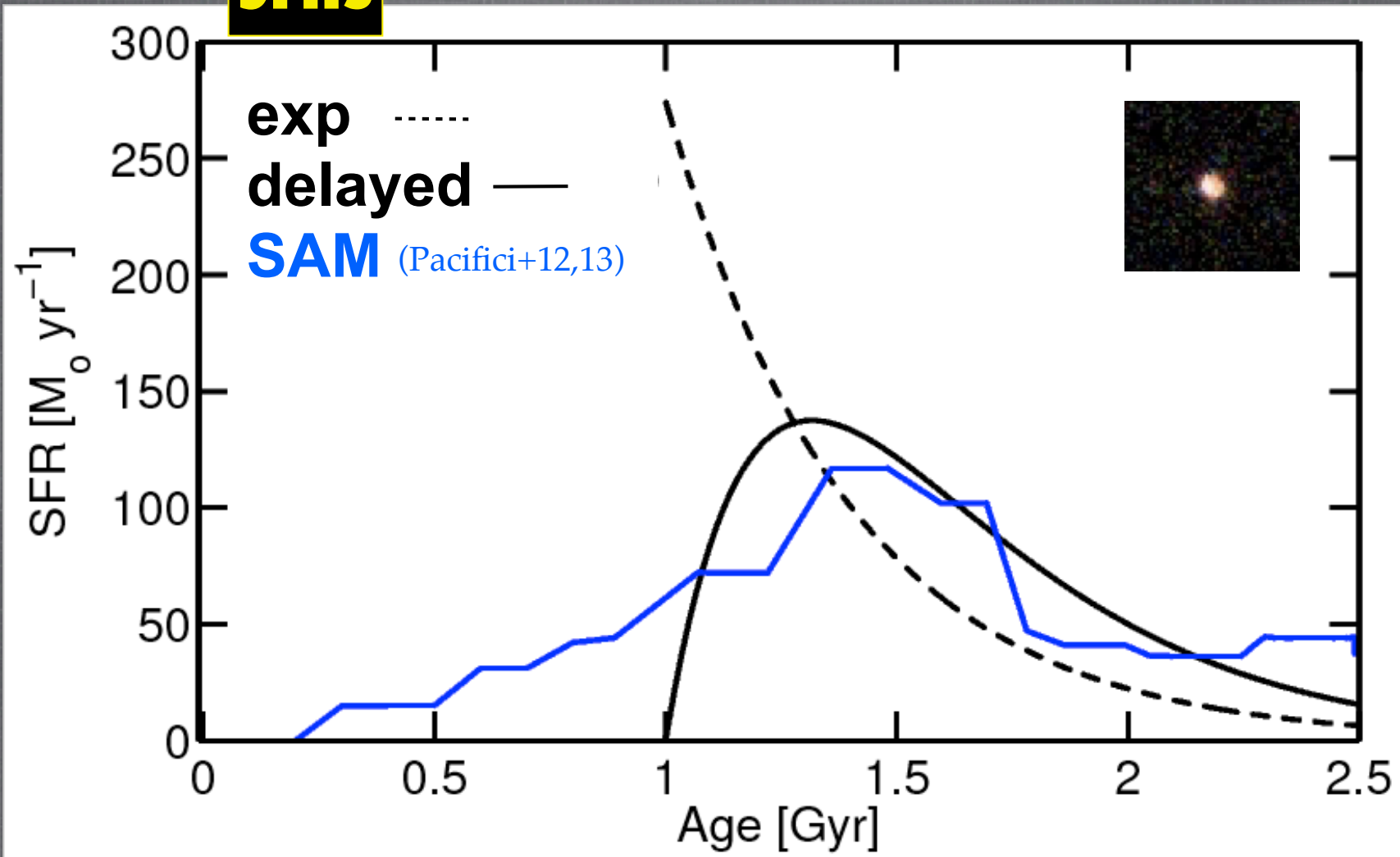


3D-HST (Brammer+12), NIR (1.1-1.7 microns) grism spectroscopy

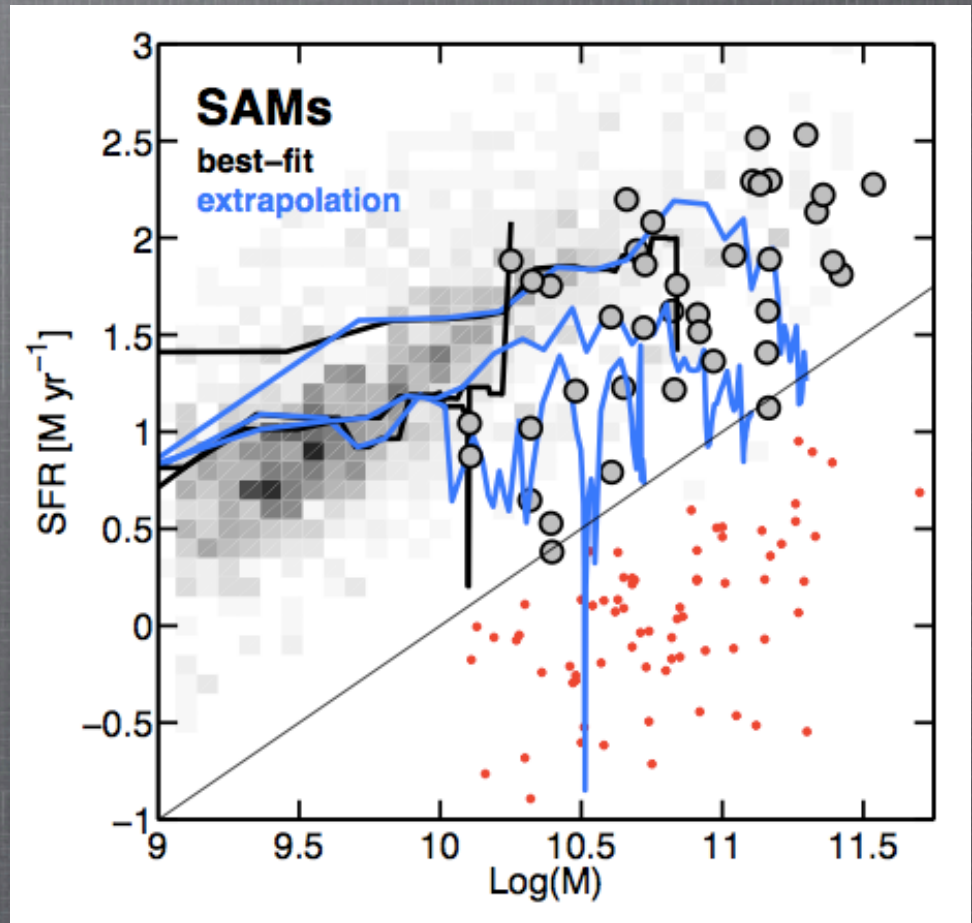
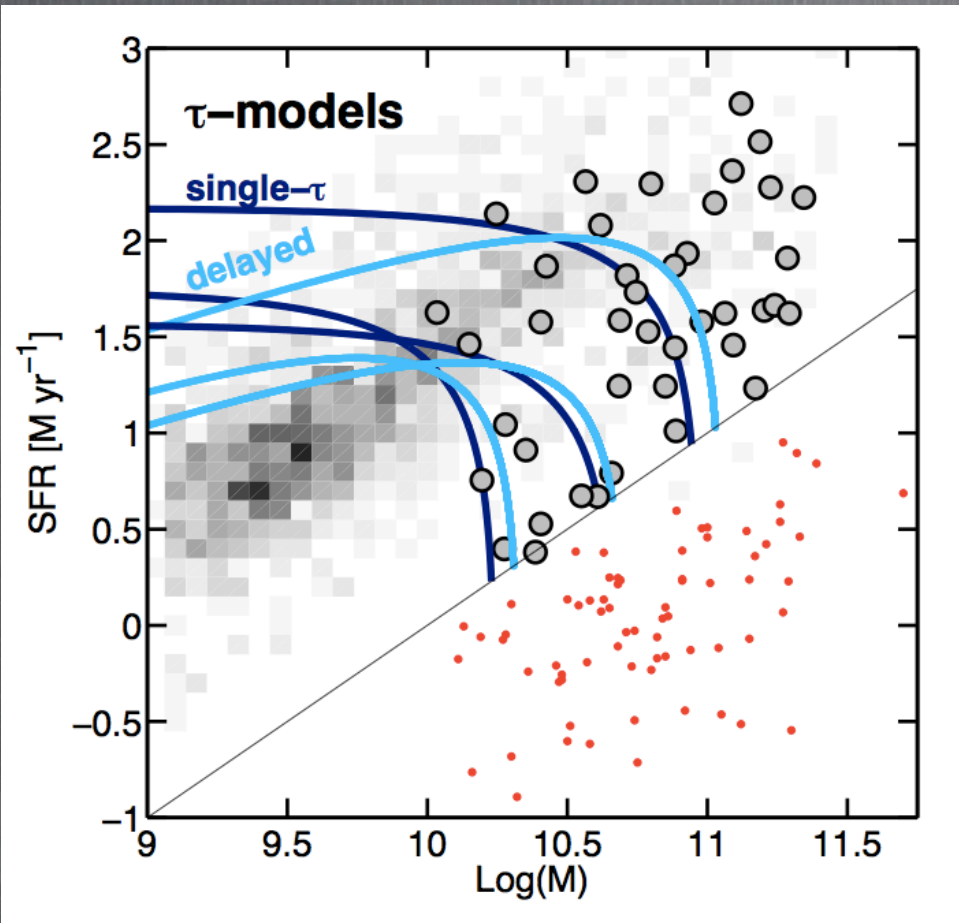
SED modeling

- Different SPS models, IMFs, metallicities and ..

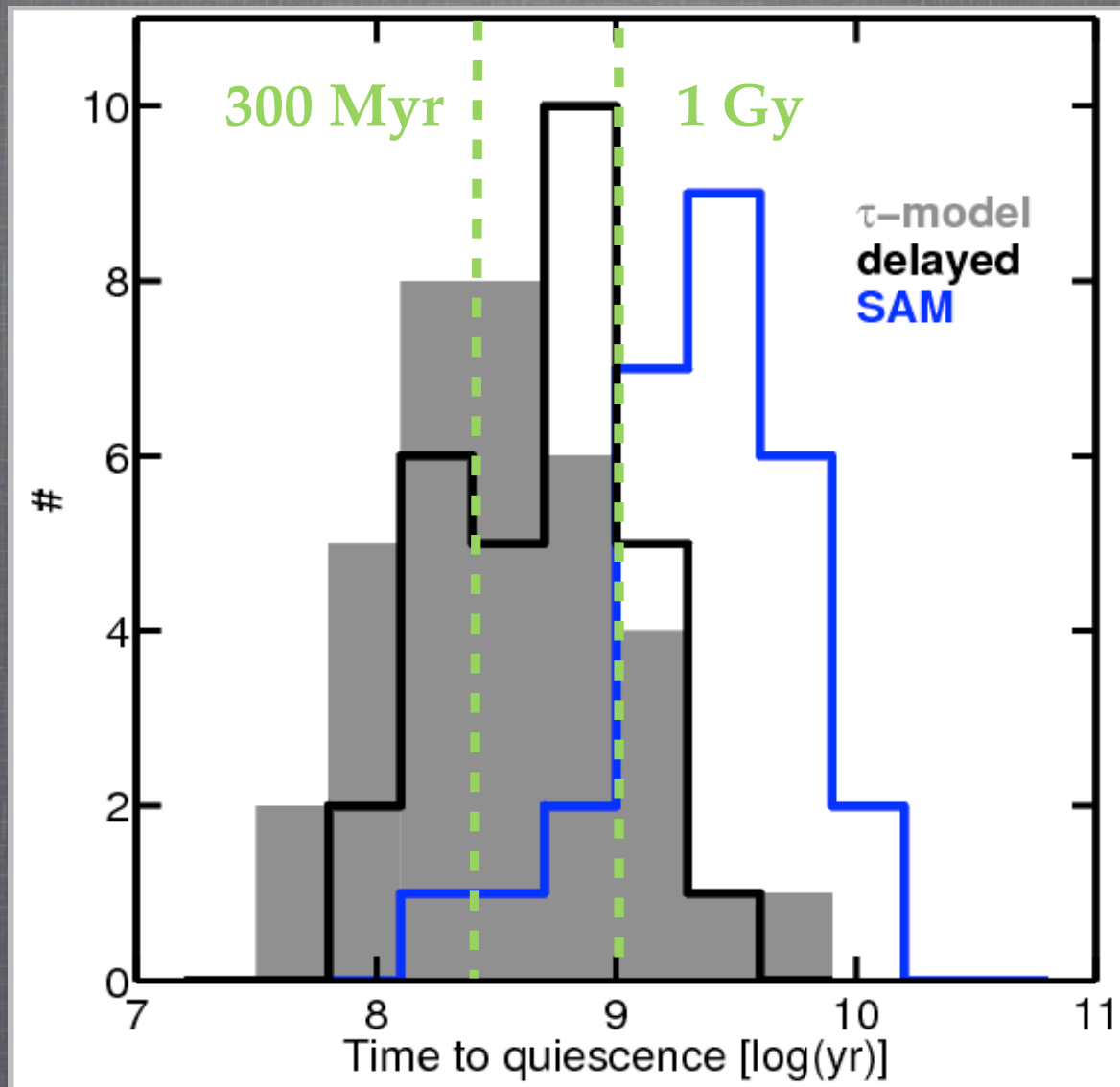
SFHs



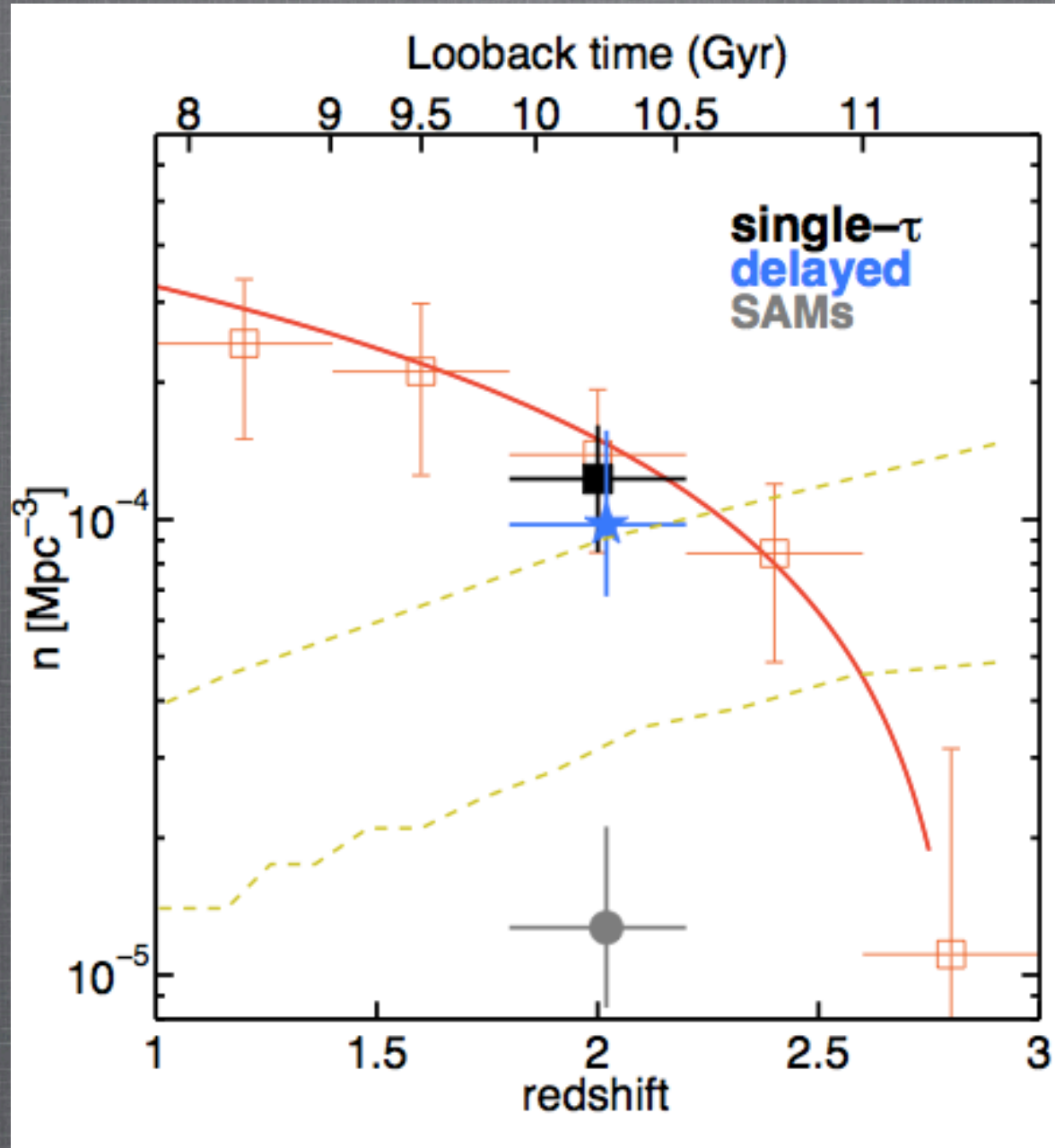
Life paths on the main-sequence



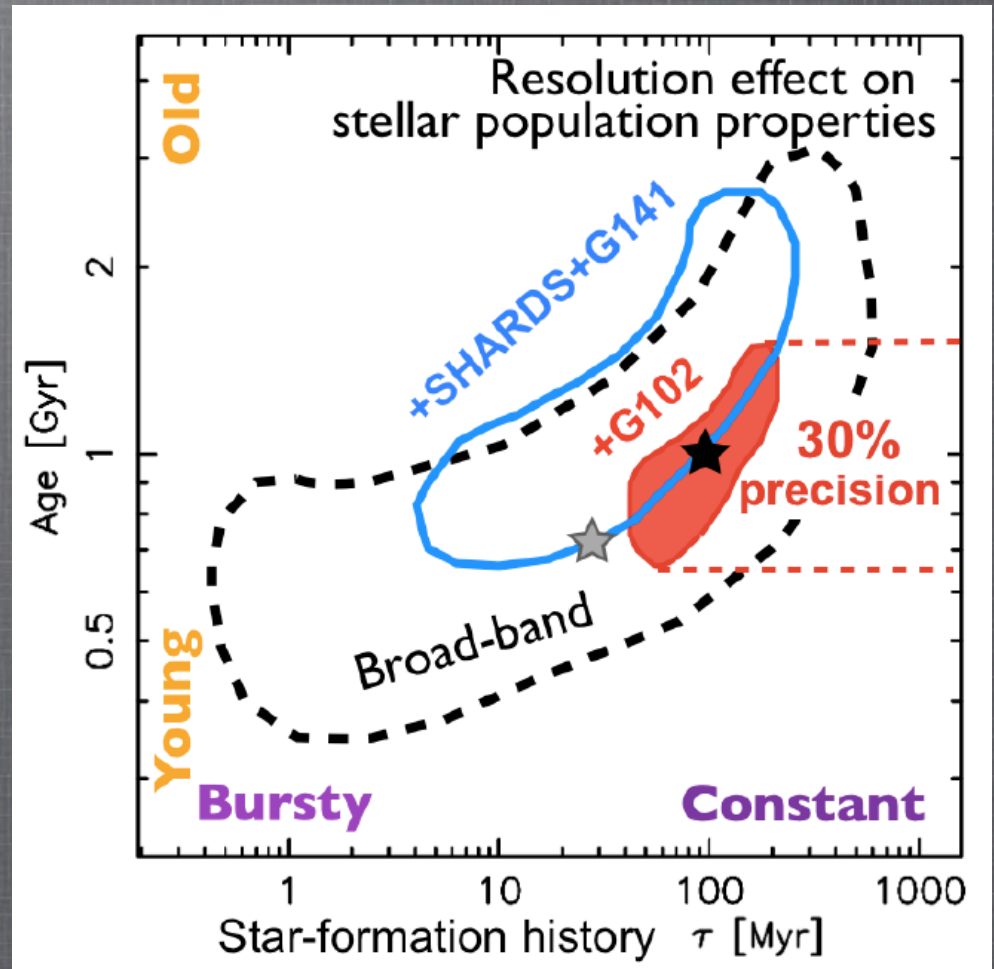
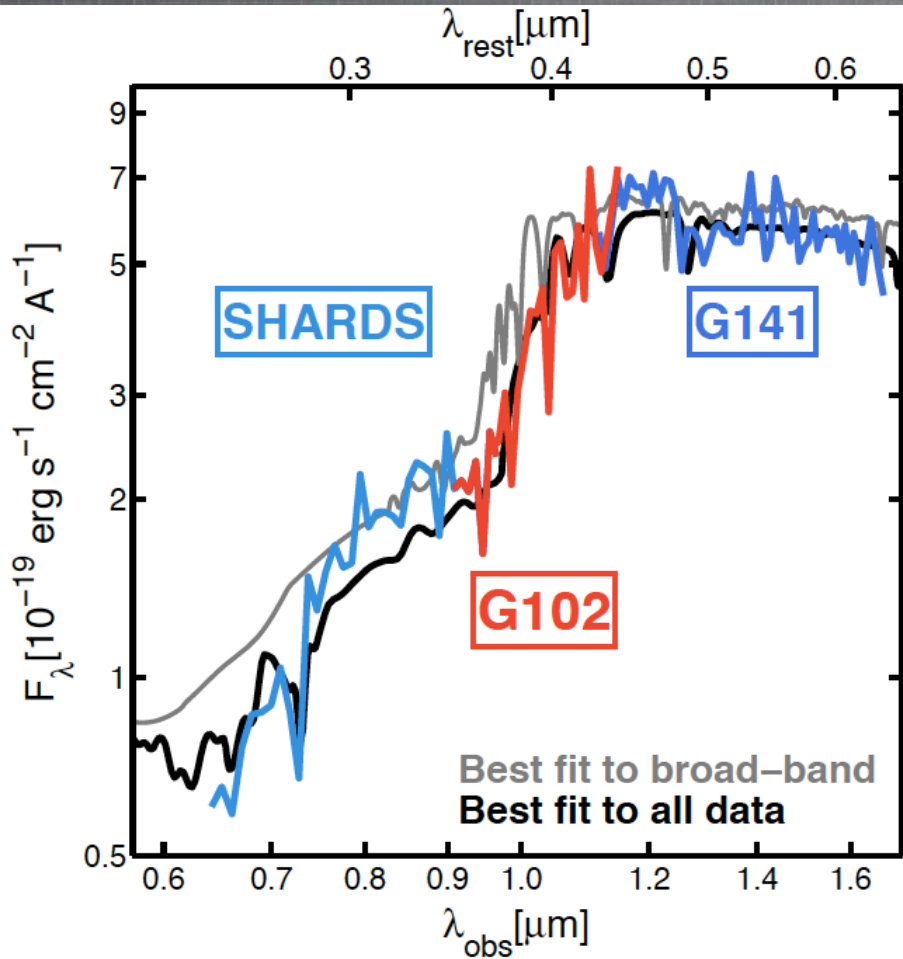
Elapsed times to quiescence



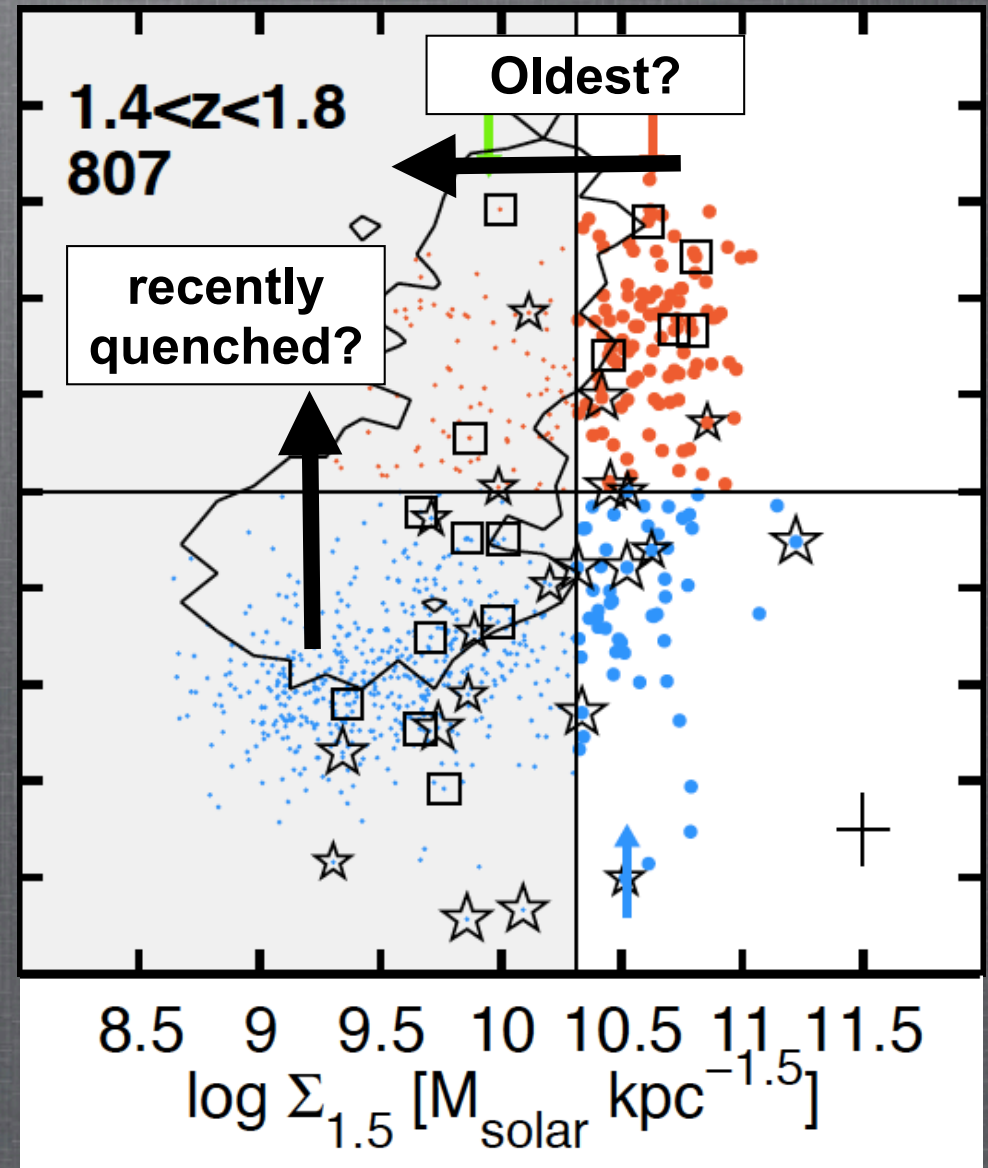
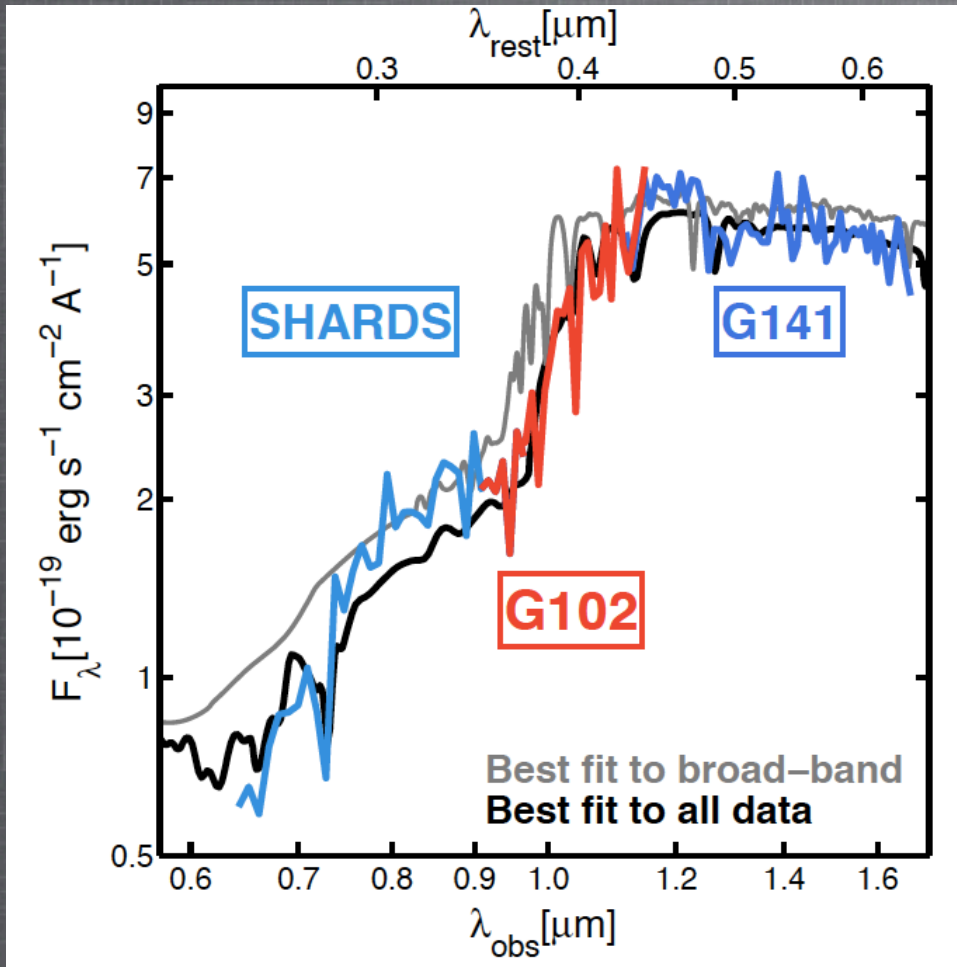
Predicted number densities



SHARDS + HST/GRISM at $1 < z < 2$

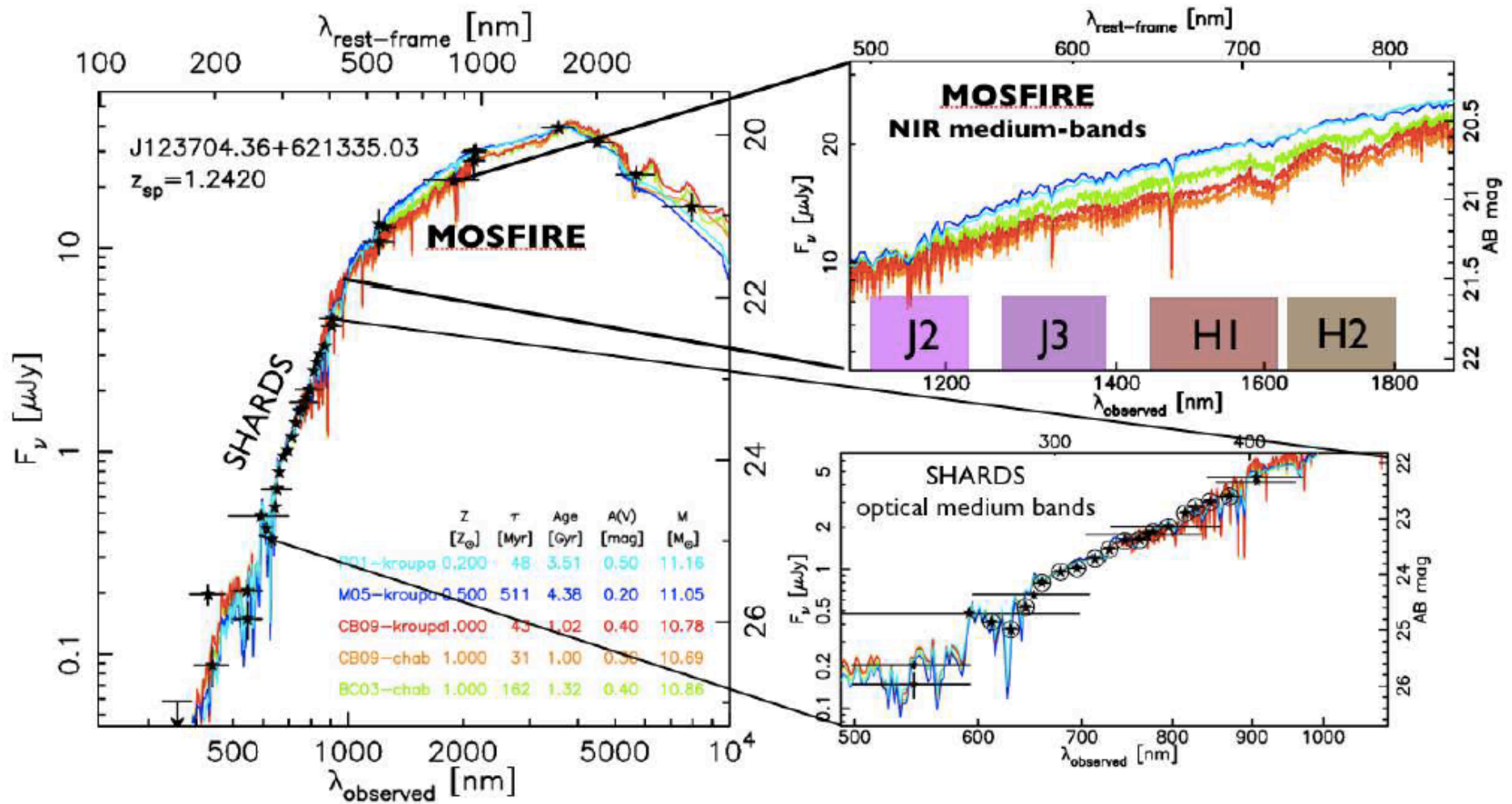


SHARDS + HST/GRISM at $1 < z < 2$

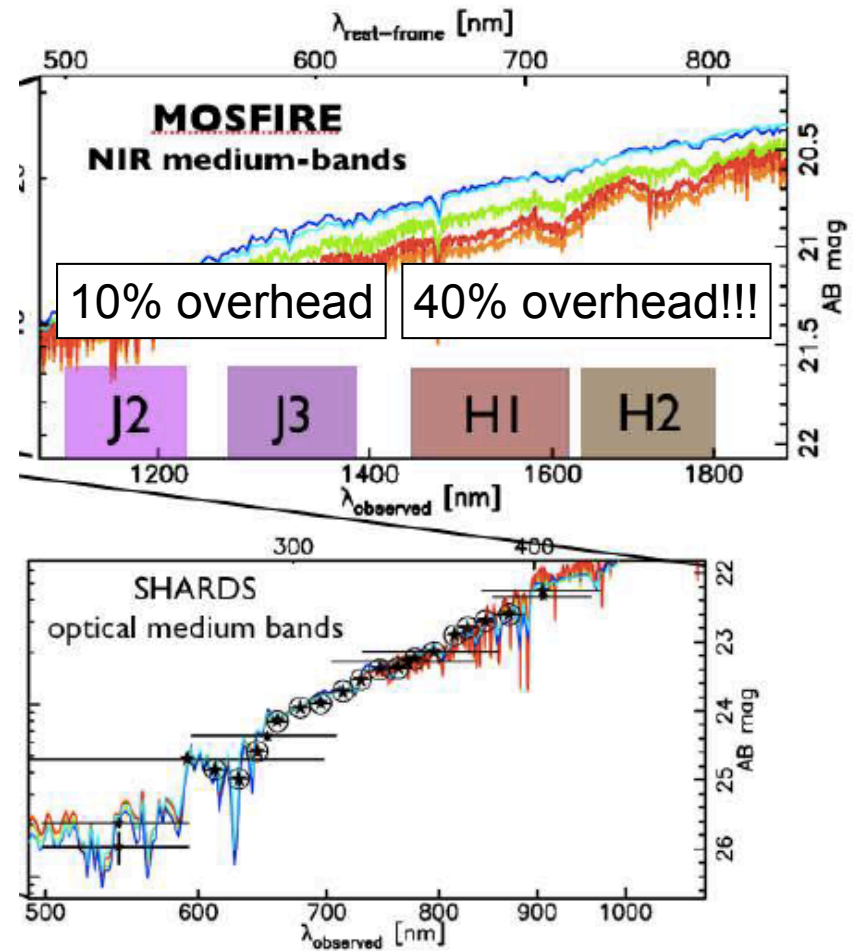
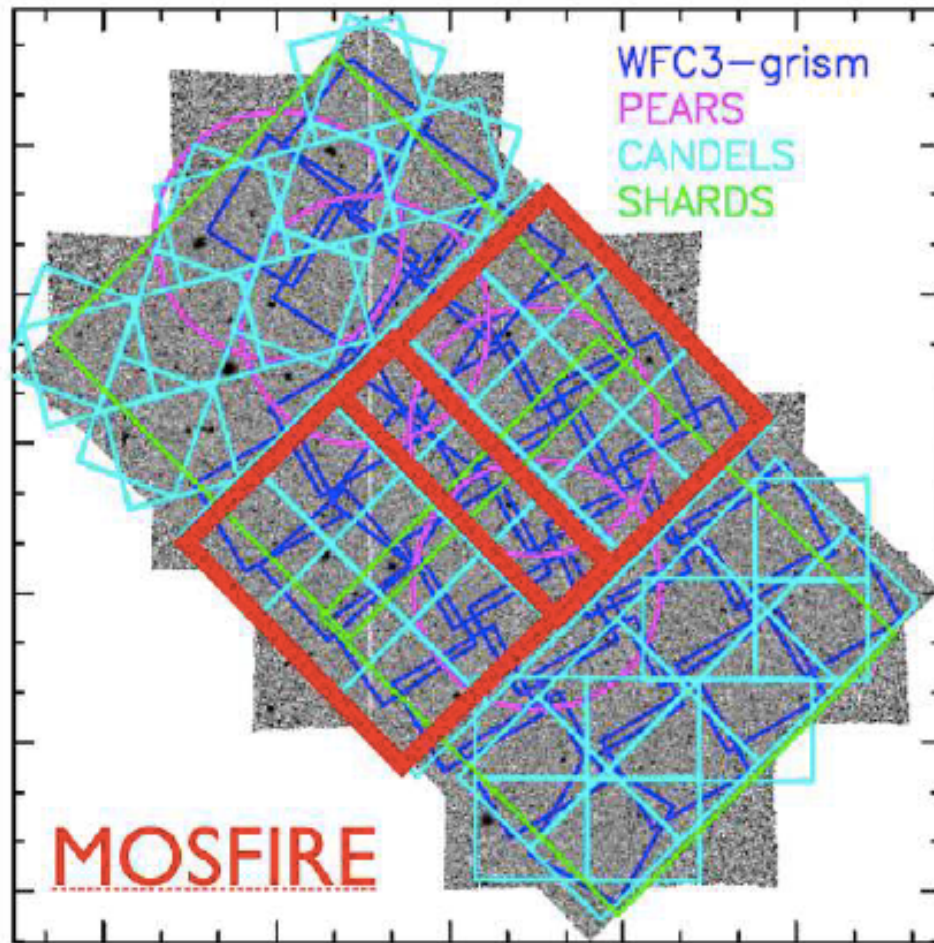


- ❖ Precision ages
- ❖ Correlation with structural properties/visual appearances

SHARDS + Keck/MOSFIRE (3 nights)

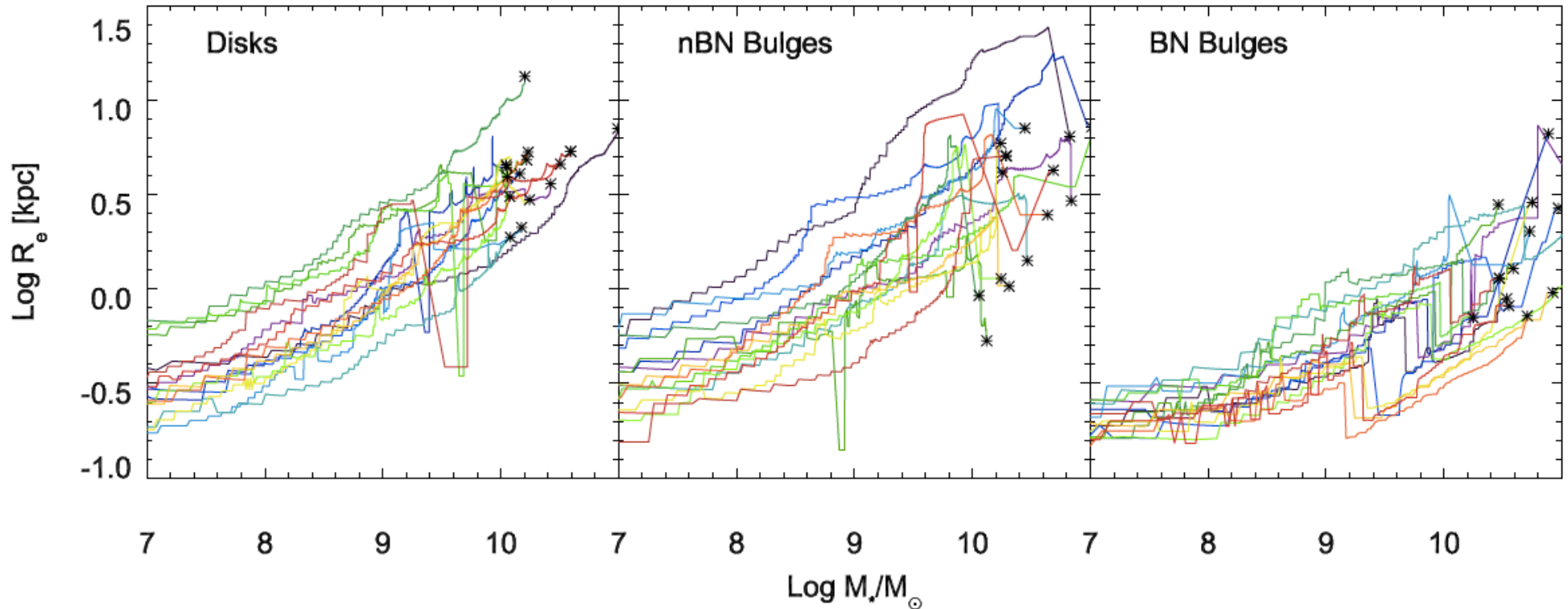


SHARDS + Keck/MOSFIRE (3 nights)



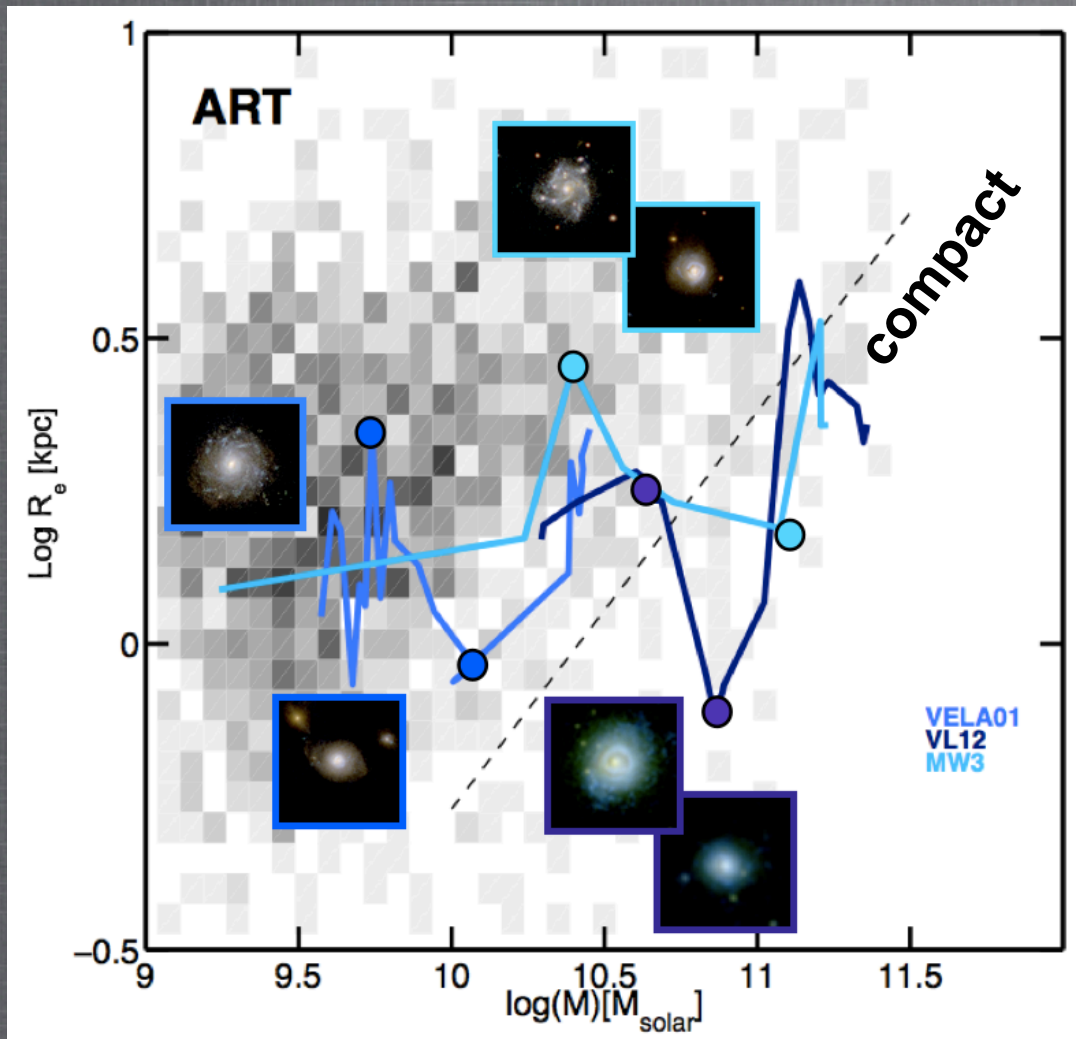
Life-paths of α SFGs from SAMs

Porter in prep.



- Sharp truncations are caused by disk instabilities more often (62%) than mergers

Life-paths of cSFGs from ART-Hydro



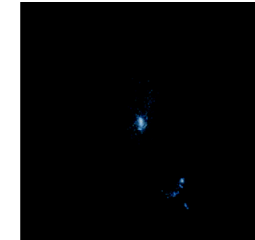
Moody in prep.

SUNRISE + PSF-match + noise

Kollipara in prep.

GALFIT

VL12



- ART-Hydro simulations (Ceverino+10,12, Dekel+13b)
- Violent disk instabilities, clump migration bulge growth

