

3193

3473

3867

3975

5149

5639

5879

6511

6610

6664

6669

6877

8466

UM417



Local XMP galaxies

Cool Flows

driven SF in *tadpoles*

*Casiana Muñoz-Tuñón*



- *Jorge Sanchez Almeida (IAC)*
- *Bruce Elmegreen (IBM- NY)*
- *Debra Elmegreen (Vassar Colege, NY)*

## Morphologies in the UHDF

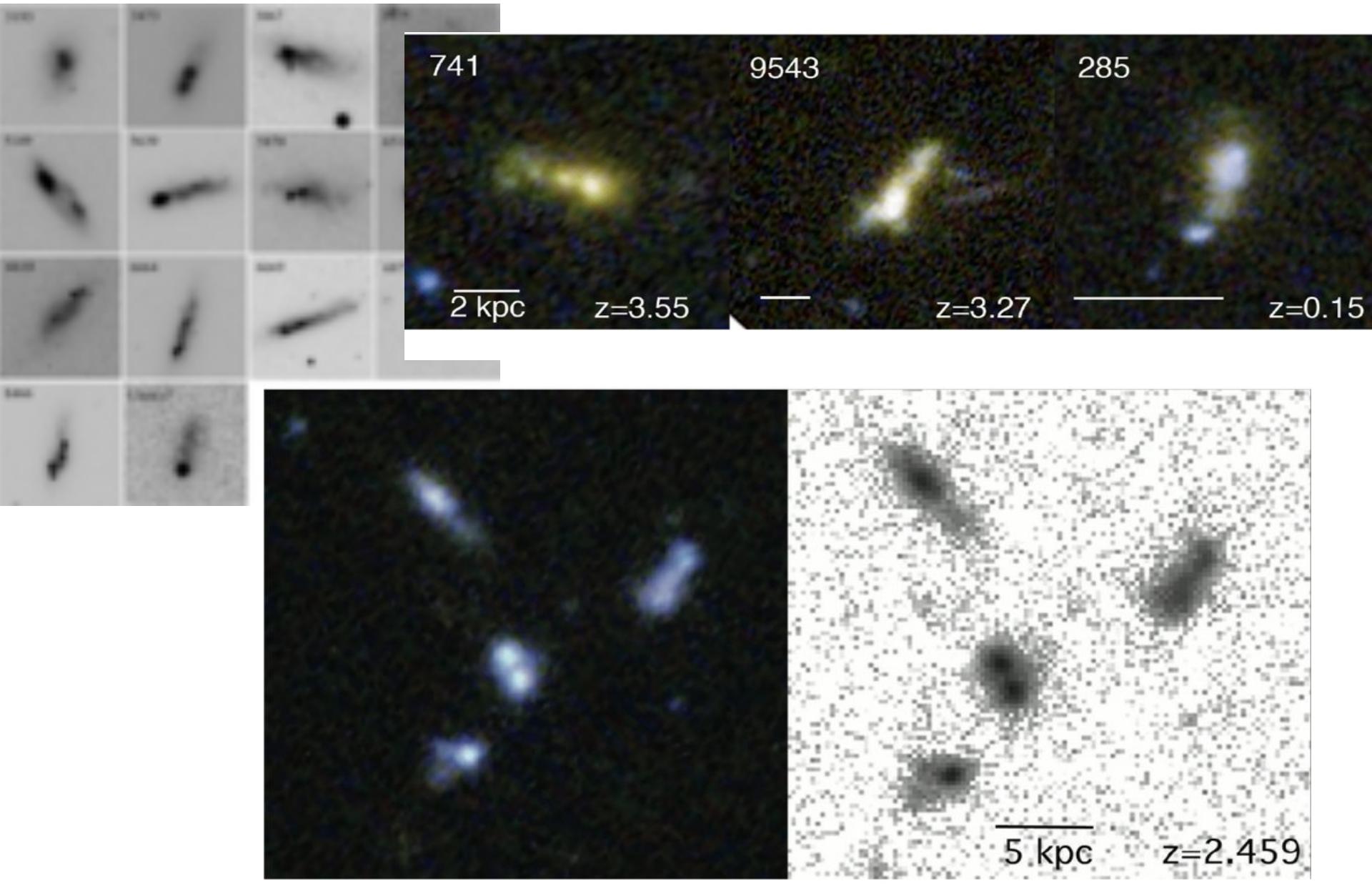
Tadpoles

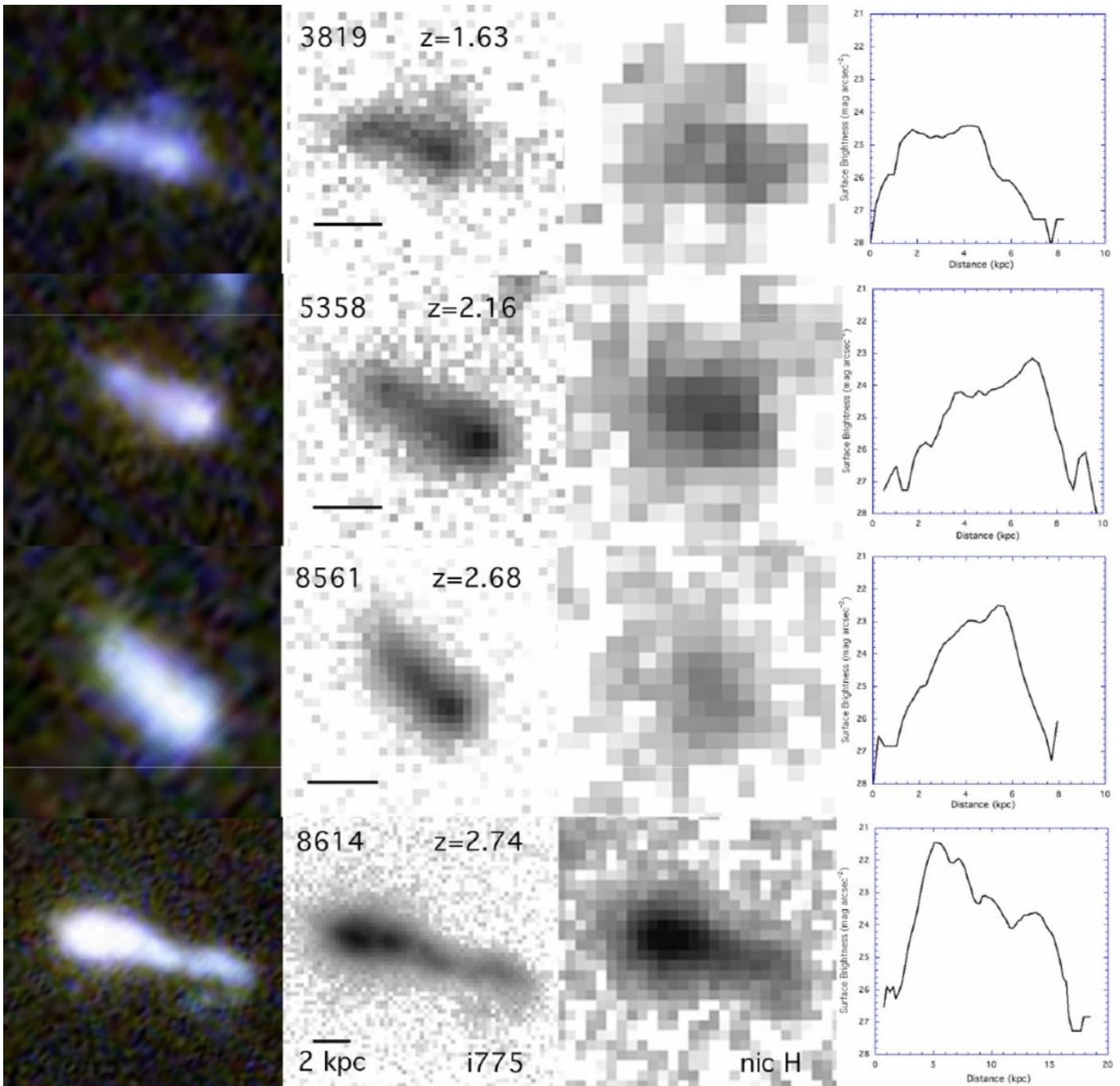


10%

Figure 12 from Tadpole Galaxies in the Hubble Ultra Deep Field

Bruce G. Elmegreen and Debra Meloy Elmegreen 2010 ApJ 722 1895 doi:10.1088/0004-637X/722/2/1895





6/20/13

Figure 1 from Tadpole Galaxies in the Hubble Ultra Deep Field 4  
Bruce G. Elmegreen and Debra Meloy Elmegreen 2010 ApJ 722 1895.

3193

3473

3867

3975

5 kpc

5149

5870

6511

6610

6664

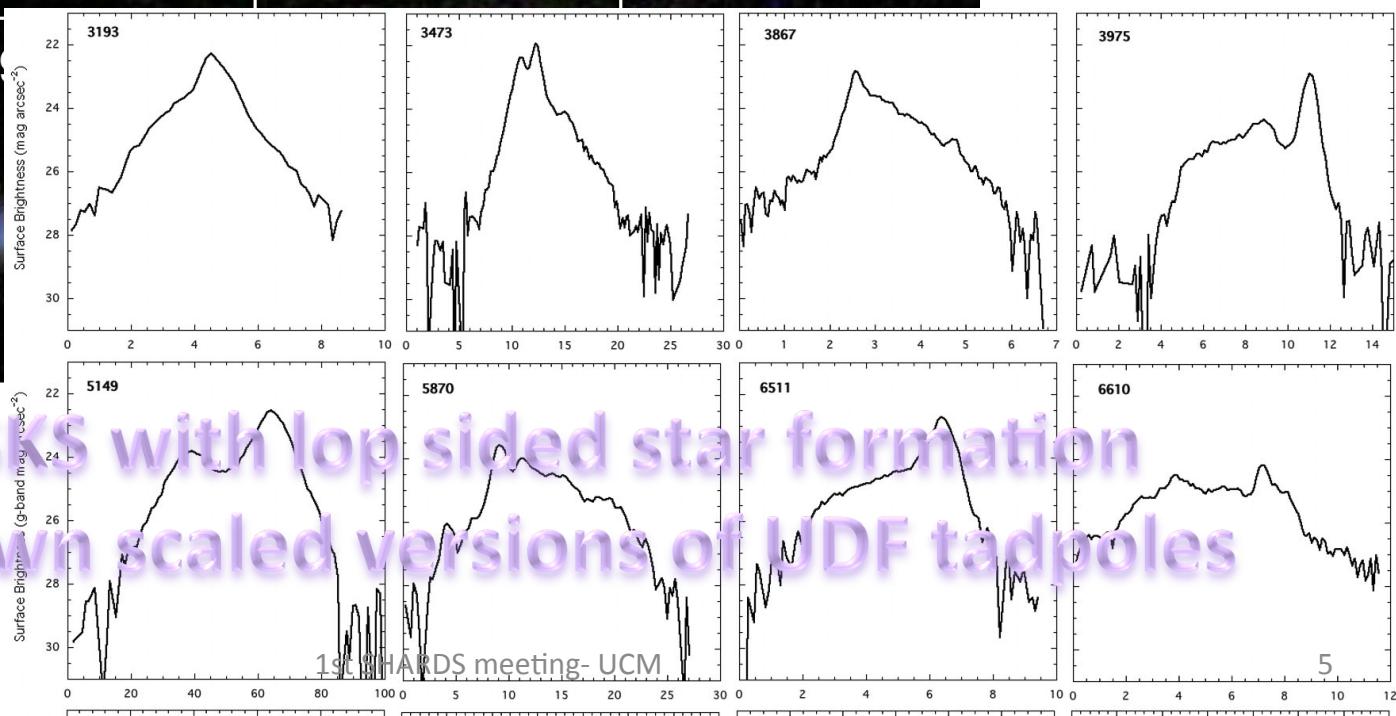
666

UM417

6/20/13

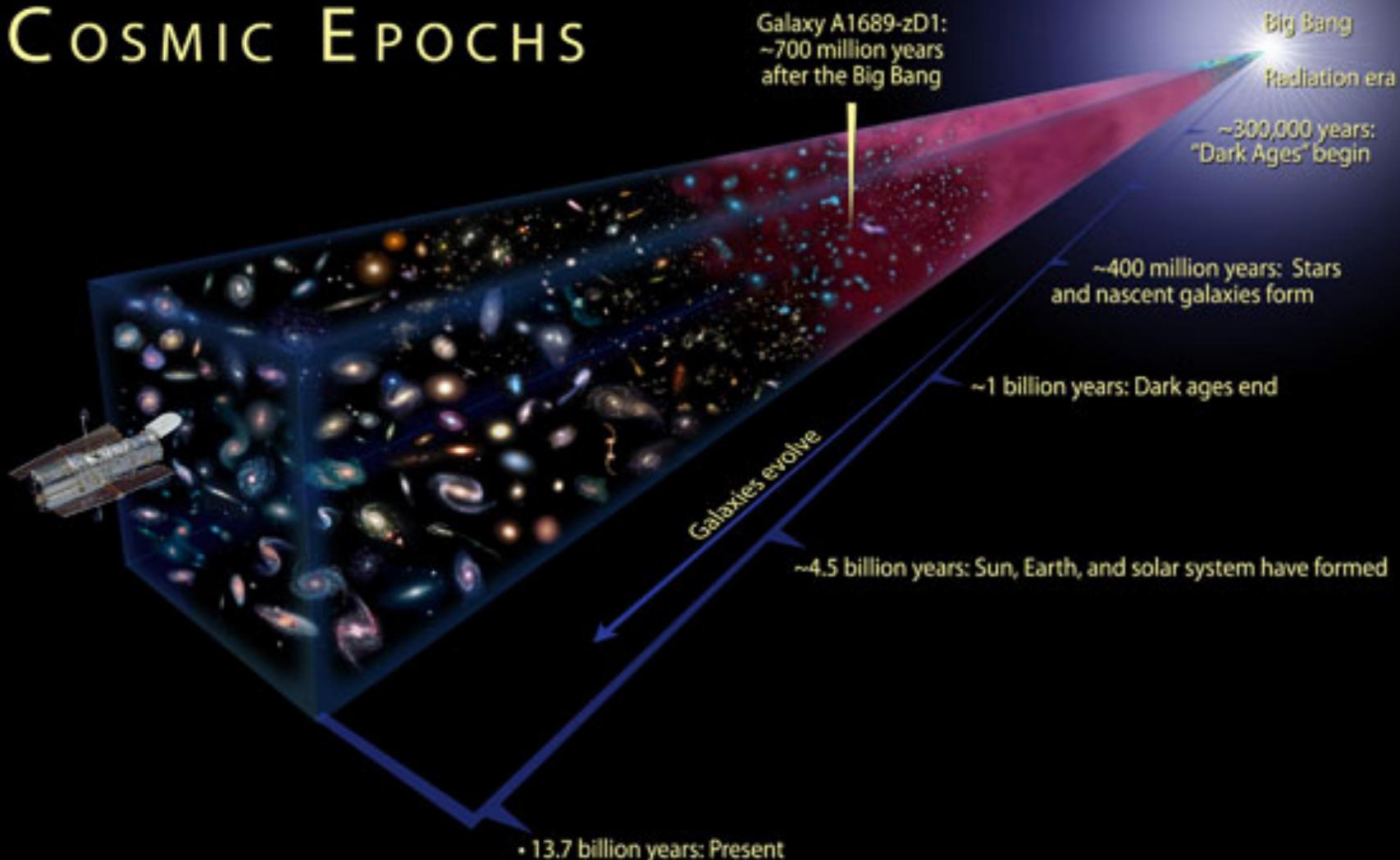
- UM and KISO archive of bright UV galaxies.
- Sloan and H-UDF

Elmegreen, D. et al.,  
2012, Ap.J., 750, 95.

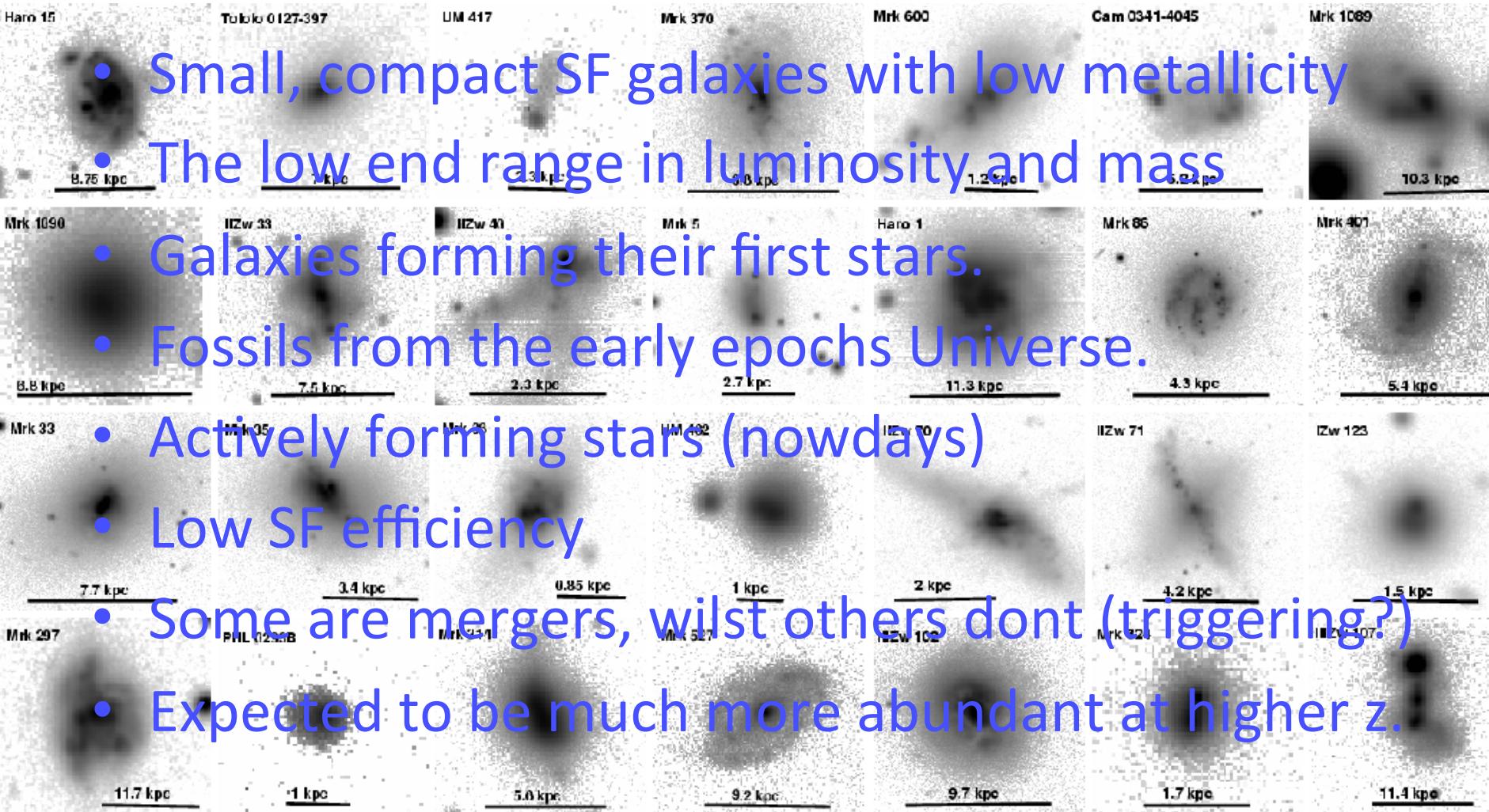


# Dwarf Galaxies

## COSMIC EPOCHS

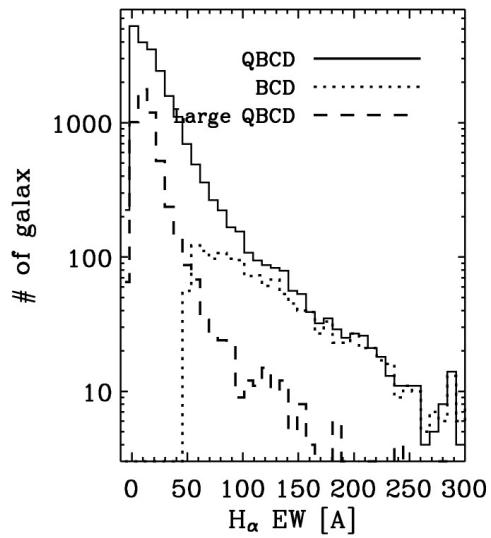
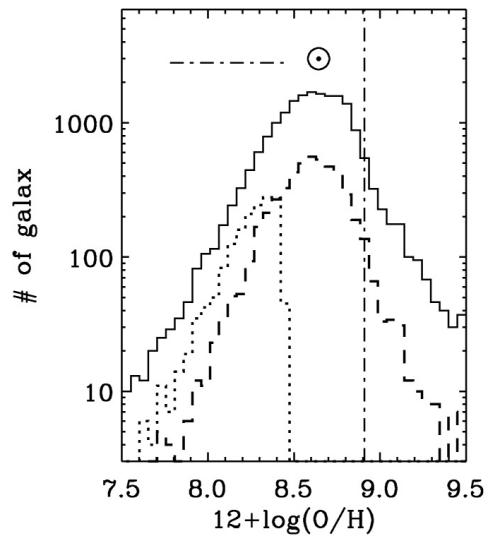
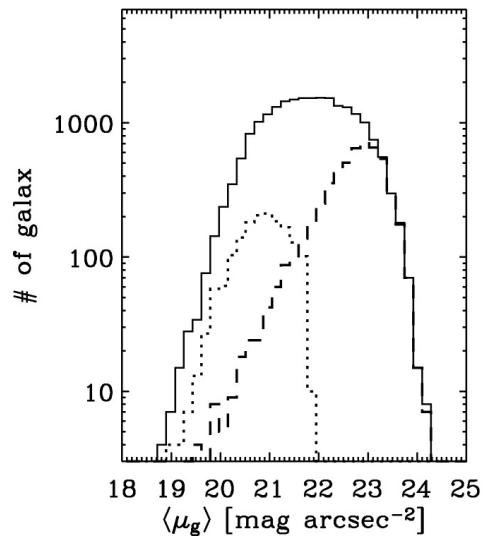
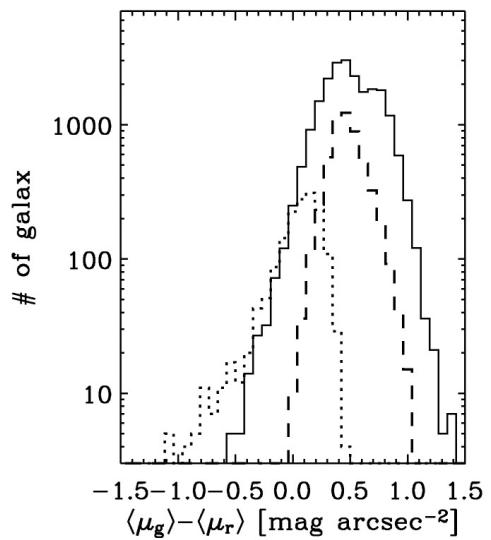
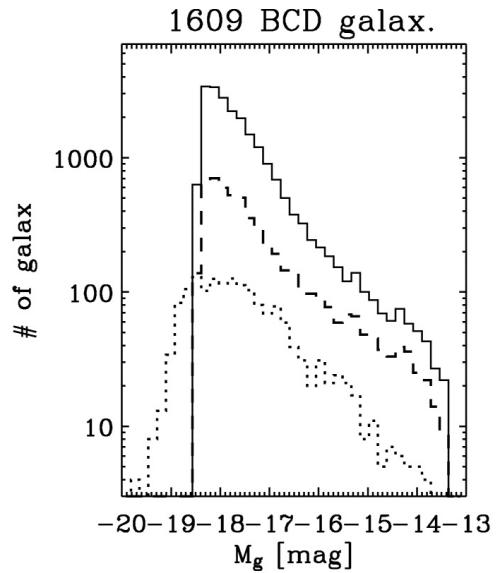
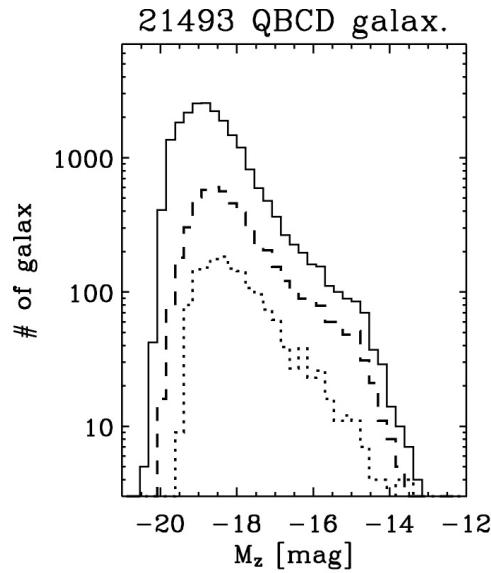


# Blue Compact Dwarfs (BCDs) (Cairós sample; Amorín)



# Search using the SDSS

- Blue Compact Dwarf Galaxies (BCDs) (DR6)
- Systematic Search for Quiescent BCDs (21.500 glxs)
- Automatic Classification of Sloan Spectra (DR7)
- Search for XMP targets.
- Complete census of the nearby BCDs, QBCDs and XMPs

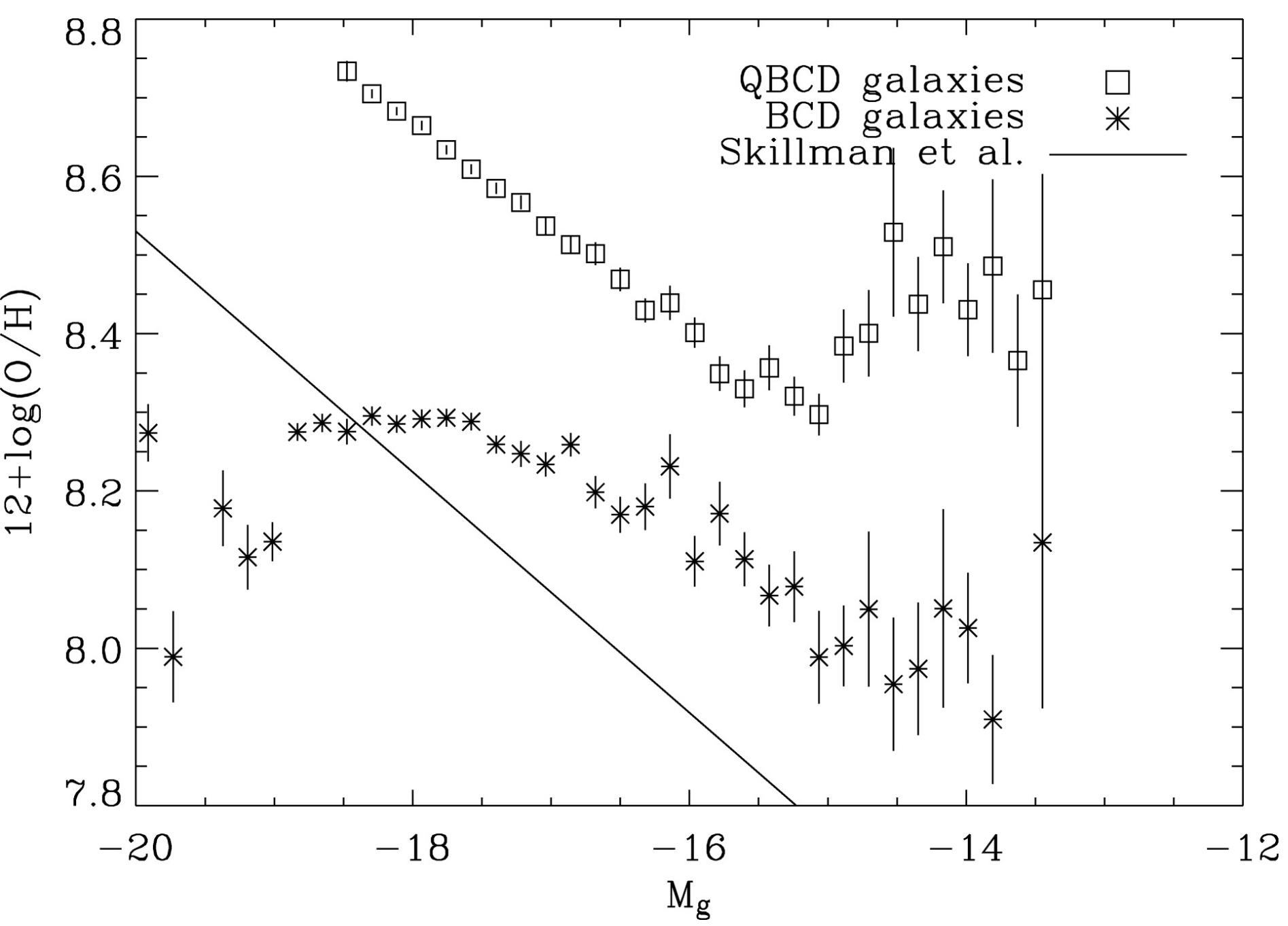


## BCDs and QBCDs in SSDS DR6

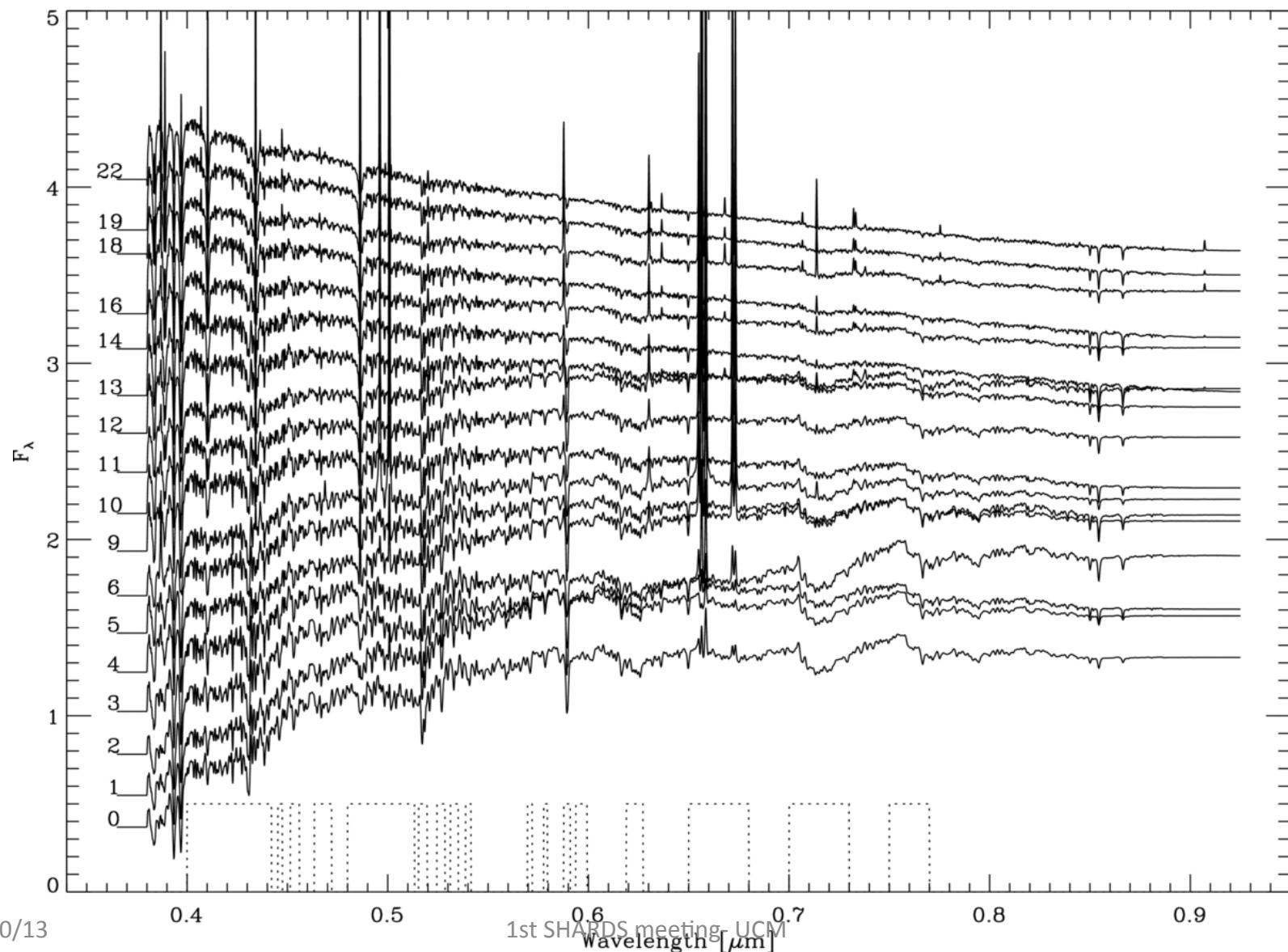
(SA, MT et al., 2008, ApJ., 685, 198-210)

6/20/13

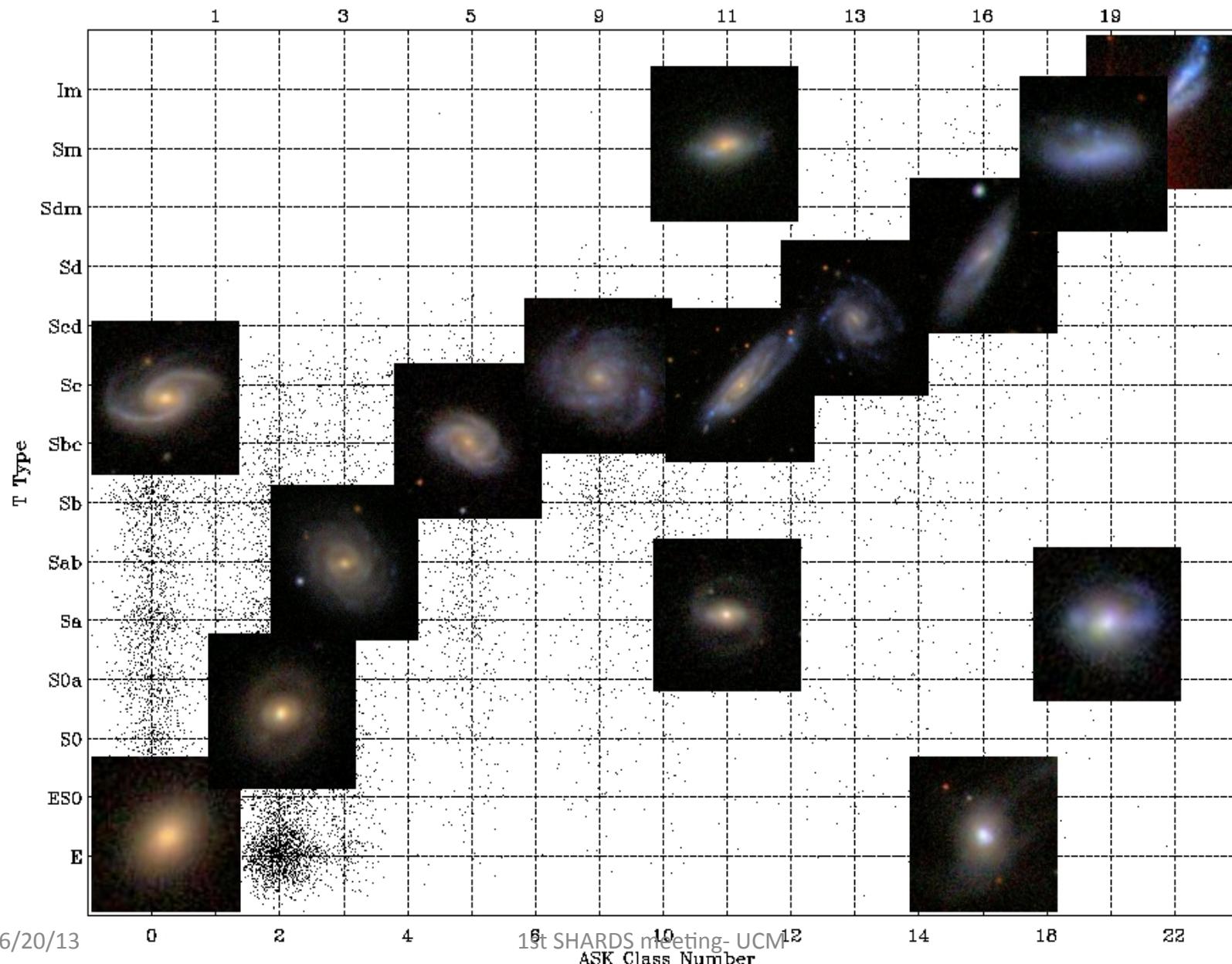
1st SHARDS meeting- UCM



Automatic Unsupervised Classification of All Sloan Digital Sky Survey DR 7 Galaxy Spectra  
SA et al. 2010 ApJ 714 487  
*Some 930.000 galaxies Using K-means cluster analysis algorithm*

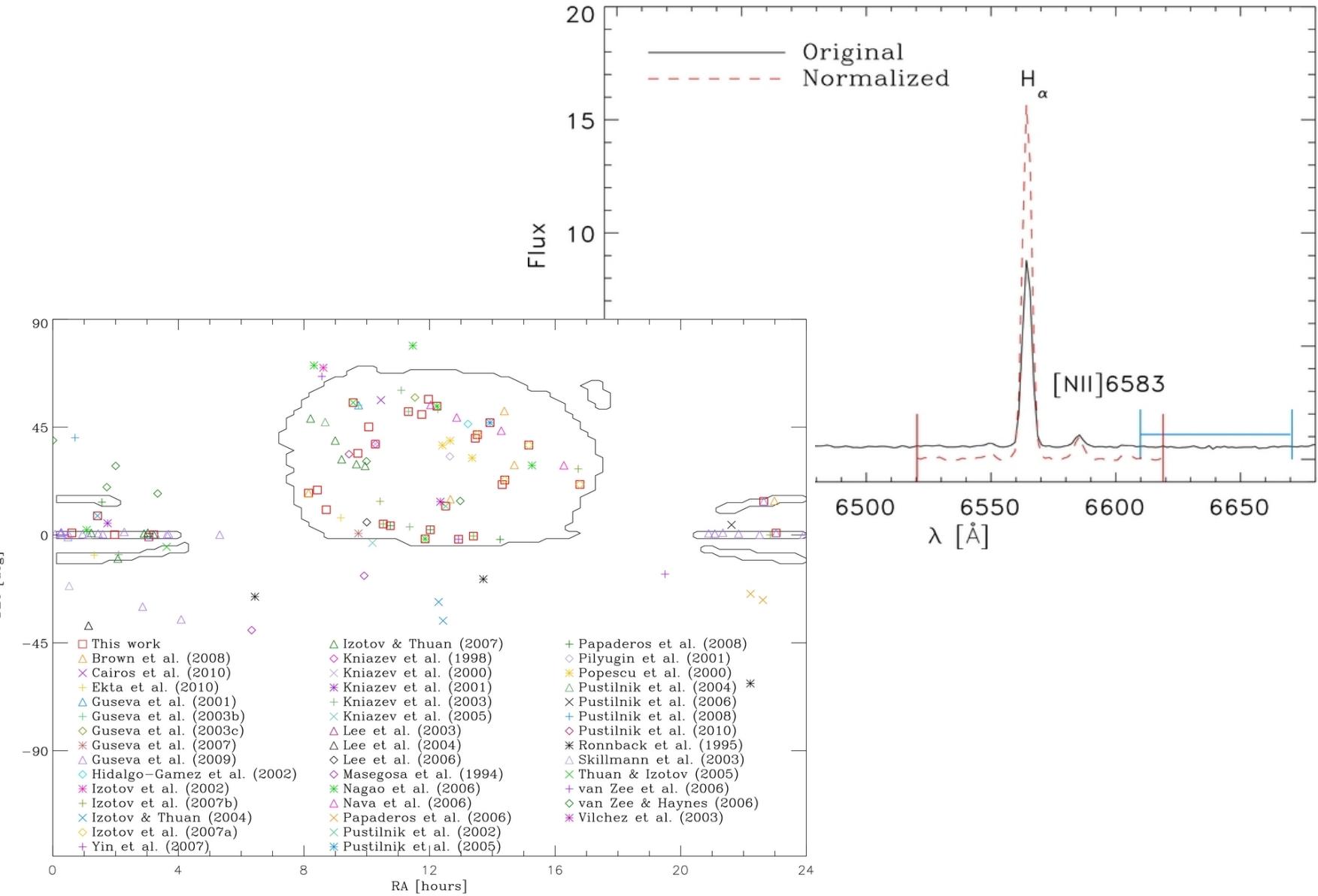


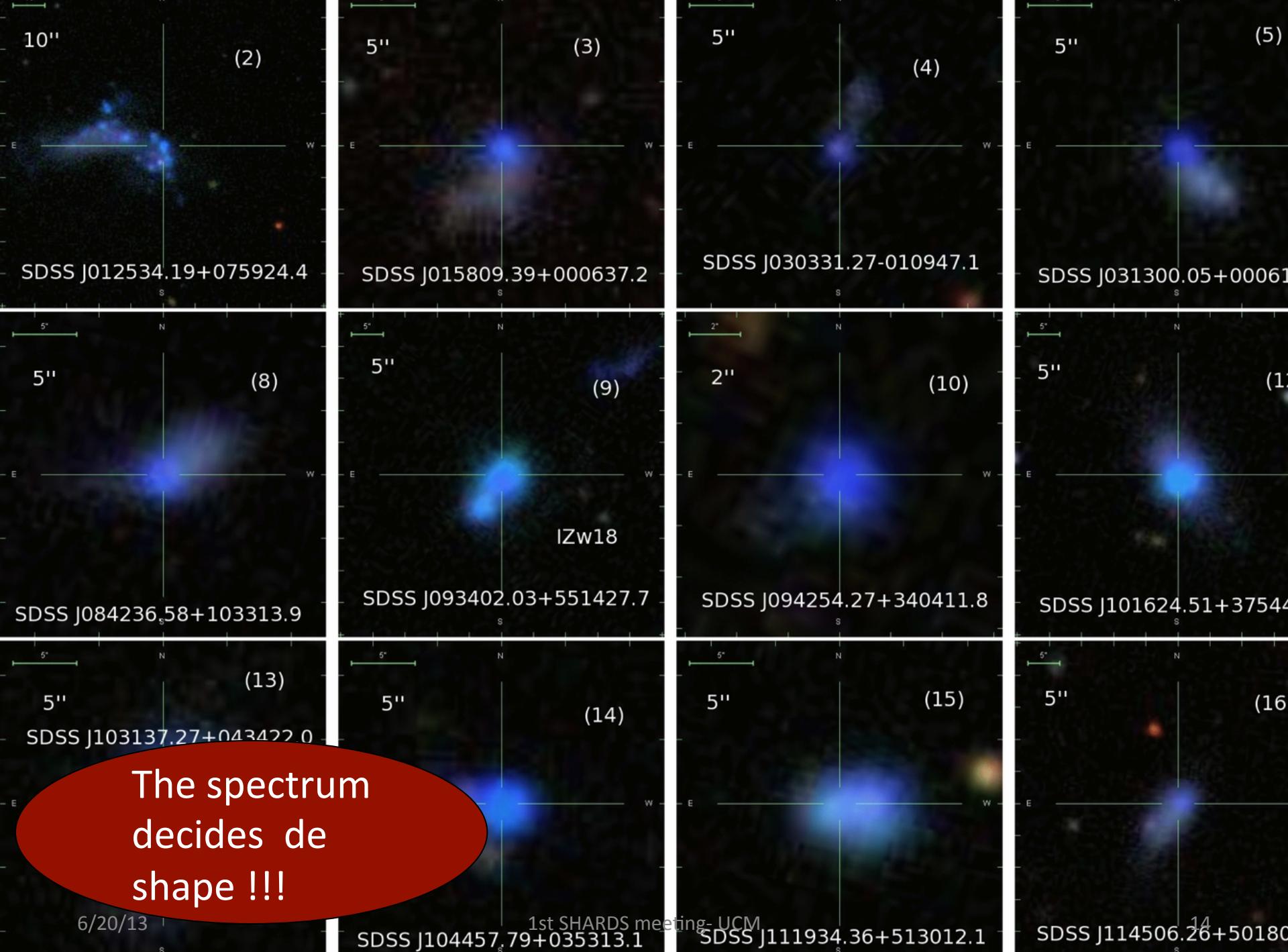
# Spectra and morphological types match



# Systematic Search for Extremely Metal-poor Galaxies in the Sloan Digital Sky Survey (DR9)

A. B. Morales-Luis et al. 2011 ApJ 743 77

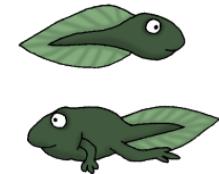




# Why are XMP galaxies tadpole-cometary galaxies?

We do not know yet ...

but answering the question is of interest beyond  
the field of XMP galaxies



Why is this question interesting at all?  
(more than a mere curiosity)

- Tadpole (cometary) galaxies are rare objects in the local universe (0.1% of the Kiso galaxies; Elmegreen et al. 2012), but very common at high redshift (10% of galaxies larger than 10 pix in the UDF Elmegreen et al. 2007).

- They may be disks in an early phase in the process of formation: gas falling in, self excited gravitational instability, interaction with DM clumps or dwarf galaxies, ram pressure striping ...

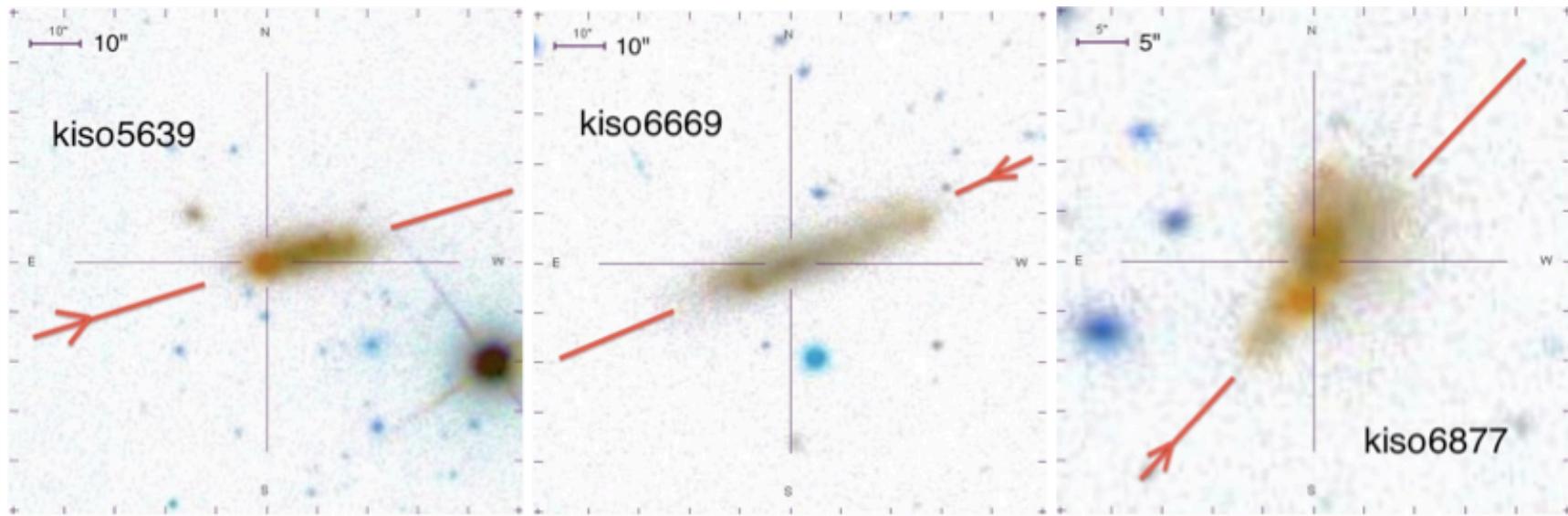
- XMP are primitive objects from a chemical point of view, but they may also be primitive with a dynamical point of view.

- Do XMP Gs represent disk-forming systems?

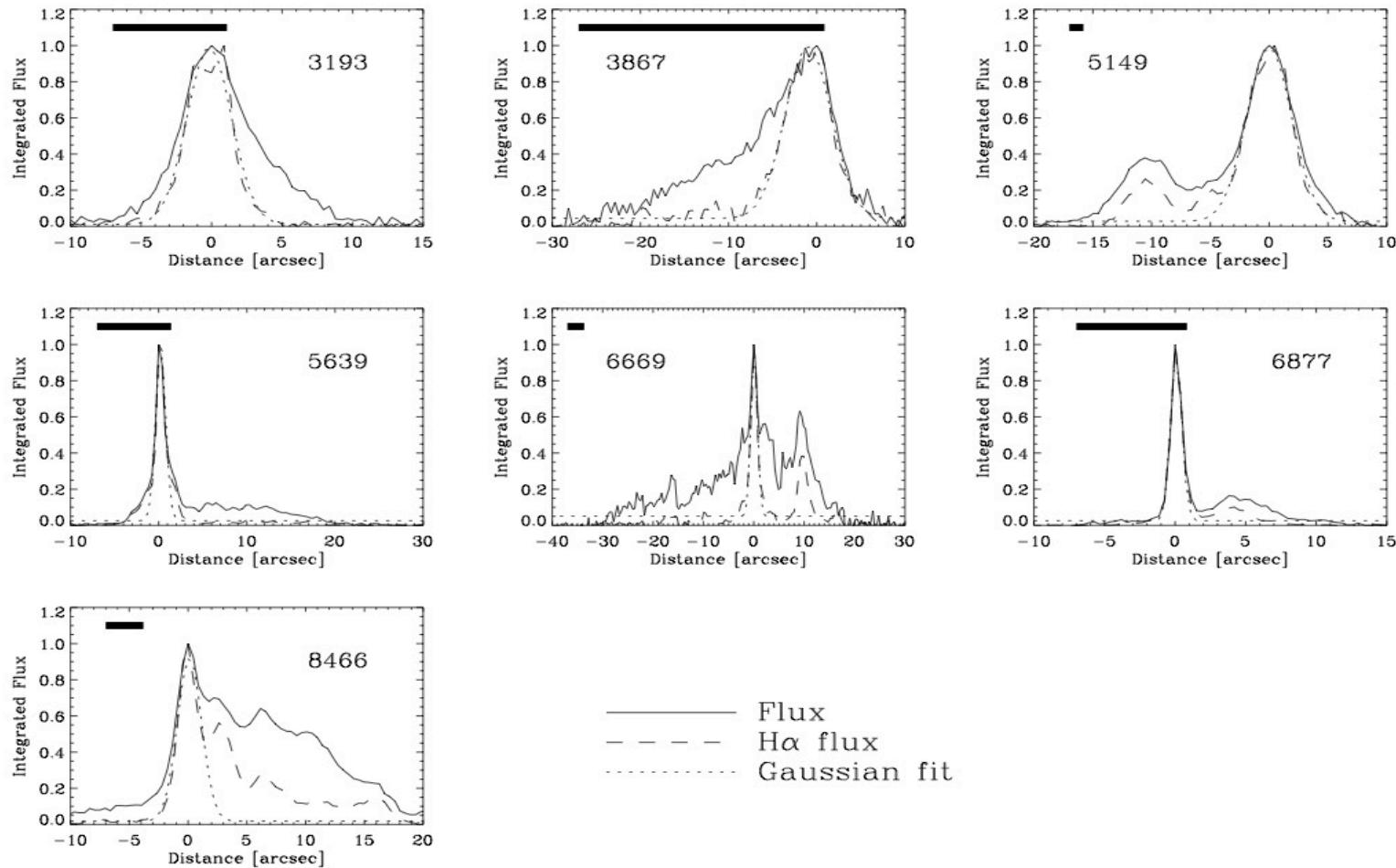
- Do XMP Gs represent nearby laboratories to study disk formation? (in pristine conditions?)

# *local tadpole galaxies (chemo-dynamics)*

- High spectral resolution 1D spectra around H $\alpha$  ( $R \sim 10000$ )
- IDS@INT and ALFOSC@NOT
- 7 tadpoles (50%) of the sample
- spectrograph slit along then head-tail direction

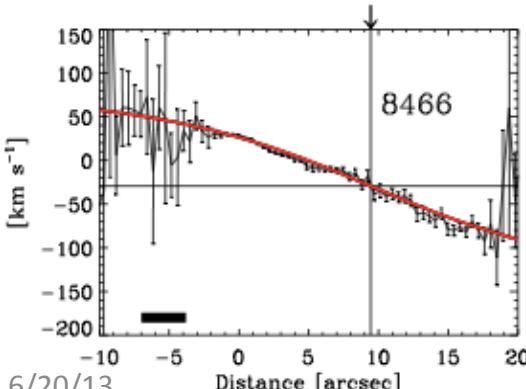
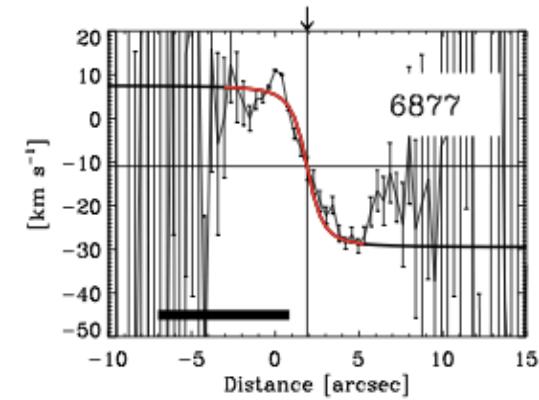
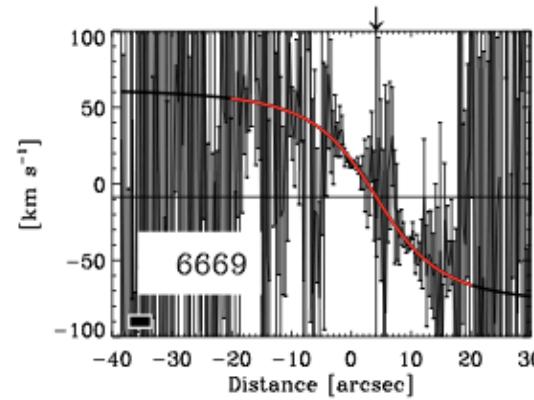
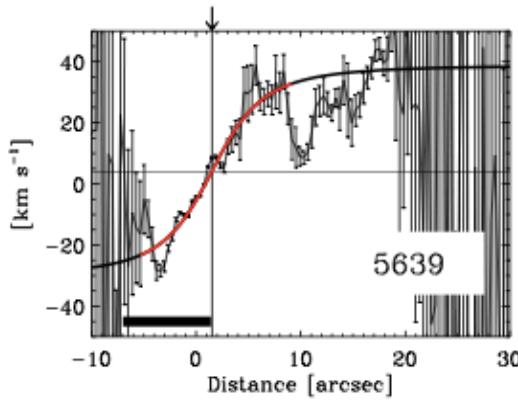
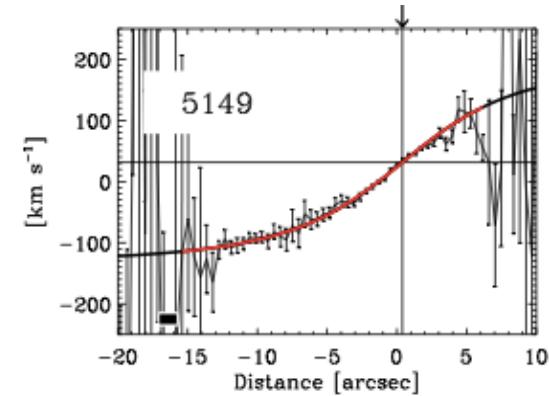
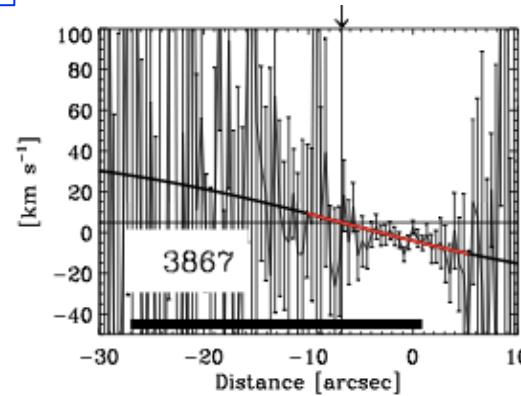
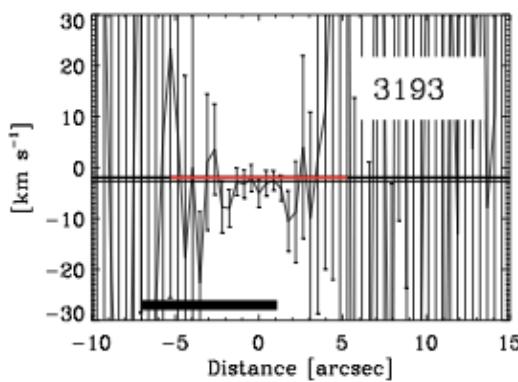


# Brightness Profiles



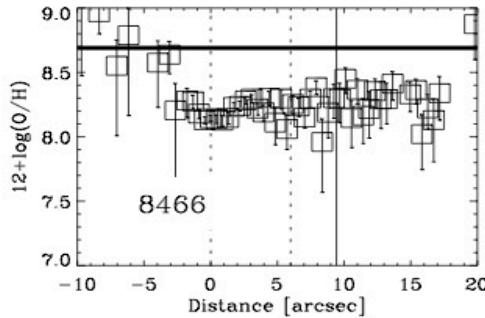
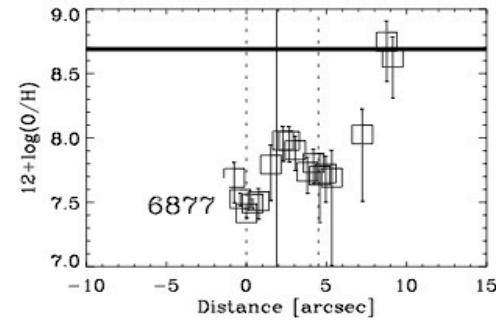
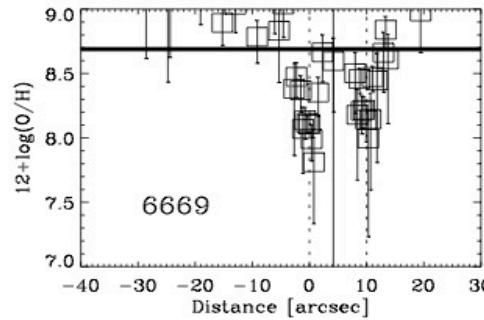
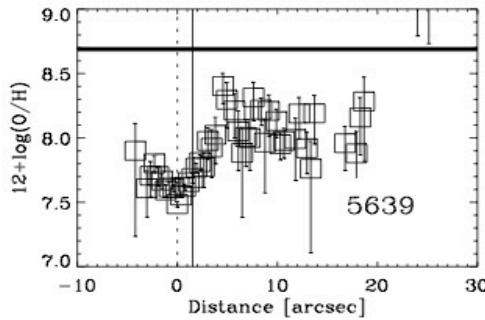
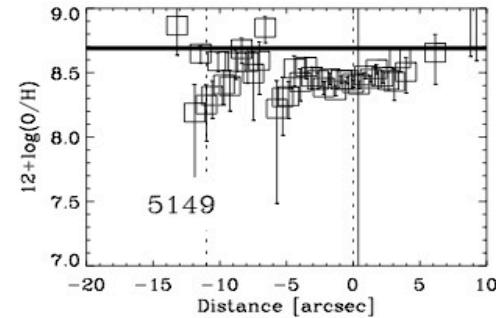
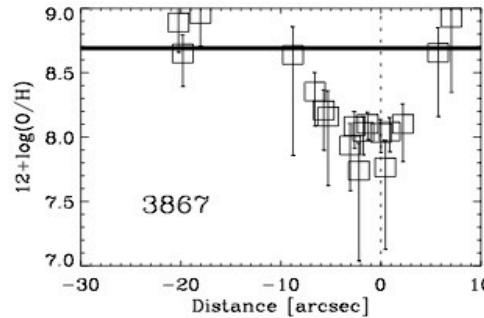
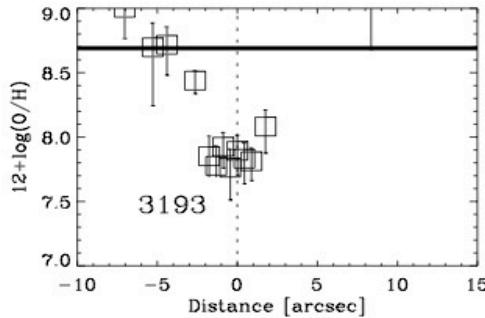
- lopsided light distribution (as expected)
- Tadpole head defined to be at zero distance

# Velocity Curves



- 5(or 6) out of 7 **rotate**
- Tadpole head often off-centered with respect to the center of rotation

# Metallicity (Oxygen abundance)



- O/H based on [NII]6583/H $\alpha$

----- Ha Flux Peaks  
————— Rotation Center  
————— Solar Abundance

- Metallicity **gradients**, lower at center and **increasing outwards**
- two of them XMP (extremely metal poor) galaxy!
- outliers of the mass-metallicity relationship

# *Cold-flow accretion as a mode of galaxy growth*

UDF galaxies are different from the Local Universe's ones, more patchy and “amorphous” (no big mergers).

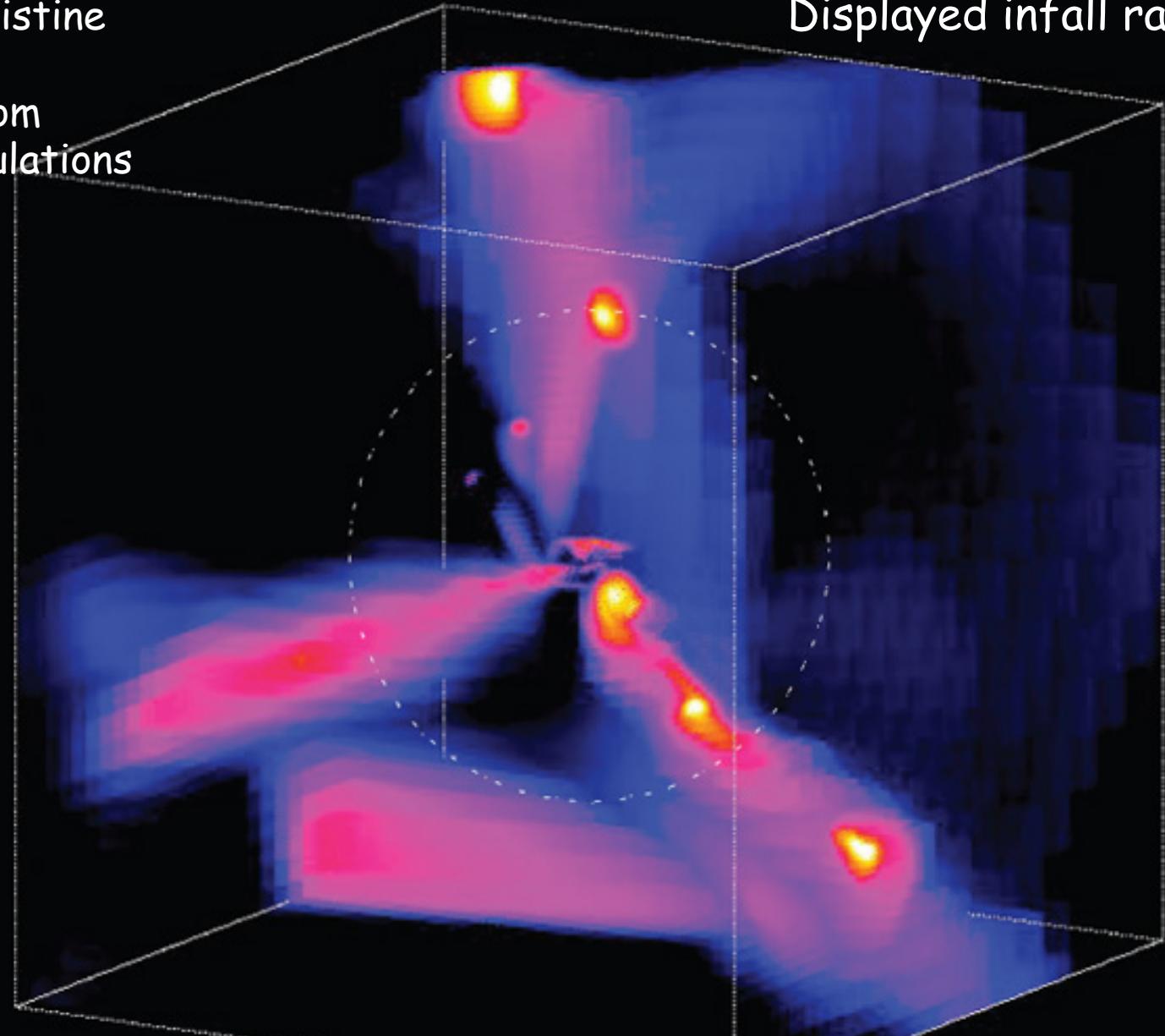
The luminosity profiles – thick clumpy disk's like (e.g., Elmegreen et al. 2005), with some rotation (e.g., Genzel et al. 2008)

Accretion of external flows of pristine gas (cold-flows), which numerical simulations predicts is thought to be the main mode of disks growth at high z..

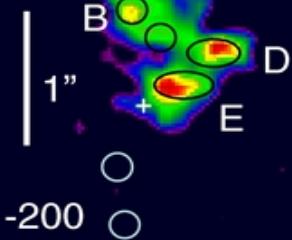
The incoming gas hit and heats the proto-disk, and is expected to form giant clumps that spiral in and merge into a central spheroid (Noguchi 1999; Genzel et al. 2008; Elmegreen et al. 2008).

Streams of pristine  
gas in three  
dimensions from  
numerical simulations

Displayed infall rate



D3a15504  
 $z=2.39$



-200

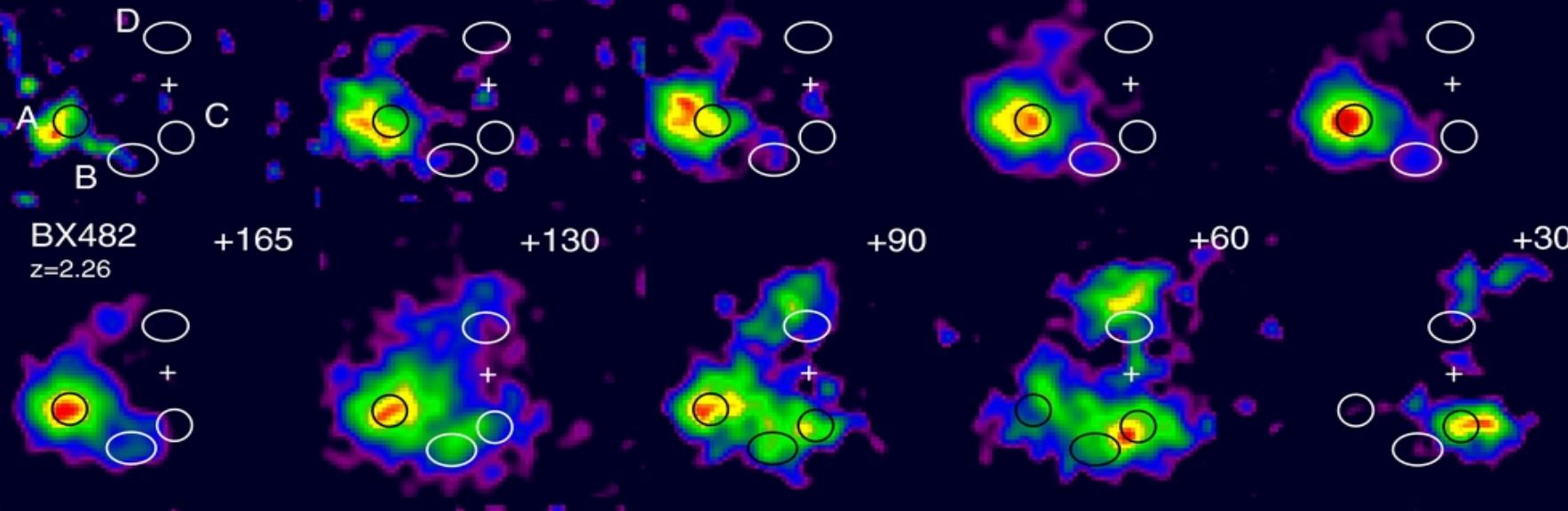
-160

-130

-90

+130

+200



BX482  
 $z=2.26$

+165

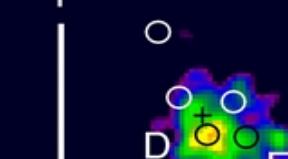
+130

+90

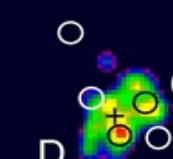
+60

+30

ZC782941  
 $z=2.18$

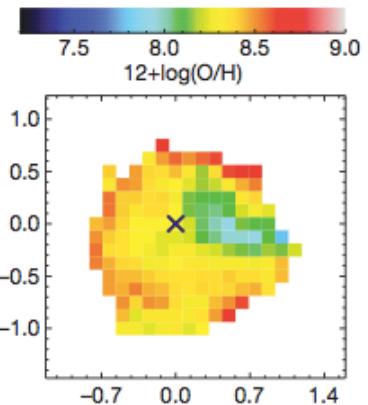
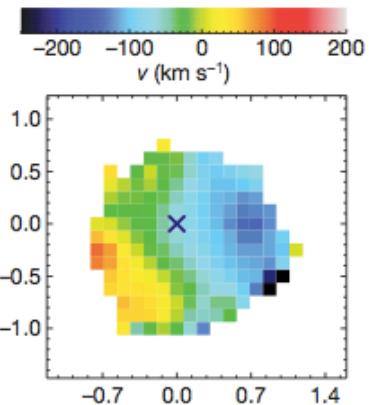
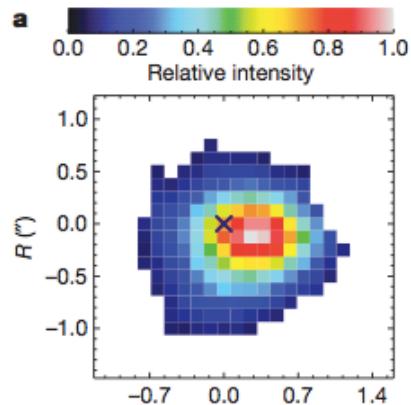


-195  
6/20/13

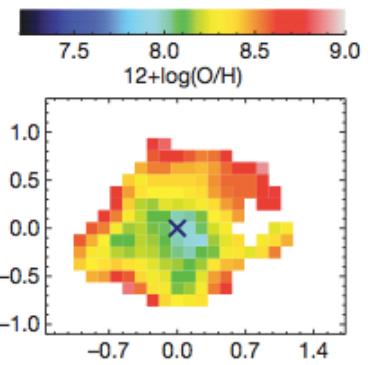
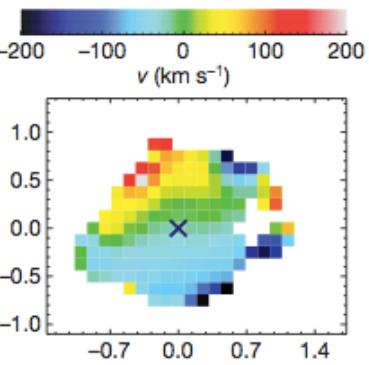
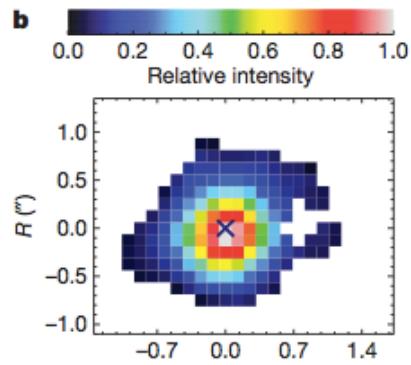


*Genzel + sins, 2011, Ap.J., 733*

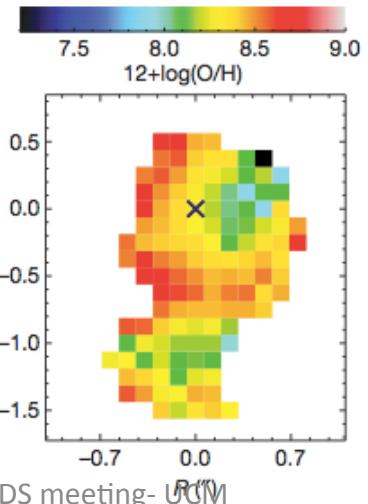
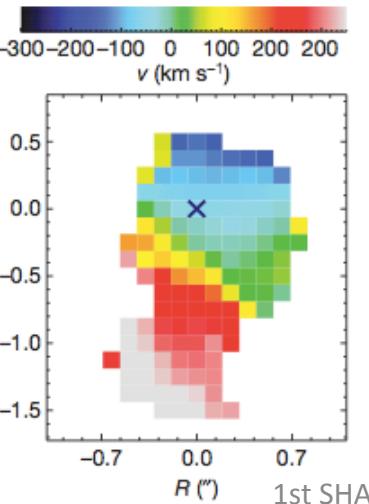
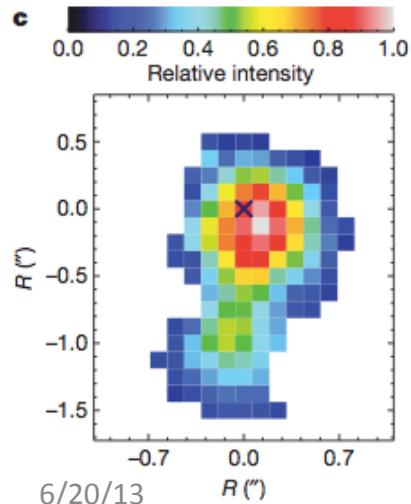
1st SHARDS meeting - UCM



Evidence for cold-flows at high redshift



Cresci et al., Nat. 2010,



$z$	$\log(M_*/M_\odot)$
3.065	$10.68^{+0.16}_{-0.54}$
3.219	$10.03^{+0.40}_{-0.08}$
3.288	$10.86^{+0.18}_{-0.41}$

# Conclusions

- Tadpoles are rare in local universe, but common at high-z, where they are commonly interpreted as disks being assembled
- Extremely metal poor (XMP) are chemically primitive objects, and, surprisingly, they turn out to have tadpole or cometary morphology
- Local Tads follow the scaling relations of high-z ones
- Local Tads rotate with the head displaced from the rotation center.
- Metallicity drops at the head (largest SFR)

★ All this results combined suggest that **XMP-tadpoles** are **disks in early stages of assembling** with its **star-formation sustained by accretion of external metal poor gas**.

- **Cold-flows**, expected from numerical simulations of galaxy formation, observed for the very first time the local universe

# COMING SOON

- Extremely Metal-Poor Galaxies: The HI content.  
(Fihlo et al. Submitted)
- Search for the cool flows in HI maps.
- Complete characterization of a representative local sample (GTC observations).
- Revise previous Sloan results with new eyes.

Our extended team: *M. Fihlo, A.B. Morales-Luis, J.A.L. Aguerri, R. Amorin, P. Papaderos, Y. Ascasibar, J.M. Gomes, P. Lagos, A. Humphery, J.M. Vilchez, J. Méndez-Abreu.*  
*(IAC, Porto- Portugal; UAM- Madrid, IAA- Granada)*

Thanks  
shardt



Fishing TADs in Shards  
shards ni SHARDT